Report on the International Training in Tenerife

Masashi NASHIMOTO

(Astronomical Institute)

GP-PU Progress Status Presentation — Spring 2019 —

International Training

Date: 4 September - 18 December 2018

Place : Instituto de Astrofísica de Canarias Tenerife island, Spain



Teide Observatory



Canary Islands



International Training







I Training











enjoyed Tenetie life

tiona





Anomalous Microwave Emission

AME is a component of diffuse Galactic radiation observed at frequencies in the range ~10-60GHz.

Polarized AME has not detected yet.

The origin of AME is unknown.



Purpose of My Staying in Tenerife

Candidates of the AME Source

- spinning dust emission
- magnetic dipole emission
- thermal emission from amorphous dust

the relation between AME and the physical mechanism of amorphous dust emission has not been studied yet

we studied whether the amorphous dust model is able to explain diagnostics of whole frequency range spectrum provided by **QUIJOTE** and other ancillary experiments.

Amorphous Dust Emission

Two-Level Systems model



at higher temperature --> absorption is dominated by relaxation processes

- phonon assisted tunneling
- hopping over the barrier



energy difference ~ 1K ~ 10GHz

Purpose of My Staying in Tenerife

Candidates of the AME Source

- spinning dust emission
- magnetic dipole emission
- thermal emission from amorphous dust

the relation between AME and the physical mechanism of amorphous dust emission has not been studied yet

we studied whether the amorphous dust model is able to explain diagnostics of whole frequency range spectrum provided by **QUIJOTE** and other ancillary experiments.

Q-U-I JOint Tenerife Experiment



QUIJOTE is an observational experiment operating in the microwave range.

--> QUIJOTE provides information about AME



Amorphous Dust & AME

W43



Our model provides consistent results both with intensity and polarization SEDs



frequency [Hz]











Summary

I had been in Tenerife from September to December

I studied amorphous dusts and AME with QUIJOTE members

 We showed the amorphous dust emission model can explain AME

 We will study toward CMB component separation using our amorphous dust model