

Multi-Pomeron Repulsion and Mass of Hyperon-Mixed Neutron Star

Y. Yamamoto

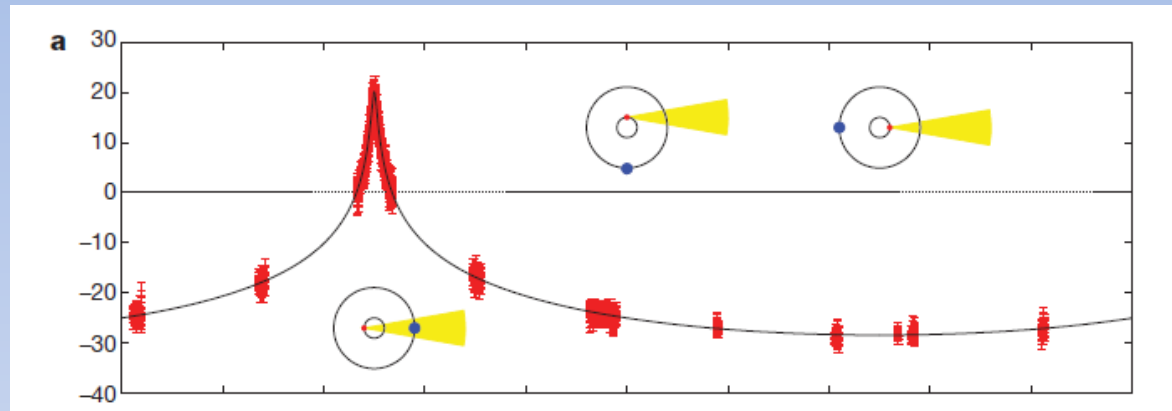
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2010 PSR J1614-2230 $(1.97 \pm 0.04) M_{\odot}$



Shapiro delay measurement

2013 PSR J0348-0432 $(2.01 \pm 0.04) M_{\odot}$

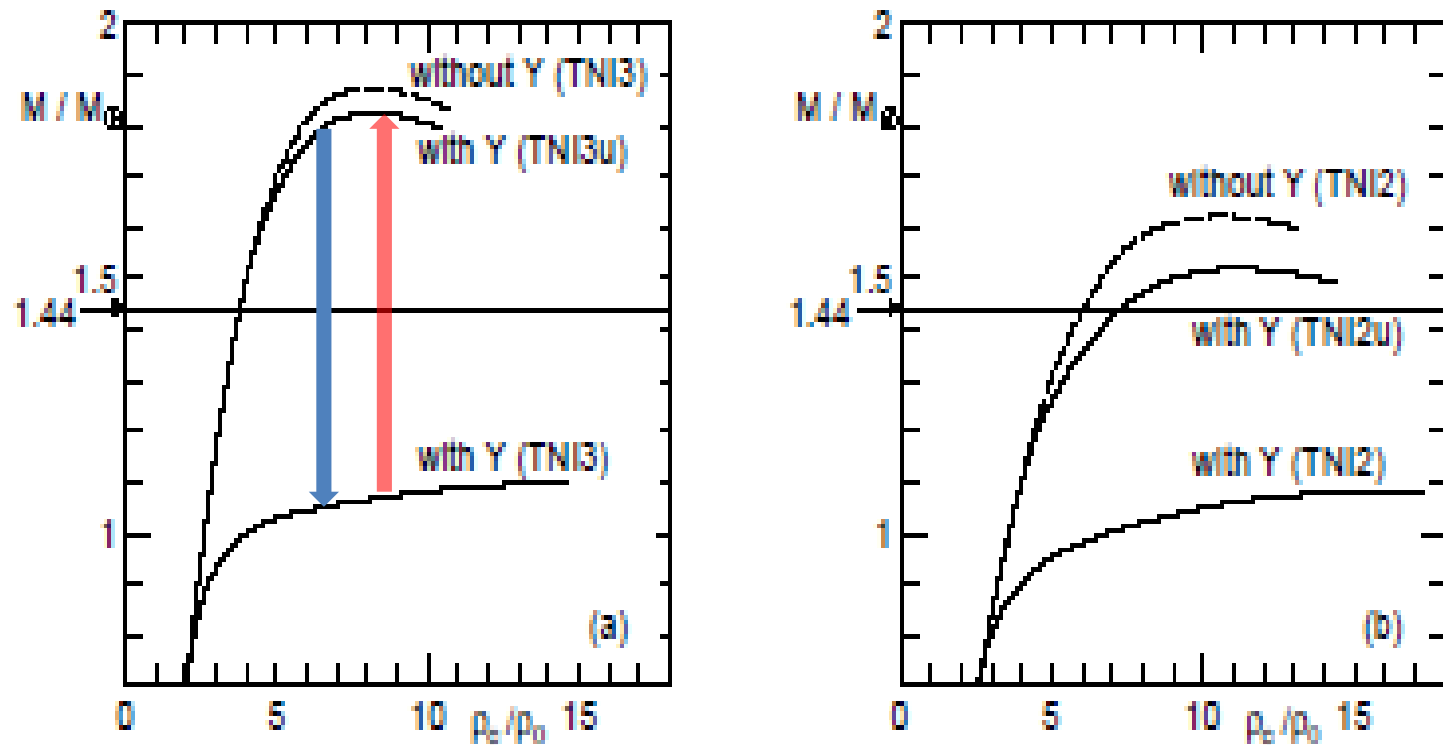


Fig. 9. The mass of a neutron star in units of the solar mass $M_\odot$ as functions of the central baryon density ρ_c with use of (a) TNI3 and (b) TNI2. The notation here is the same as in Fig. 8.

Softening by hyperon mixing to neutron-star matter

Massive ($2M_{\odot}$) neutron stars

Softening of EOS by hyperon mixing

Compatible ?

An idea is Universal Three-Baryon Repulsion (TBR)
by Takatsuka

Modeling of TBR in ESC = Multi-Pomeron exchange Potential

in this talk

Our strategy for neutron stars

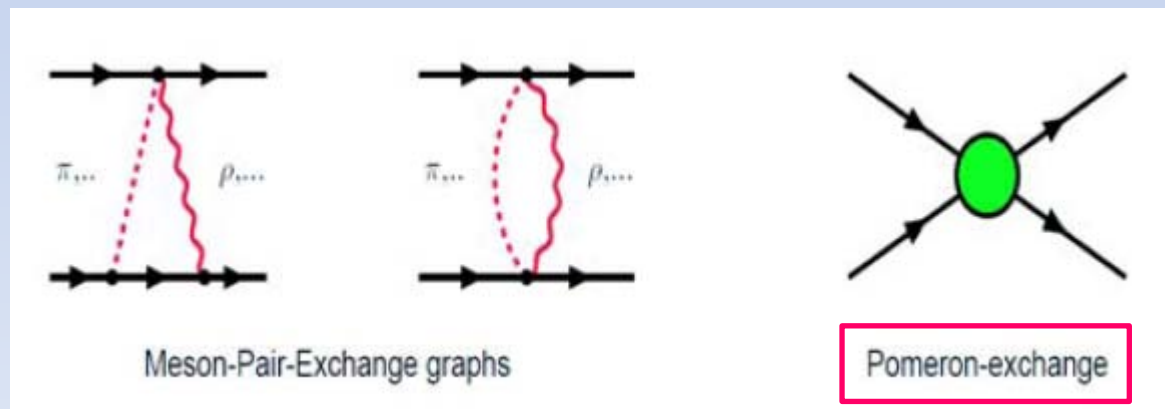
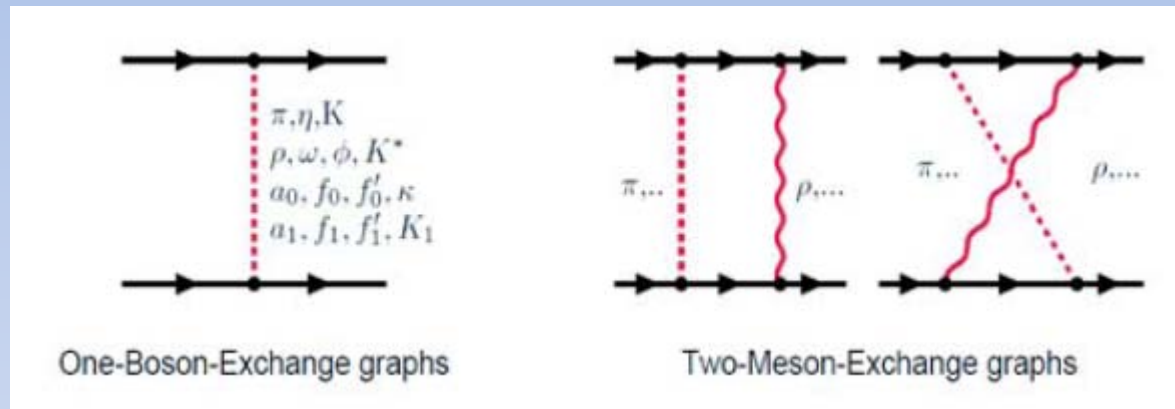
Neutron-star EOS derived from
Baryon-Baryon interaction model
in relation to Earth-based experiments

without ad hoc parameter for stiffness of EOS

on the basis of G-matrix theory

Extended Soft-Core Model (ESC)

- Two-meson exchange processes are treated explicitly
- Meson-Baryon coupling constants are taken consistently with Quark-Pair Creation model



repulsive cores

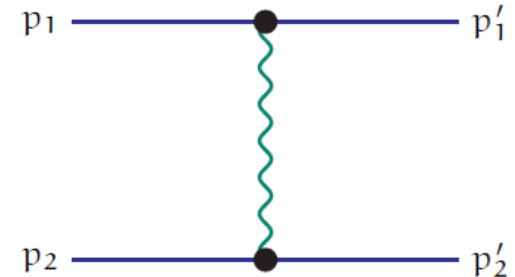
Two-body Potential from Pomeron-exchange

Lagrangian & Propagator

$$\mathcal{L}_{\text{PNN}} = g_{\text{P}} \bar{\psi}(x) \psi(x) \sigma_{\text{P}}(x)$$

$$\Delta_{\text{F}}^{\text{P}}(k^2) = + \exp(-\mathbf{k}^2/4m_{\text{P}}^2)/\mathcal{M}^2$$

scaling mass $\mathcal{M} = 1\text{GeV}$.



$$\begin{aligned} -iM_{\text{P}}(p'_1, p'_2; p_1, p_2) &= (+i)^2 g_{\text{P}}^2 [\bar{u}(p')u(p)] [\bar{u}(-p')u(-p)] \cdot i\Delta_{\text{F}}^{\text{P}}[(p' - p)^2] \\ M_{\text{P}}(p'_1, p'_2; p_1, p_2) &= g_{\text{P}}^2 [\bar{u}(p')u(p)] [\bar{u}(-p')u(-p)] \cdot \Delta_{\text{F}}^{\text{P}}[(p' - p)^2] \\ &\approx g_{\text{P}}^2 \exp(-\mathbf{k}^2/4m_{\text{P}}^2)/\mathcal{M}^2, \end{aligned}$$

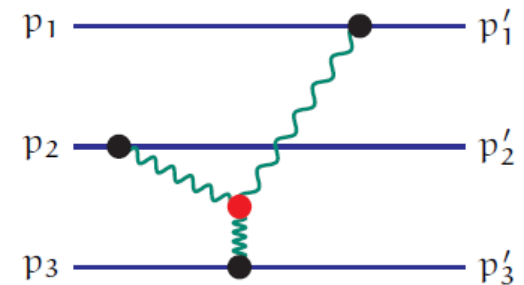
$$\begin{aligned} V_{\text{P}}(r) &= \int \frac{d^3k}{(2\pi)^3} e^{i\mathbf{k} \cdot \mathbf{x}} M_{\text{P}}(\mathbf{p}' - \mathbf{p}) \delta(\mathbf{k} - \mathbf{p}' + \mathbf{p}) \\ &= \frac{g_{\text{P}}^2}{4\pi} \frac{4}{\sqrt{\pi}} \frac{m_{\text{P}}^3}{\mathcal{M}^2} \exp(-m_{\text{P}}^2 r_{12}^2). \end{aligned}$$

Two-body repulsive core

Three(Four) -body Potential from the Triple(Quadruple) -pomeron vertex

$$\mathcal{L}_{\text{PPP}} = g_{3\text{P}} \mathcal{M} \sigma_{\text{P}}^3(\chi)/3!$$

$$\left[\mathcal{L}_{\text{PPPP}} = g_{4\text{P}} \sigma_{\text{P}}^4(\chi)/4! \right]$$



$$\begin{aligned} M_{3\text{P}}(p'_1, p'_2, p'_3; p_1, p_2, p_3) &= g_{3\text{P}} g_{\text{P}}^3 \prod_{i=1}^3 \{ [\bar{u}(p'_i) u(p_i)] \Delta_{\text{F}}^{\text{P}}[(p'_i - p_i)^2] \} \\ &\approx g_{3\text{P}} g_{\text{P}}^3 \prod_{i=1}^3 \Delta_{\text{F}}^{\text{P}}[(p'_i - p_i)^2] . \end{aligned}$$

$$\begin{aligned} V(\mathbf{x}'_1, \mathbf{x}'_2, \mathbf{x}'_3; \mathbf{x}_1, \mathbf{x}_2, \mathbf{x}_3) &= \int \prod_{i=1}^3 \frac{d^3 p'_i}{(2\pi)^3} \frac{d^3 p_i}{(2\pi)^3} \cdot \\ &\times \prod_{i=1}^3 e^{-i(\mathbf{p}'_i \cdot \mathbf{x}'_i - \mathbf{p}_i \cdot \mathbf{x}_i)} \cdot \delta \left(\sum \mathbf{p}'_i - \sum \mathbf{p}_i \right) \\ &\times M_{3\text{P}}(p'_1, p'_2, p'_3; p_1, p_2, p_3) . \end{aligned}$$

$$V(\mathbf{x}'_1, \mathbf{x}'_2, \mathbf{x}'_3; \mathbf{x}_1, \mathbf{x}_2, \mathbf{x}_3) \equiv V(\mathbf{x}_1, \mathbf{x}_2, \mathbf{x}_3) \delta(\mathbf{x}'_1 - \mathbf{x}_1) \delta(\mathbf{x}'_2 - \mathbf{x}_2) \delta(\mathbf{x}'_3 - \mathbf{x}_3) \quad (15)$$

$$\begin{aligned} V(\mathbf{x}_1, \mathbf{x}_2, \mathbf{x}_3) &= g_{3\text{P}} g_{\text{P}}^3 \prod_{i=1}^3 \int \frac{d^3 \mathbf{k}_i}{(2\pi)^3} \prod_{i=1}^3 e^{-i\mathbf{k}_i \cdot \mathbf{x}_i} \cdot (2\pi)^3 \delta(\mathbf{k}_1 + \mathbf{k}_2 + \mathbf{k}_3) \cdot \\ &\times \exp(-\mathbf{k}_1^2/4m_{\text{P}}^2) \exp(-\mathbf{k}_2^2/4m_{\text{P}}^2) \exp(-\mathbf{k}_3^2/4m_{\text{P}}^2) \cdot \mathcal{M}^{-5} , \end{aligned}$$

Three- and Four-body repulsions with parameters $g_{3\text{P}}$ & $g_{4\text{P}}$

The effective two-body potential

$$V_{\text{eff}}(\mathbf{x}_1, \mathbf{x}_2) = \rho_{\text{NM}} \int d^3x_3 V(\mathbf{x}_1, \mathbf{x}_2, \mathbf{x}_3)$$

$$\begin{aligned} V_{\text{eff}}(\mathbf{x}_1, \mathbf{x}_2) &= g_{3\text{P}} g_{\text{P}}^3 \frac{\rho_{\text{NM}}}{\mathcal{M}^5} \cdot \int \frac{d^3k_1}{(2\pi)^3} \int \frac{d^3k_2}{(2\pi)^3} e^{-i\mathbf{k}_1 \cdot \mathbf{x}_1} e^{-i\mathbf{k}_2 \cdot \mathbf{x}_2} \cdot \\ &\quad \times (2\pi)^3 \delta(\mathbf{k}_1 + \mathbf{k}_2) \exp\left(-\mathbf{k}_1^2/4m_{\text{P}}^2\right) \exp\left(-\mathbf{k}_2^2/4m_{\text{P}}^2\right) \\ &= g_{3\text{P}} g_{\text{P}}^3 \frac{\rho_{\text{NM}}}{\mathcal{M}^5} \cdot \int \frac{d^3k_1}{(2\pi)^3} e^{-i\mathbf{k}_1 \cdot (\mathbf{x}_1 - \mathbf{x}_2)} \cdot \exp\left(-\mathbf{k}_1^2/2m_{\text{P}}^2\right) \\ &= g_{3\text{P}} g_{\text{P}}^3 \frac{\rho_{\text{NM}}}{\mathcal{M}^5} \cdot \frac{1}{4\pi} \frac{4}{\sqrt{\pi}} \left(\frac{m_{\text{P}}}{\sqrt{2}}\right)^3 \exp\left(-\frac{1}{2}m_{\text{P}}^2 r_{12}^2\right) \cdot \end{aligned}$$

$$\left[g_{4\text{P}} g_{\text{P}}^4 \frac{\rho_{\text{NM}}^2}{\mathcal{M}^8} \cdot \frac{1}{4\pi} \frac{4}{\sqrt{\pi}} \left(\frac{m_{\text{P}}}{\sqrt{2}}\right)^5 \exp\left(-\frac{1}{2}m_{\text{P}}^2 r_{12}^2\right) \right]$$

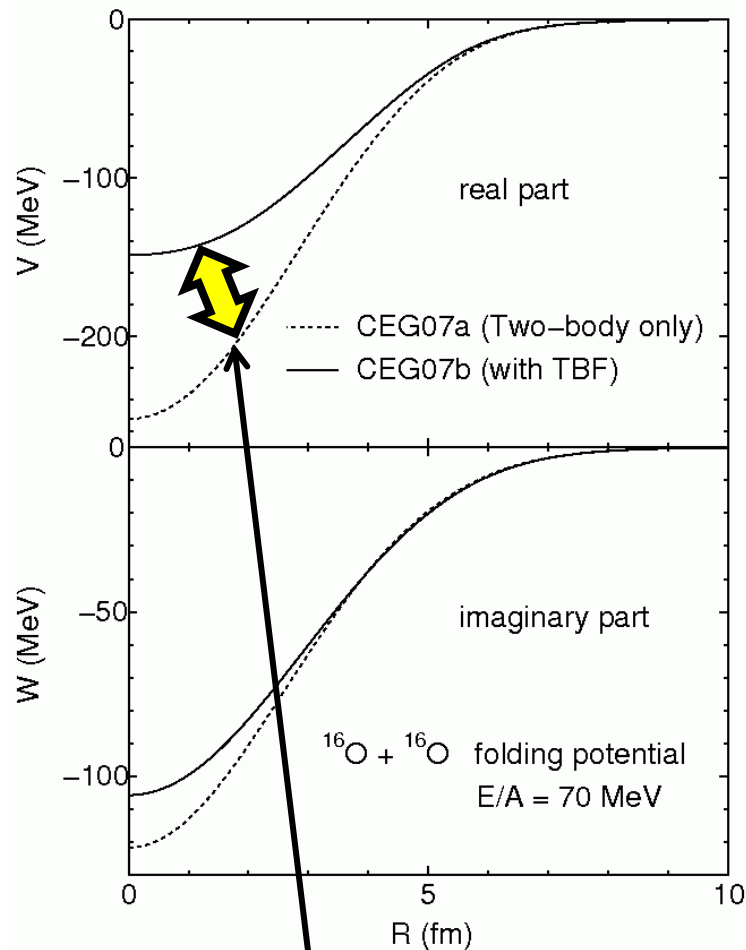
How to determine coupling constants g_{3P} and g_{4P} ?



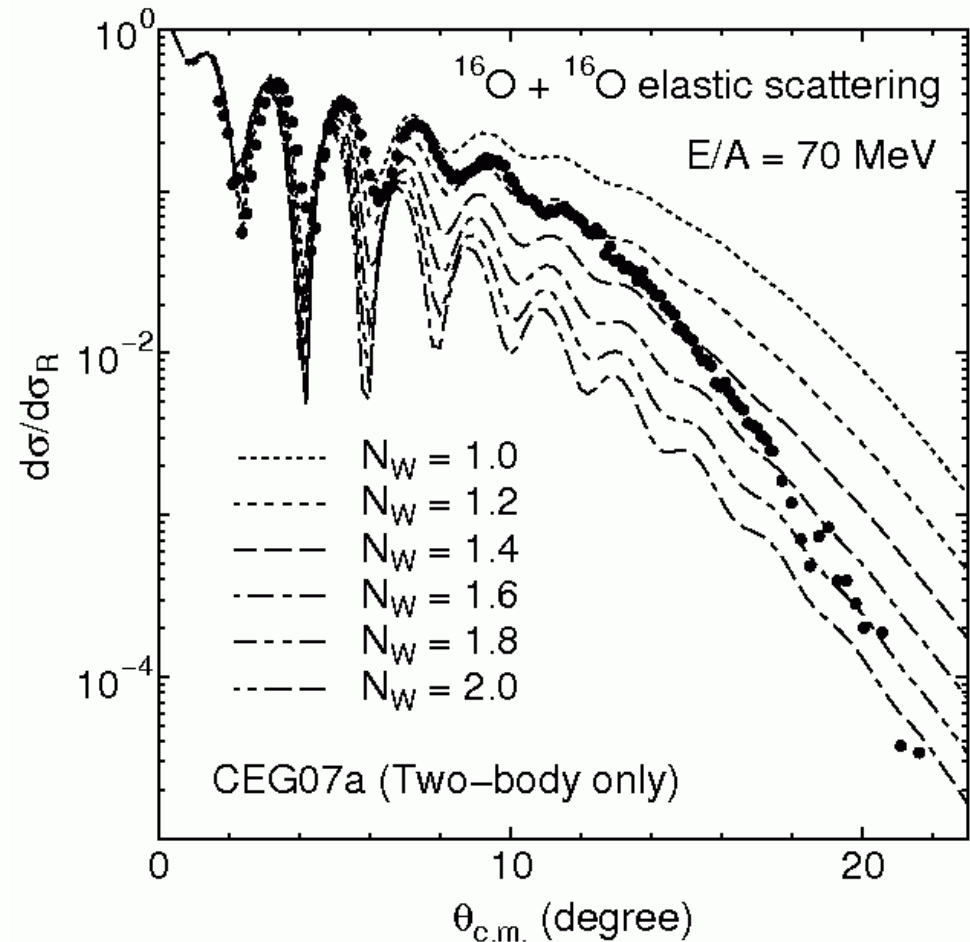
Nucleus-Nucleus scattering data

$^{16}\text{O} + ^{16}\text{O}$ elastic scattering $E/A = 70 \text{ MeV}$

with G-matrix
folding model



Effect of three-body force



$$U_{DFM} = V_{DFM} + iN_W W_{DFM}$$

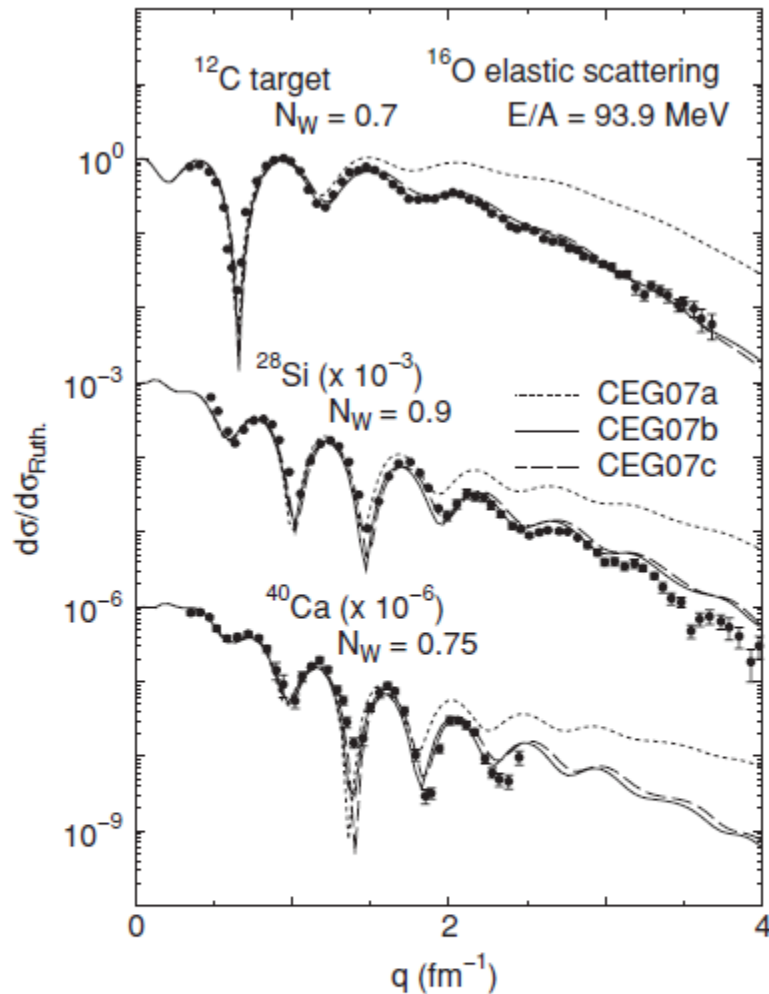


FIG. 13. Rutherford ratio of the cross sections for elastic scattering of ^{16}O by the ^{12}C , ^{28}Si , and ^{40}Ca targets at $E/A = 93.9$ MeV, calculated with the three types of complex G -matrix interactions, which are compared with the experimental data from Ref. [35]. The abscissa is the momentum transfer q defined as $q = 2k \sin \frac{\theta}{2}$, where k is the asymptotic momentum.

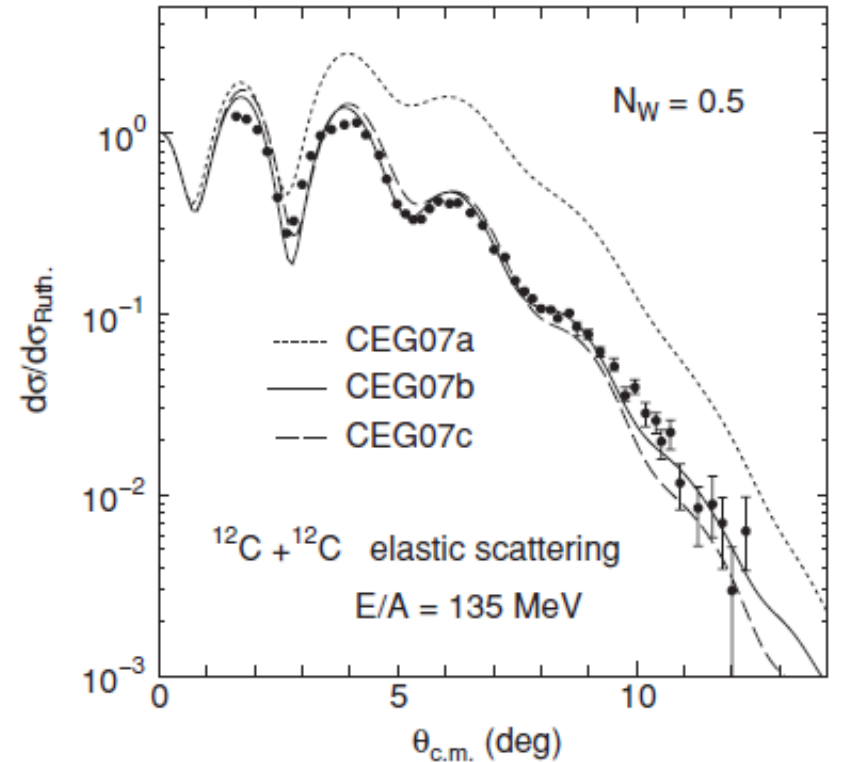


FIG. 15. Rutherford ratio for the $^{12}\text{C} + ^{12}\text{C}$ system at $E/A = 135$ MeV calculated with the three types of complex G -matrix interactions, which are compared with the experimental data from Ref. [36].

ESC08c + MPP + TNA

repulsive attractive

phenomenological

$$V_{TNA}(r; \rho) = V_0 \rho \exp(-\eta \rho) \exp(-(r/2.0)^2) (1 + P_r)/2$$

V_0 and η are determined so as to reproduce saturation density/energy

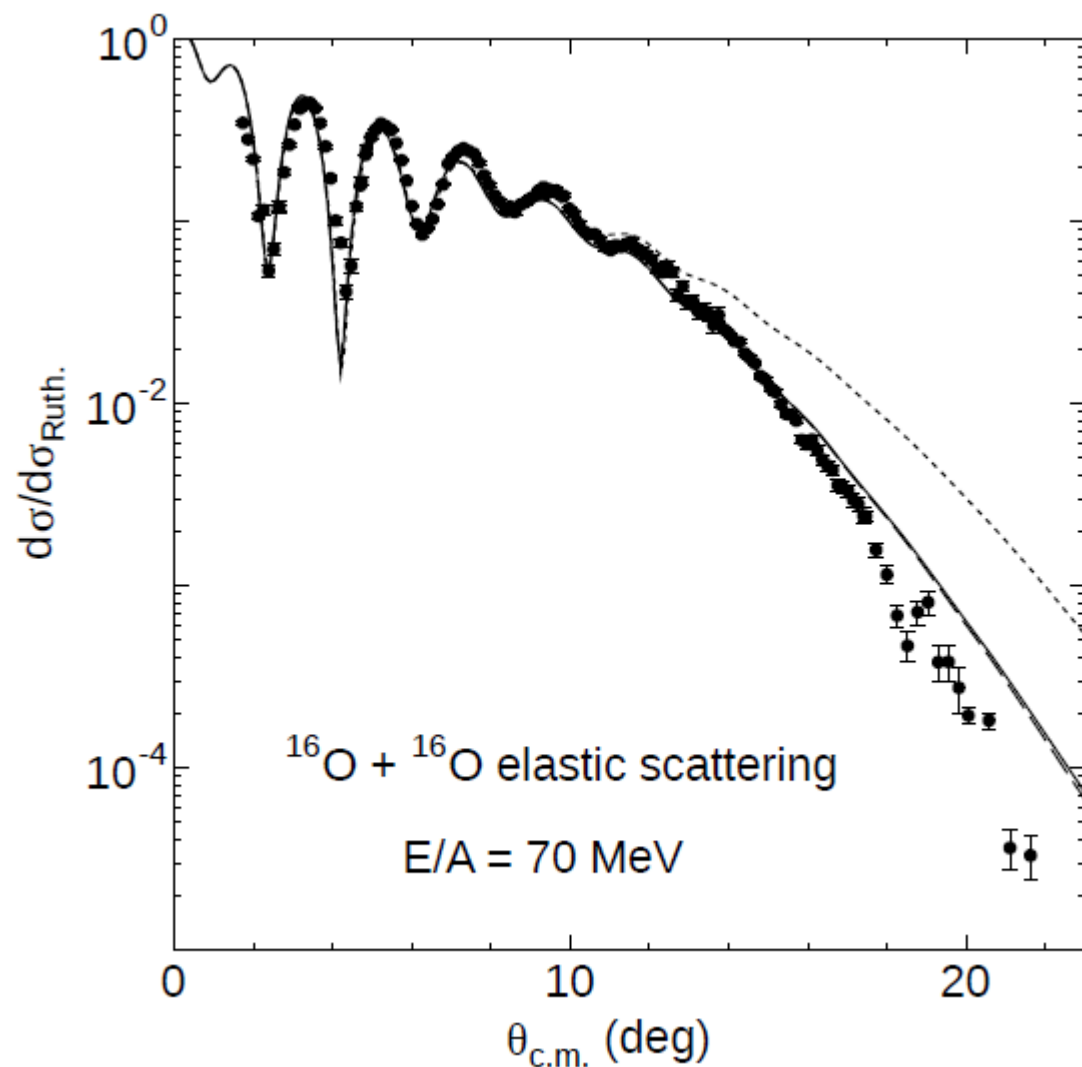
<u>MP1a</u> :	$(g_P^{(3)}, g_P^{(4)}) = (2.34, 30.)$	$V_0 = -32.8$	$\eta = 3.5$
<u>MP2a</u> :	$(g_P^{(3)}, g_P^{(4)}) = (2.94, 0.0)$	$V_0 = -45.0$	$\eta = 5.4$

MPP

TNA

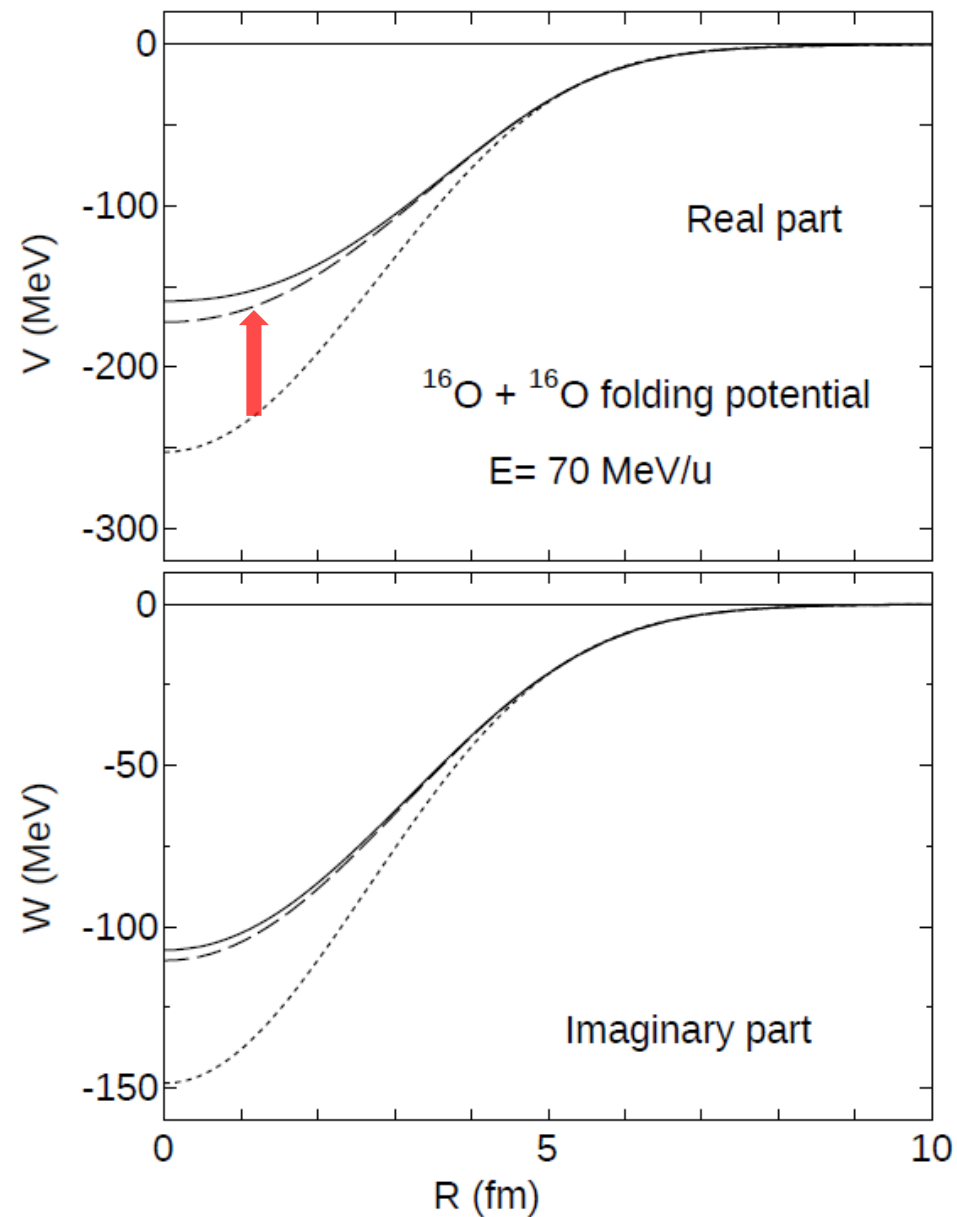
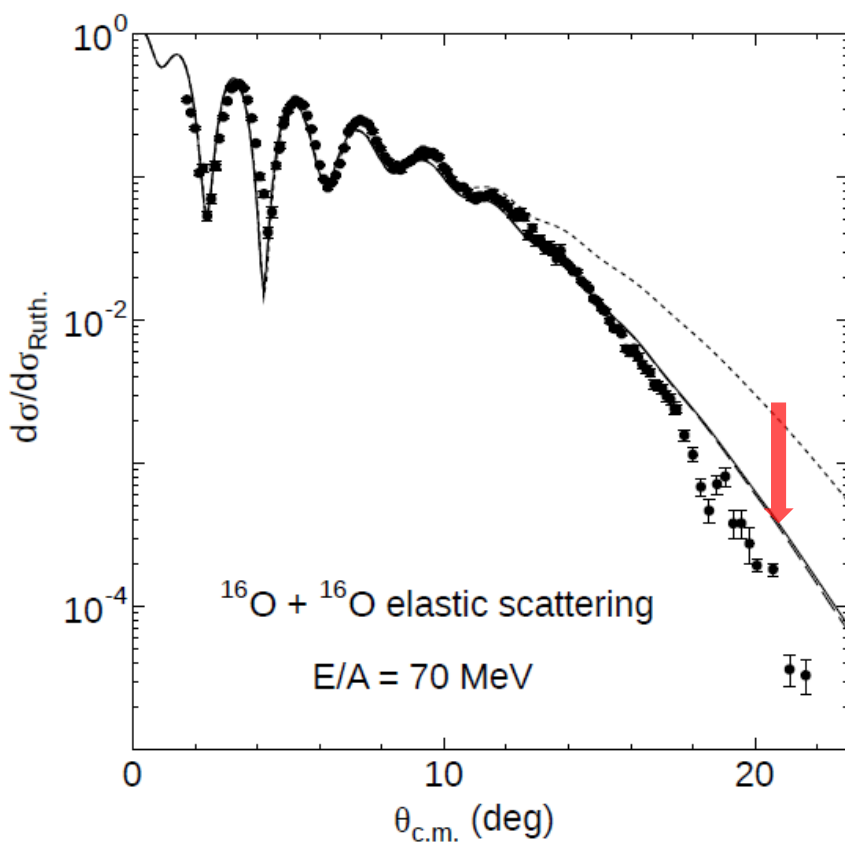
Ratio g_{4P}/g_{3P} is not determined in our analysis ---→ two versions MP1a & MP2a

with G-matrix folding model



solid: ESC08c+MP1a+TNA
dashed: ESC08c+MP2a+TNA
dotted: ESC08c only

MP1a (MP2a): with (without) 4-body repulsion



Frozen-Density Approximation

Two Fermi-spheres separated in momentum space

can overlap in coordinate space without disturbance of Pauli principle

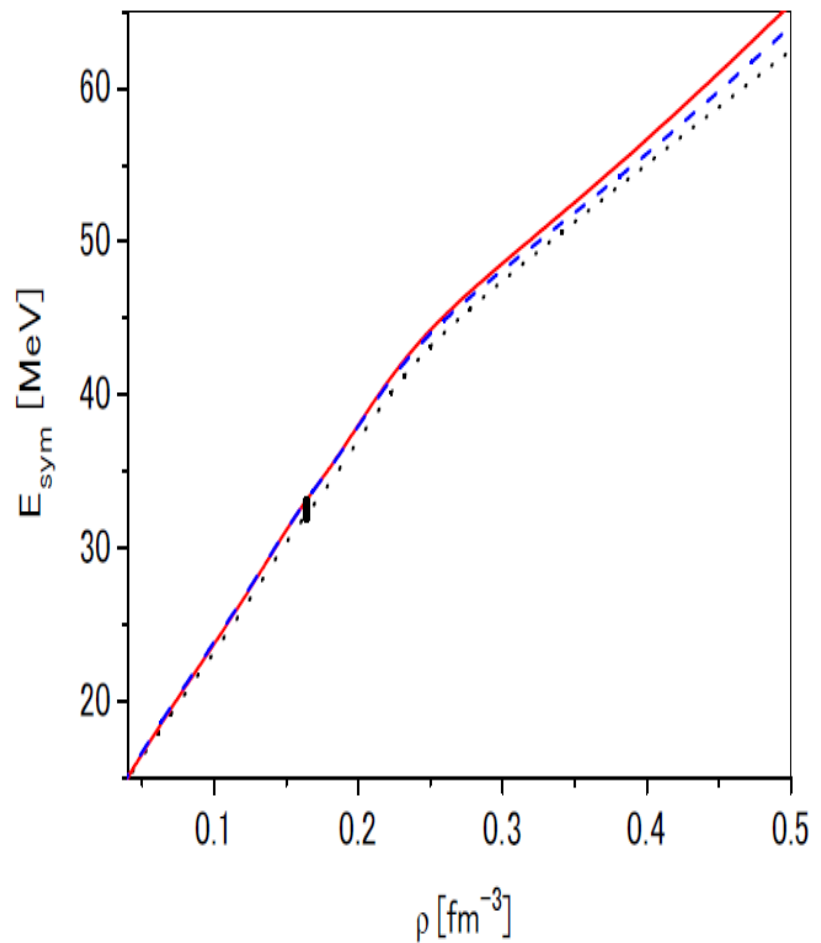
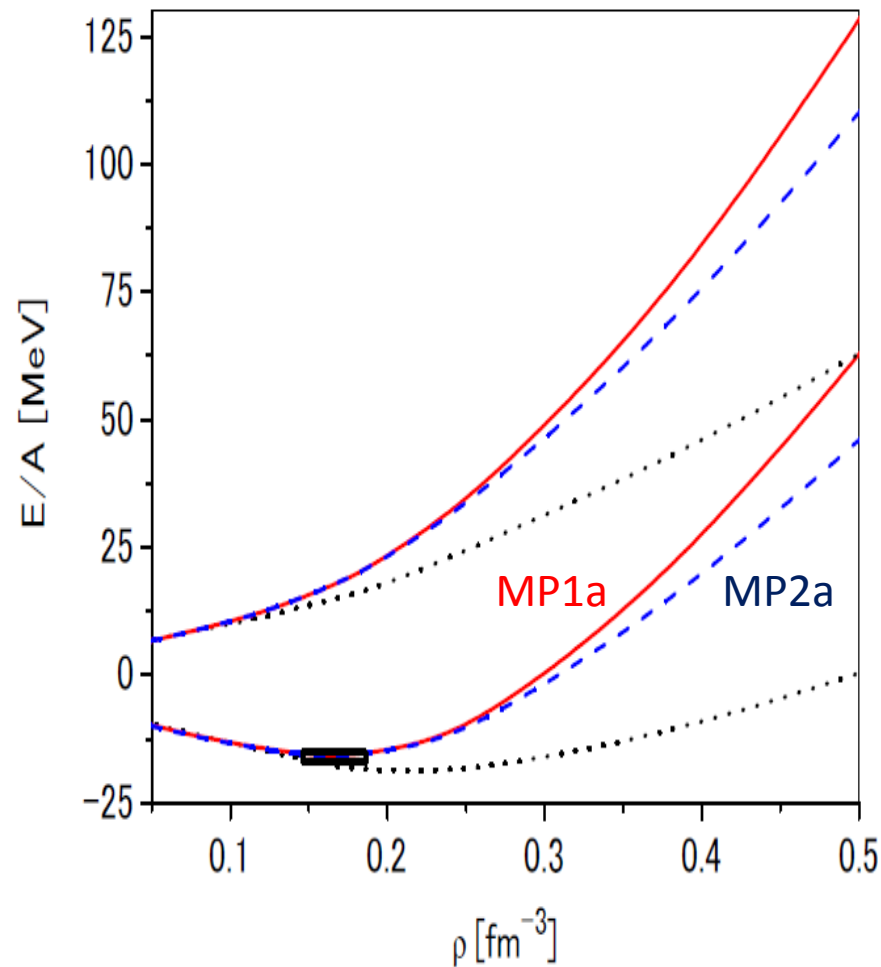
Nuclear Matter Properties are specified by

$$E_{sym}(\rho) = \frac{E}{A}(\rho, \beta = 1) - \frac{E}{A}(\rho, \beta = 0)$$

$$L = 3\rho_0 \left[\frac{\partial E_{sym}(\rho)}{\partial \rho} \right]_{\rho=\rho_0}$$

$$K = 9\rho_0^2 \left[\frac{d^2}{d\rho^2} \frac{E}{A}(\rho, \beta = 0) \right]_{\rho=\rho_0}$$

$$\beta = \frac{\rho_n - \rho_p}{\rho_n + \rho_p}$$



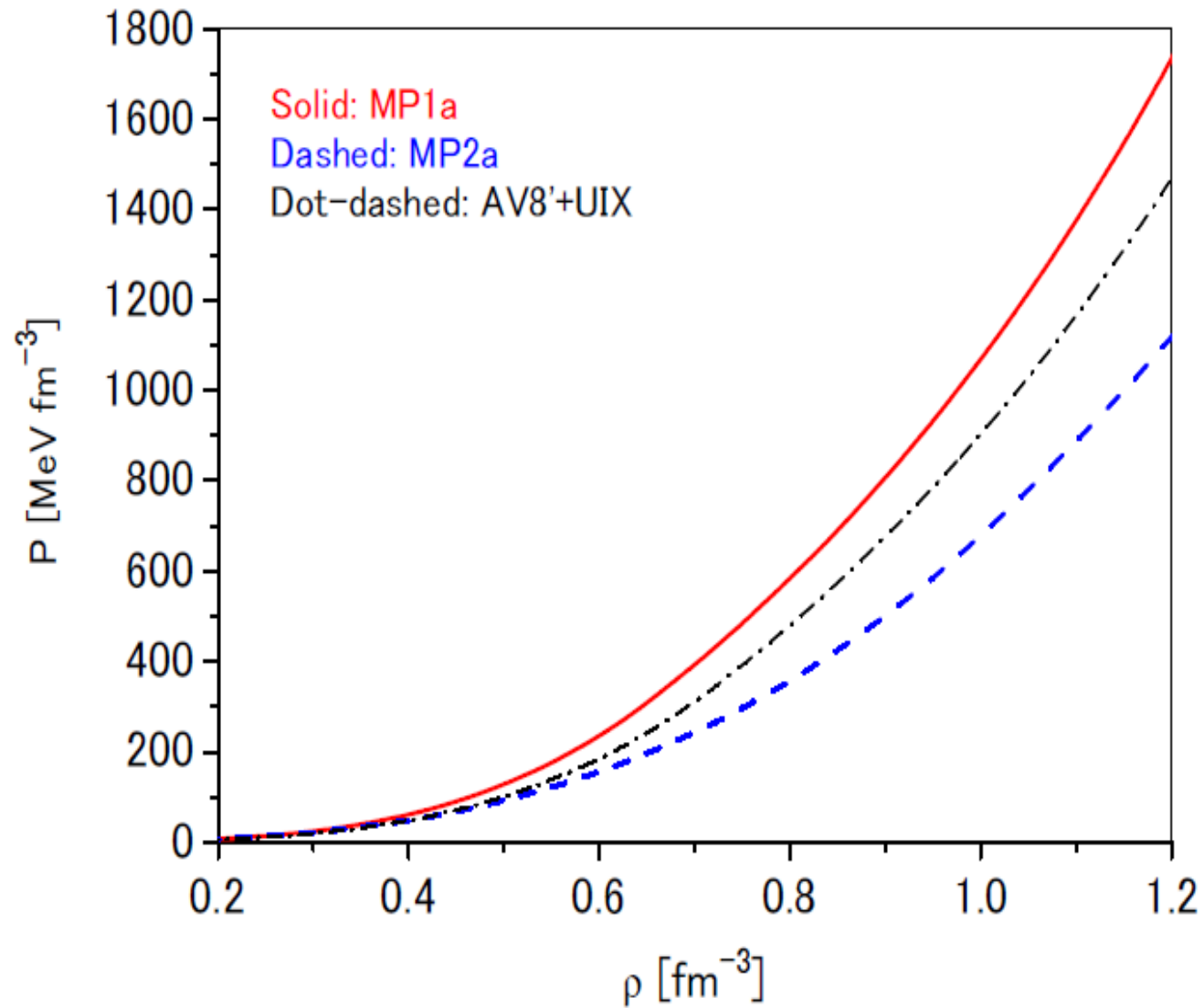
$\rho_0 = 0.164 \text{ fm}^{-3}$ $E/A(\rho_0) = -15.8 \text{ MeV}$ both for MP1a and MP2a

$E_{\text{sym}}(\rho_0) = 33.1 \text{ MeV}$ $L(\rho_0) = 70.4 \text{ MeV}$ $K = 285 \text{ MeV}$ for MP1a

$E_{\text{sym}}(\rho_0) = 33.1 \text{ MeV}$ $L(\rho_0) = 69.2 \text{ MeV}$ $K = 267 \text{ MeV}$ for MP2a

AV8' + UIX : $E_{\text{sym}} = 35.1 \text{ MeV}$ $L = 63.6 \text{ MeV}$ (Gandolfi et al.)

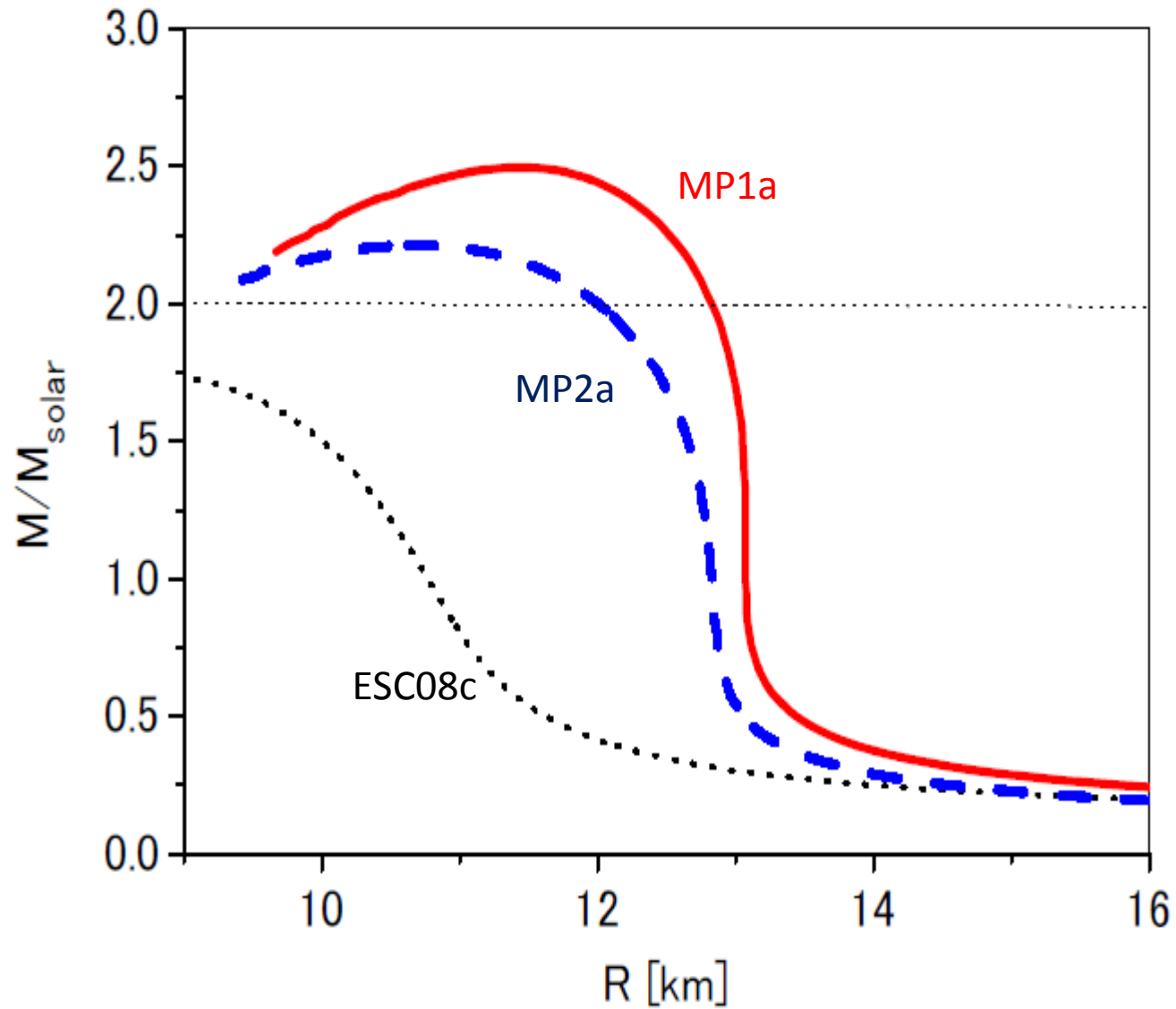
EOS (neutron matter)



Tolman-Oppenheimer-Volkoff equation

$$\frac{dP}{dr} = -\frac{GM\varepsilon}{r^2c^2} \left(1 + \frac{P}{\varepsilon}\right) \left(1 + \frac{4\pi r^3 P}{Mc^2}\right) \left(1 - \frac{2GM}{c^2 r}\right)^{-1}$$

with neutron-matter EOS



Summarizing

ESC08c + MPP + TNA

nuclear part

MPP strength determined by
analysis for $^{16}\text{O}+^{16}\text{O}$ scattering

TNA adjusted phenomenologically
To reproduce $E/A(\rho_0) = -15.8 \text{ MeV}$
with $\rho_0 = 0.164 \text{ fm}^{-3}$

No ad hoc parameter for massive neutron star (stiff EOS)
on the basis of terrestrial experiments

Hyperon-Mixed Neutron-Star Matter

ESC08c	defined in $S=0, -1, -2$ channels
MPP	universal in all BB channels
TNA	TBA ???

(ESC08c+MPP+TBA) model should be test in hypernuclei
hyperonic sector

Λ & Σ states based on
ESC08c + MPP + TBA

↑
TNA

Table 1: Values of U_Λ at normal density and partial wave contributions in $^{2S+1}L_J$ states for ESC08a/b/c from the G-matrix calculations with CON prescriptions (in MeV). The value specified by D gives the sum of $^{2S+1}D_J$ contributions. Contributions from S -state spin-spin interactions are given by $U_{\sigma\sigma} = (U(^3S_1) - 3U(^1S_0))/12$. m_Λ^* is defined by M_Λ^*/M_Λ , M_Λ^* being a Λ effective mass.

	1S_0	3S_1	1P_1	3P_0	3P_1	3P_2	D	U_Λ	$U_{\sigma\sigma}$	m_Λ^*
ESC08c	-13.6	-25.3	2.7	0.2	1.4	-3.2	-1.6	-39.4	1.29	0.73
ESC08c ⁺	-13.3	-25.4	3.4	0.4	2.2	-1.7	-2.7	-37.2	1.22	0.70
ESC08a	-13.2	-27.7	2.7	0.0	1.3	-3.7	-1.6	-42.1	1.00	0.73
ESC08b	-13.0	-25.4	2.4	-0.2	1.4	-4.3	-1.7	-40.9	1.12	0.75
NSC97e	-13.5	-29.6	1.9	0.4	2.9	-1.4	-1.3	-40.5	0.92	0.69
NSC97f	-15.0	-26.6	2.2	0.4	3.7	-0.9	-1.3	-37.5	1.54	0.68

$$\text{ESC08c}^+ = \text{ESC08c} + \text{MP1a} + \text{TBA}$$

$U_\Lambda(\rho_0)$ is conceptually different from $U_{\text{WS}}(-28 \text{ MeV})$!!
use G-matrix folding model

Y-nucleus folding potential derived from YN G-matrix interaction $G(\mathbf{r}; k_F)$

$$U_Y(\mathbf{r}, \mathbf{r}') = U_{dr} + U_{ex}$$

$$U_{dr} = \delta(\mathbf{r} - \mathbf{r}') \int d\mathbf{r}'' \rho(\mathbf{r}'') V_{dr}(|\mathbf{r} - \mathbf{r}''|; \langle k_F \rangle)$$

$$U_{ex} = \rho(\mathbf{r}, \mathbf{r}') V_{ex}(|\mathbf{r} - \mathbf{r}'|; \langle k_F \rangle)$$

G-matrix interactions

$$V_{dr} = \frac{1}{2(2t_Y + 1)(2s_Y + 1)} \sum_{TS} (2T + 1)(2S + 1) [G_{TS}^{(+)} + G_{TS}^{(-)}]$$

$$V_{ex} = \frac{1}{2(2t_Y + 1)(2s_Y + 1)} \sum_{TS} (2T + 1)(2S + 1) [G_{TS}^{(+)} - G_{TS}^{(-)}]$$

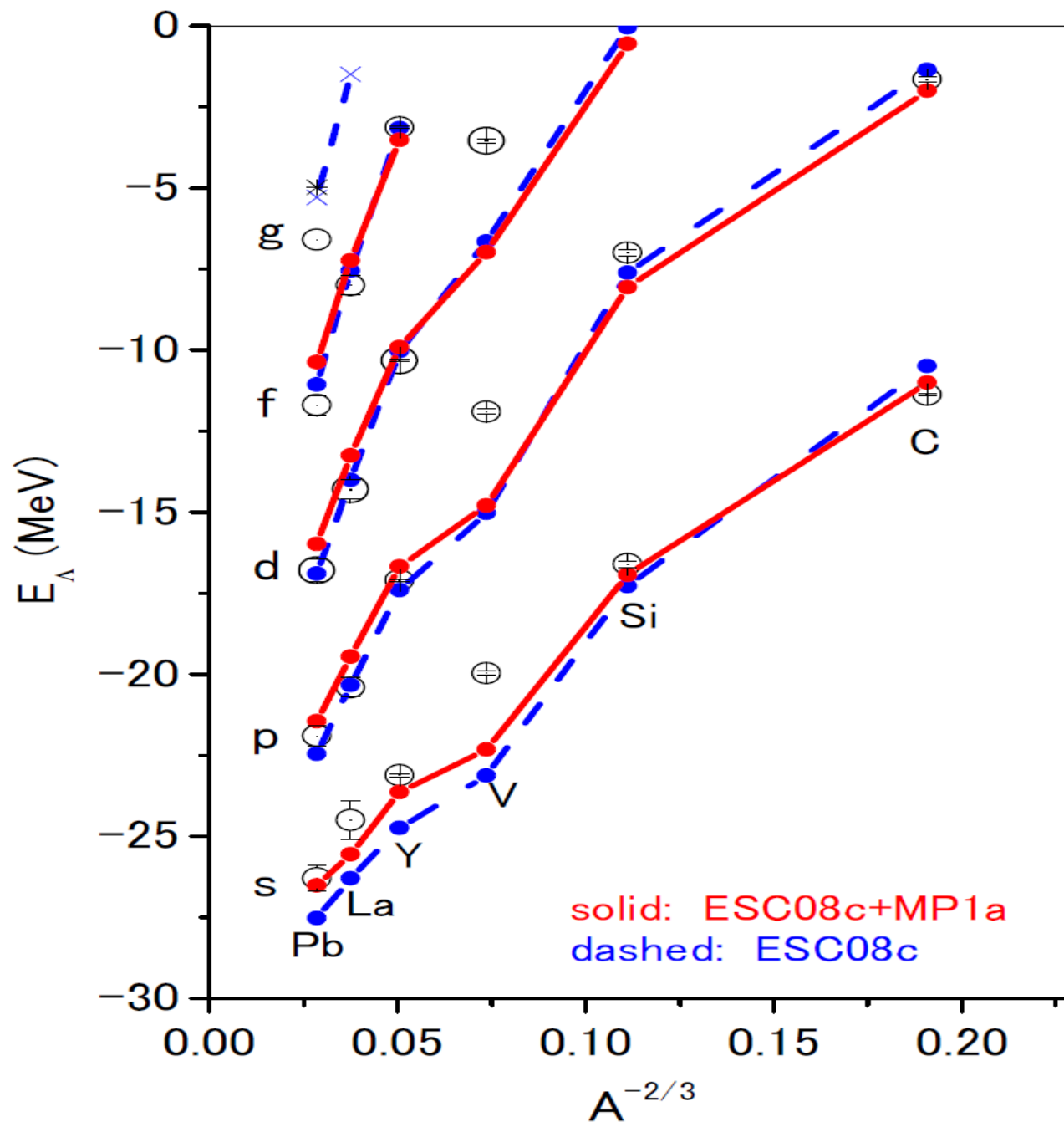
Averaged- k_F Approximation

$$\langle \rho \rangle = \langle \phi_Y(r) | \rho(r) | \phi_Y(r) \rangle$$

$$\langle k_F \rangle = (1.5\pi^2 \langle \rho \rangle)^{1/3} + \Delta k_F$$

calculated
self-consistently

Mixed density $\rho(\mathbf{r}_1, \mathbf{r}_2) = \sum_j \varphi_j^*(\mathbf{r}_1) \varphi_j(\mathbf{r}_2)$ obtained from SkHF w.f.



$\Delta k_F = -0.05$

$\Delta k_F = 0.02$

Quark-Pauli effect in ESC08 models

ESC core = pomeron + ω

Repulsive cores are similar
to each other in all channels

Assuming

“equal parts” of ESC and QM are similar to each other

Almost Pauli-forbidden states in [51] are taken into account by changing the pomeron strengths for the corresponding channels phenomenologically

$$g_p \longrightarrow \text{factor} * g_p$$

Table III. $SU(6)_{fs}$ -contents of the various potentials on the isospin, spin basis.

(S, I)	$V = aV_{[51]} + bV_{[33]}$
$(0, 1)$	$V_{NN} = \frac{4}{9}V_{[51]} + \frac{5}{9}V_{[33]}$
$(1, 0)$	$V_{NN} = \frac{4}{9}V_{[51]} + \frac{5}{9}V_{[33]}$
$(0, 1/2)$	$V_{\Lambda\Lambda} = \frac{1}{2}V_{[51]} + \frac{1}{2}V_{[33]}$
$(1, 1/2)$	$V_{\Lambda\Lambda} = \frac{1}{2}V_{[51]} + \frac{1}{2}V_{[33]}$
$(0, 1/2)$	$V_{\Sigma\Sigma} = \frac{17}{18}V_{[51]} + \frac{1}{18}V_{[33]}$
$(1, 1/2)$	$V_{\Sigma\Sigma} = \frac{1}{2}V_{[51]} + \frac{1}{2}V_{[33]}$
$(0, 3/2)$	$V_{\Sigma\Sigma} = \frac{4}{9}V_{[51]} + \frac{5}{9}V_{[33]}$
$(1, 3/2)$	$V_{\Sigma\Sigma} = \frac{8}{9}V_{[51]} + \frac{1}{9}V_{[33]}$

(S, I)	$V = aV_{[51]} + bV_{[33]}$
$(0, 0)$	$V_{\Lambda\Lambda, \Lambda\Lambda} = \frac{1}{2}V_{[51]} + \frac{1}{2}V_{[33]}$
$(0, 0)$	$V_{\Xi N, \Xi N} = \frac{1}{3}V_{[51]} + \frac{2}{3}V_{[33]}$
$(0, 0)$	$V_{\Sigma\Sigma, \Sigma\Sigma} = \frac{11}{18}V_{[51]} + \frac{7}{18}V_{[33]}$
$(0, 1)$	$V_{\Xi N, \Xi N} = \frac{7}{9}V_{[51]} + \frac{2}{9}V_{[33]}$
$(0, 0)$	$V_{\Sigma\Lambda, \Sigma\Lambda} = \frac{2}{3}V_{[51]} + \frac{1}{3}V_{[33]}$
$(0, 2)$	$V_{\Sigma\Sigma, \Sigma\Sigma} = \frac{4}{9}V_{[51]} + \frac{5}{9}V_{[33]}$
$(1, 0)$	$V_{\Xi N, \Xi N} = \frac{5}{9}V_{[51]} + \frac{4}{9}V_{[33]}$
$(1, 1)$	$V_{\Xi N, \Xi N} = \frac{17}{27}V_{[51]} + \frac{10}{27}V_{[33]}$
$(1, 1)$	$V_{\Sigma\Lambda, \Sigma\Lambda} = \frac{2}{3}V_{[51]} + \frac{1}{3}V_{[33]}$
$(1, 1)$	$V_{\Sigma\Sigma, \Sigma\Sigma} = \frac{16}{27}V_{[51]} + \frac{11}{27}V_{[33]}$

Pauli-forbidden state in $V_{[51]} \rightarrow$ strengthen pomeron coupling



$$V_{BB} = V(\text{pom}) + w_{BB}[51] * \underline{V(\text{PB})}$$

Table 1: Values of U_Σ at normal density and partial wave contributions for ESC08c models (in MeV).

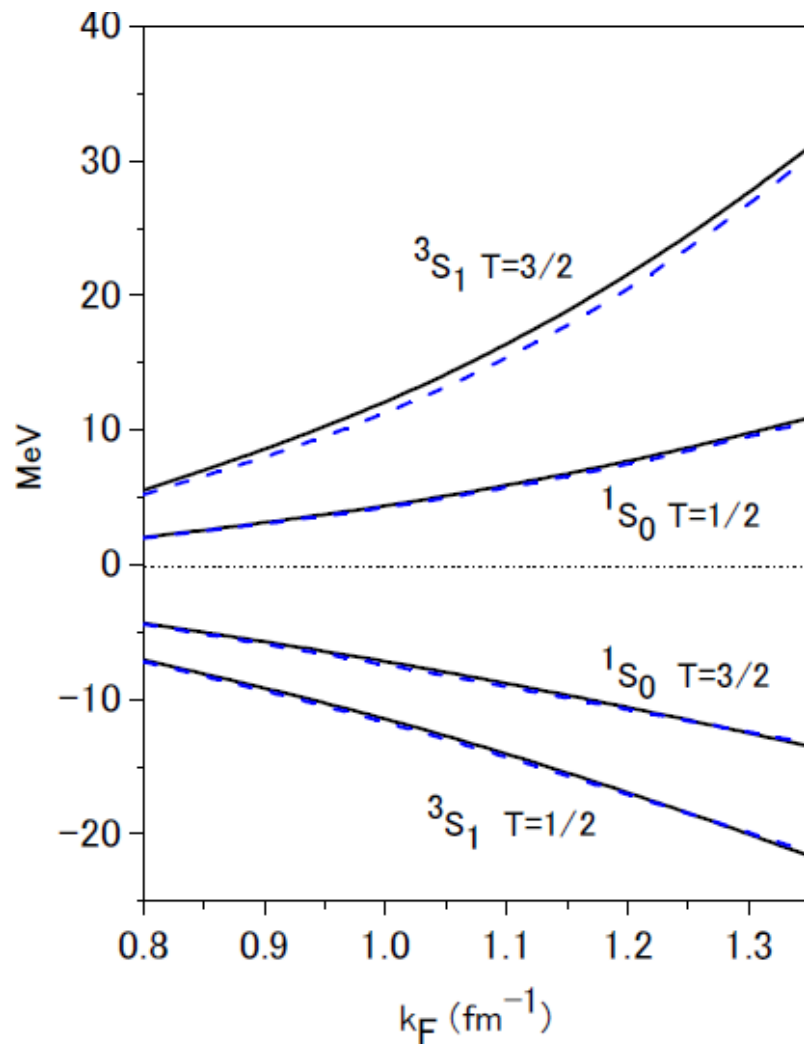
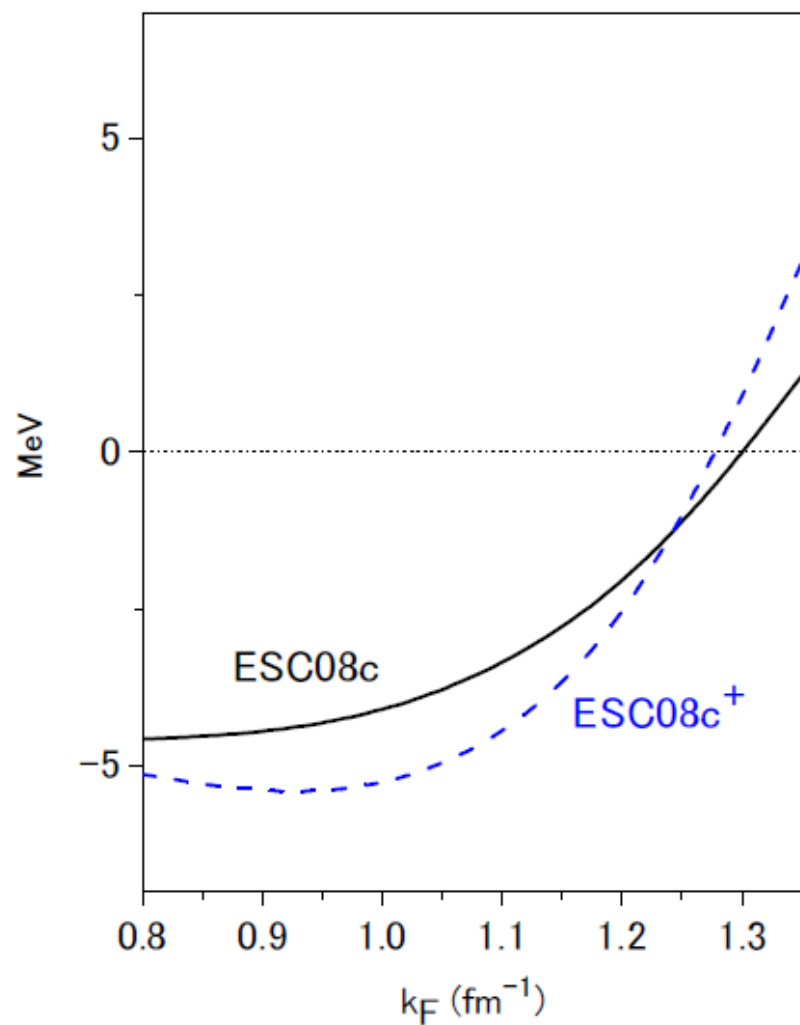
model	T	1S_0	3S_1	1P_1	3P_0	3P_1	3P_2	D	U_Σ
ESC08c	1/2	10.9	-21.6	2.4	2.1	-6.0	-1.0	-0.7	+1.2
	3/2	-13.5	31.0	-4.7	-1.8	5.9	-1.5	-0.2	

Pauli-forbidden state in QCM → strong repulsion in $T=3/2$ 3S_1 state



Σ^- in neutron matter

$$U_{\Sigma}(k_F)$$



Solid $\longrightarrow$ Dashed : Contributions from MPP+TBA

Hyperon-mixed Neutron-Star matter with universal TBR (MPP)

EoS of $n+p+\Lambda+\Sigma+e+\mu$ system

ESC08c (YN) + MPP (YNN) + TBA (YNN)



Preliminary

Hyperon-mixed neutron matter

Starting from single particle potentials calculated with the G-matrix theory:

$$U_B(k) = \sum_{B'} U_B^{(B')}(k) \quad \text{with } B, B' = n, p, \Lambda, \Sigma^-$$

$U_B^{(B')}$ means a single particle potential of B particle in B' matter

$$\begin{aligned} \varepsilon &= \varepsilon_{mass} + \varepsilon_{kin} + \varepsilon_{pot} \\ &= 2 \sum_B \int_0^{k_F^B} \frac{d^3k}{(2\pi)^3} \left[M_B - M_n + \frac{\hbar^2 k^2}{2M_B} + \frac{1}{2} U_B(k) \right] \\ \varepsilon_{mass} &= \sum_B (M_B - M_n) \rho_B \\ \varepsilon_{kin} &= \sum_B \frac{3}{5} \frac{\hbar^2 (k_F^B)^2}{2M_B} \rho_B = \sum_B \frac{3}{5} \frac{\hbar^2}{2M_B} (3\pi^2)^{2/3} (\rho_B)^{5/3} \\ \varepsilon_{pot} &= 2 \sum_B \int_0^{k_F^B} \frac{d^3k}{(2\pi)^3} \frac{1}{2} U_B(k) = \frac{1}{2} \sum_B \int_0^{k_F^B} \frac{k^2 dk}{\pi^2} U_B(k) \end{aligned}$$

$$\varepsilon_{pot} = \frac{1}{2} \sum_{BB'} \mathcal{U}_B^{(B')}$$

$$\mathcal{U}_B^{(B')} = \int_0^{k_F^B} \frac{k^2 dk}{\pi^2} U_B^{(B')}(k) \quad \mathcal{U}_B^{(B')} = \mathcal{U}_{B'}^{(B)}$$

$$\int_0^{k_F^B} \frac{k^2 dk}{\pi^2} U_B^{(B')}(k) = \int_0^{k_F^{B'}} \frac{k^2 dk}{\pi^2} U_{B'}^{(B)}(k)$$

$$\frac{\partial}{\partial \rho_B} \mathcal{U}_B^{(B')} = U_B^{(B')}(k_F^B) + \int_0^{k_F^B} \frac{k^2 dk}{\pi^2} \frac{\partial U_B^{(B')}(k)}{\partial \rho_B}$$

$$\text{Chemical potential : } \mu_B = \frac{\partial \varepsilon}{\partial \rho_B}$$

Chemical potential : $\mu_B = \frac{\partial \varepsilon}{\partial \rho_B}$

Chemical equilibrium conditions:

$$\mu_n = \mu_p + \mu_e$$

$$\mu_e = \mu_\mu$$

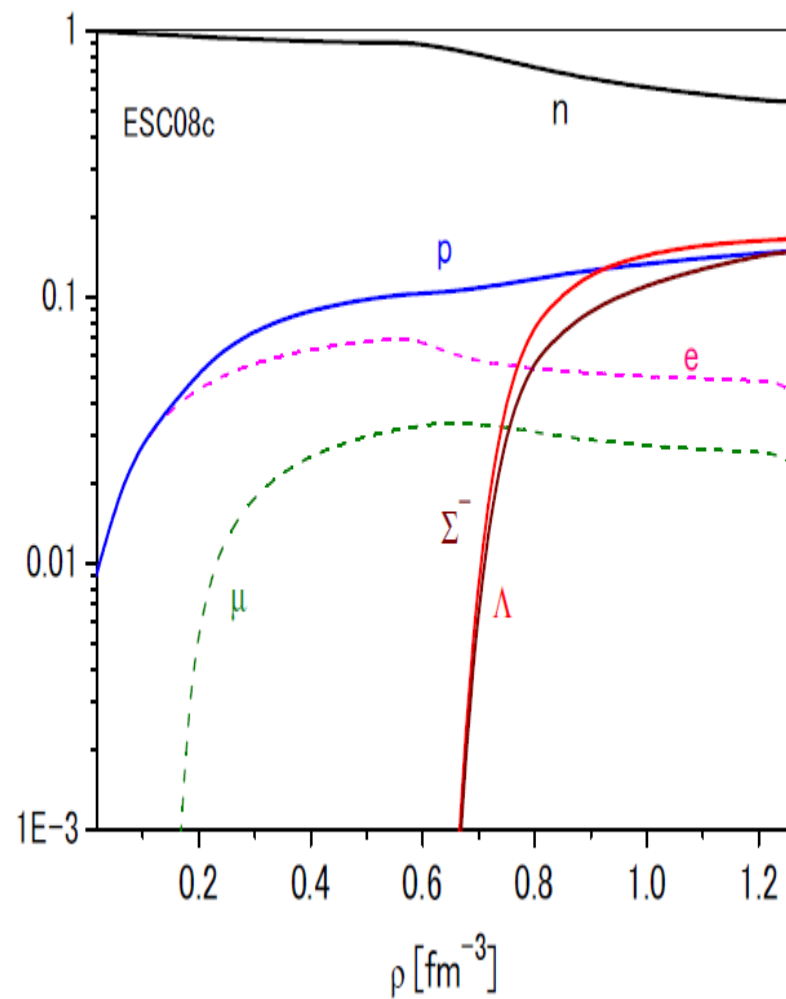
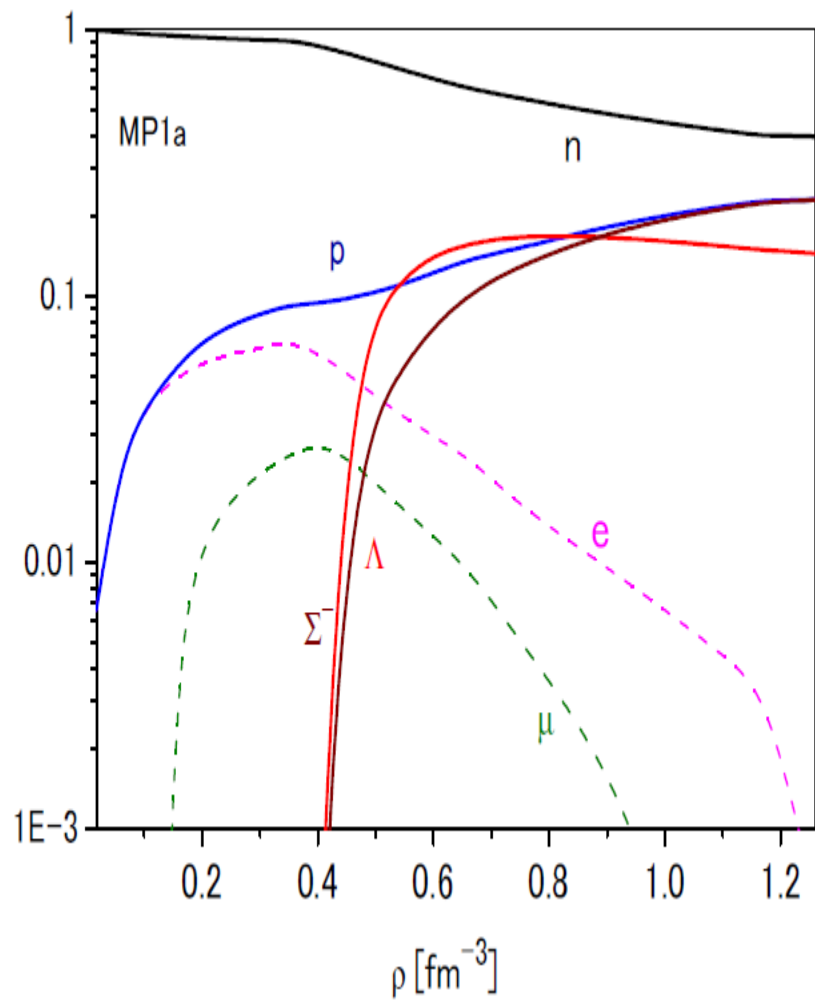
$$\mu_{\Sigma^-} = \mu_n + \mu_e$$

$$\mu_\Lambda = \mu_n$$

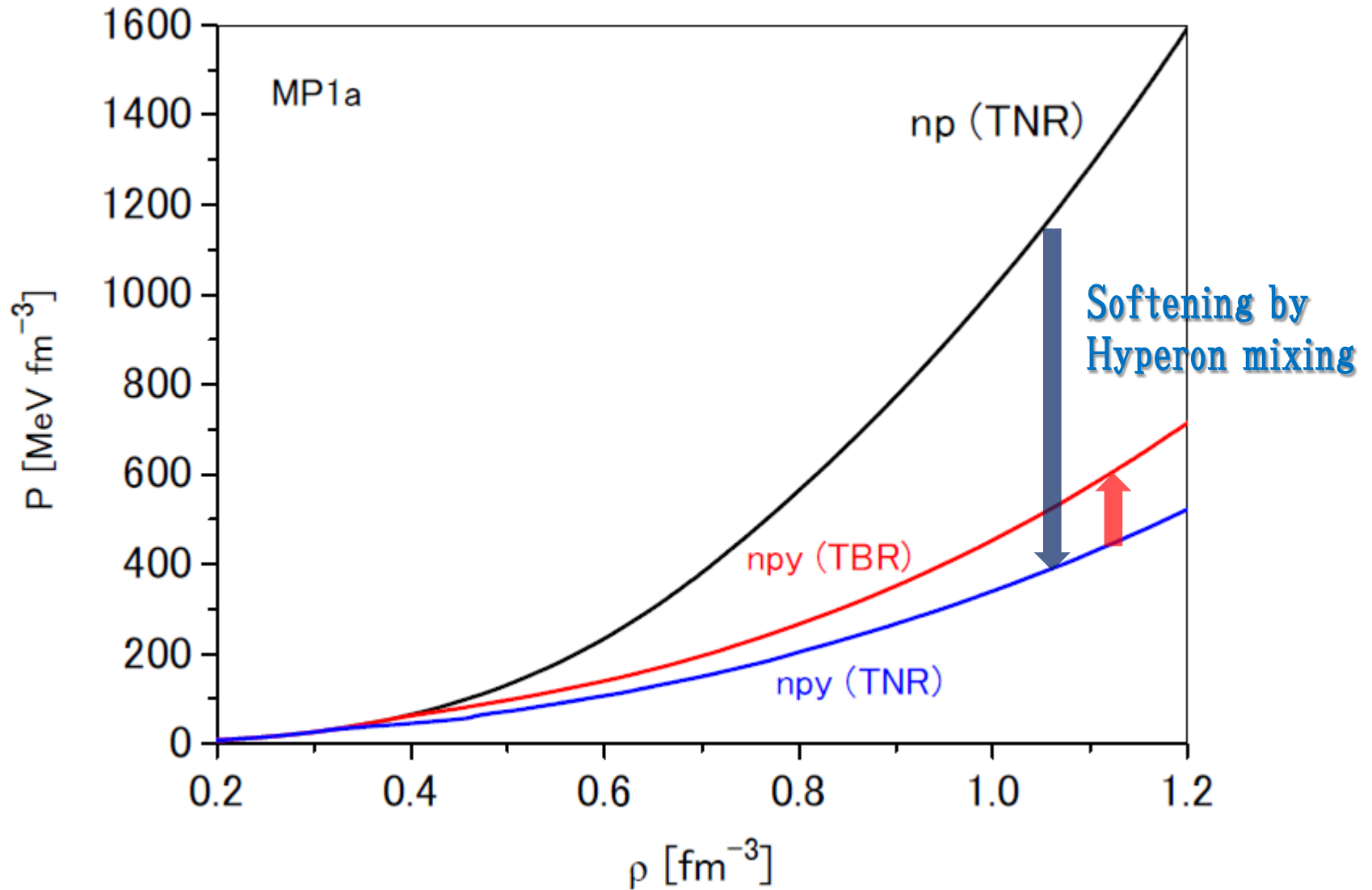
Baryon-number conservation : $y_n + y_p + y_\Lambda + y_{\Sigma^-} = 1$

Charge neutrality : $y_p = y_{\Sigma^-} + y_e + y_\mu$

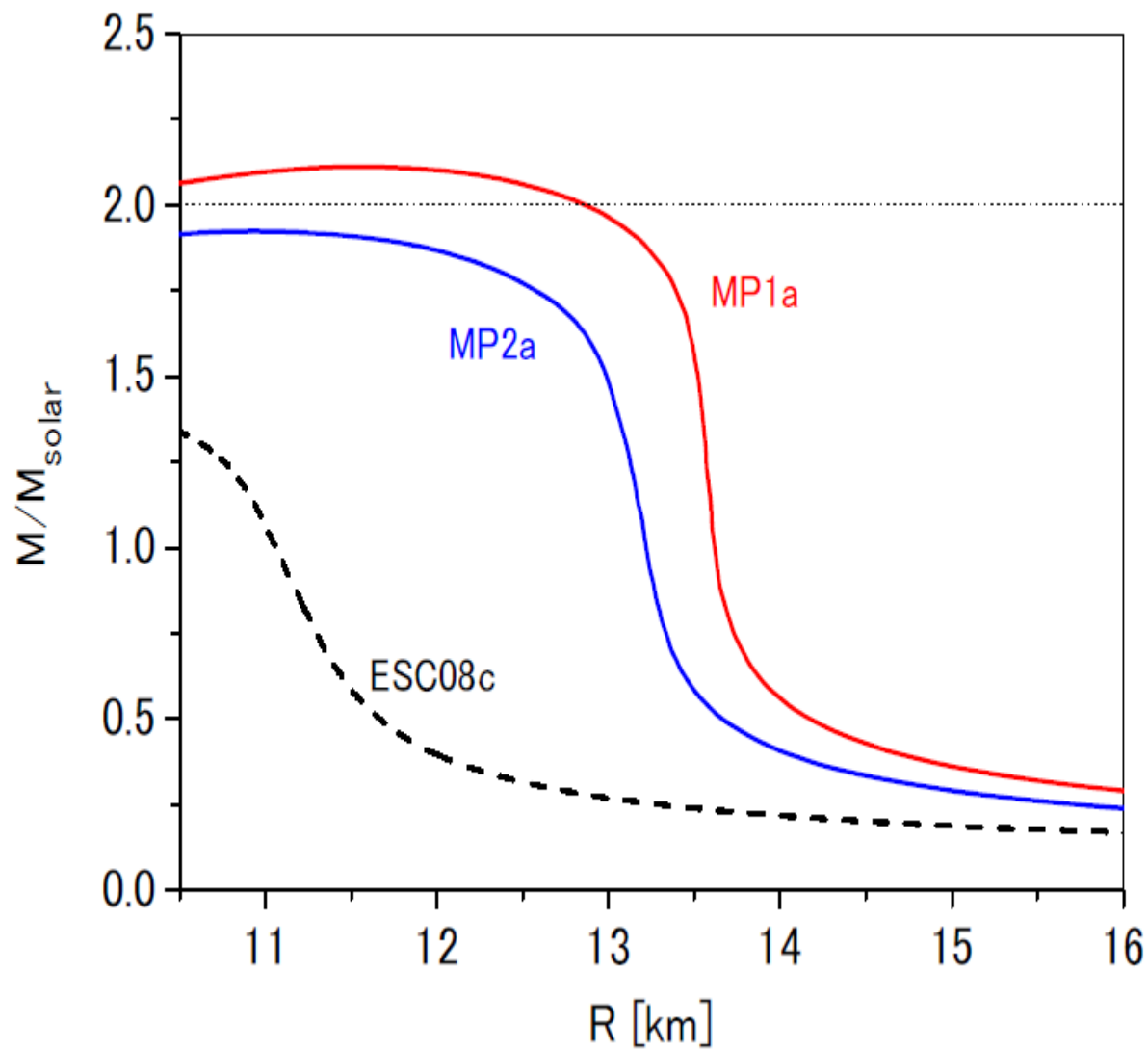
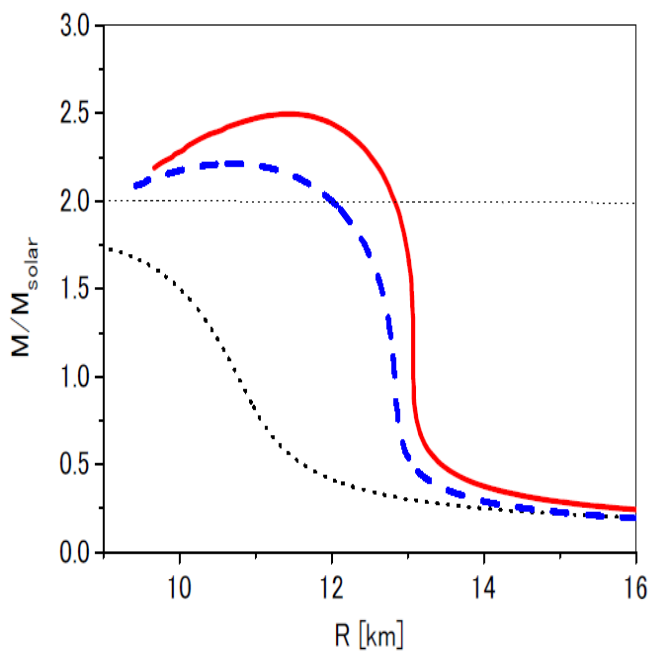
β -stable $n+p+\Lambda+\Sigma^-$ matter



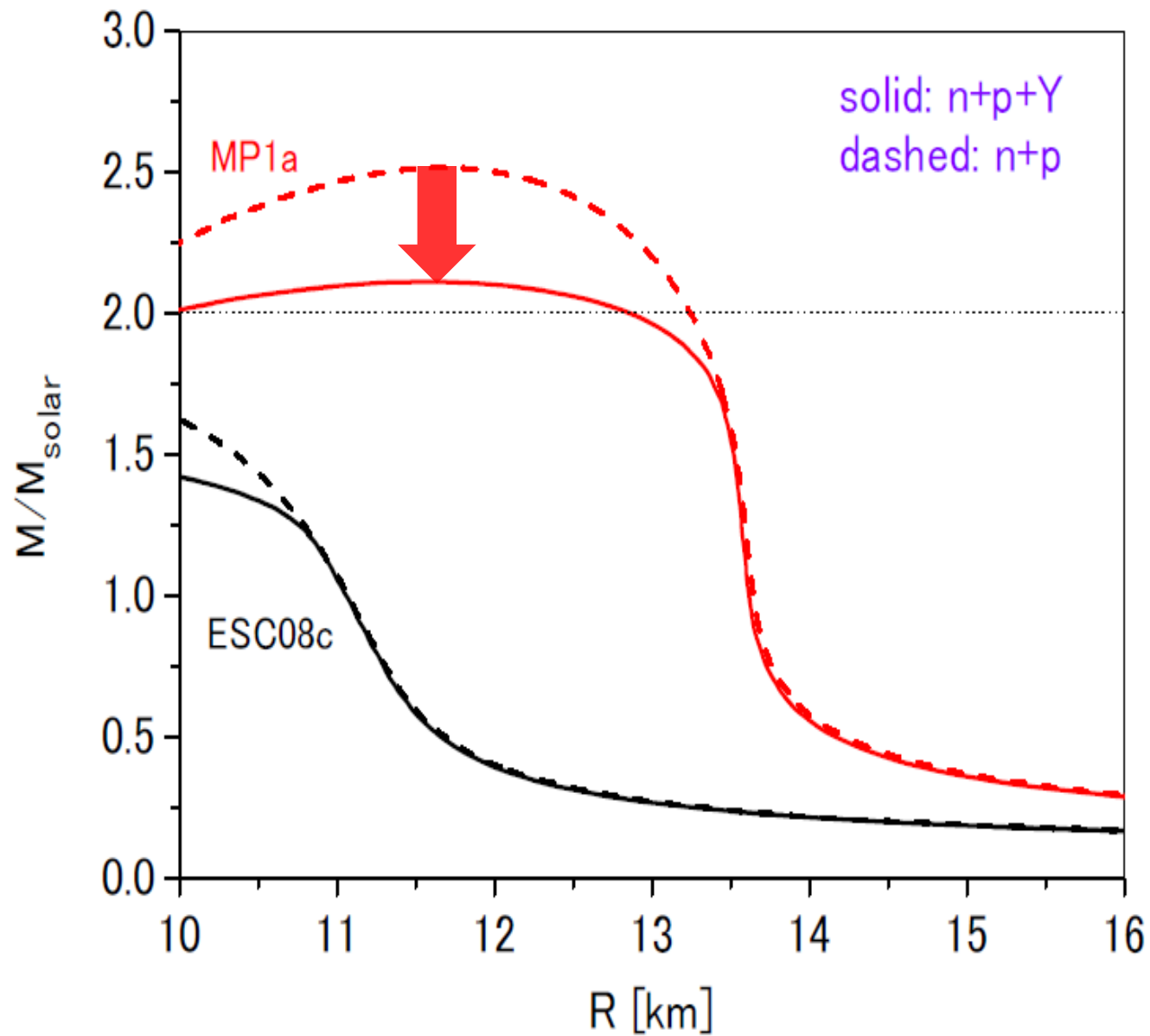
EOS



with EOS of $n+p+\Lambda+\Sigma^-$ matter



Softening by hyperon mixing



Conclusion

ESC08c+MPP+TBA model

- * MPP strength determined by analysis for $^{16}\text{O}+^{16}\text{O}$ scattering
- * TNA adjusted phenomenologically to reproduce
 $E/A(\rho_0) = -15.8 \text{ MeV}$ with $\rho_0 = 0.164 \text{ fm}^{-3}$
- * Consistent with hypernuclear data
- * No ad hoc parameter to stiffen EOS
BB interactions based on on-Earth experiments

MP1a set including 3- and 4-body repulsions leads to massive neutron stars with $2M_\odot$ in spite of significant softening of EOS by hyperon mixing