

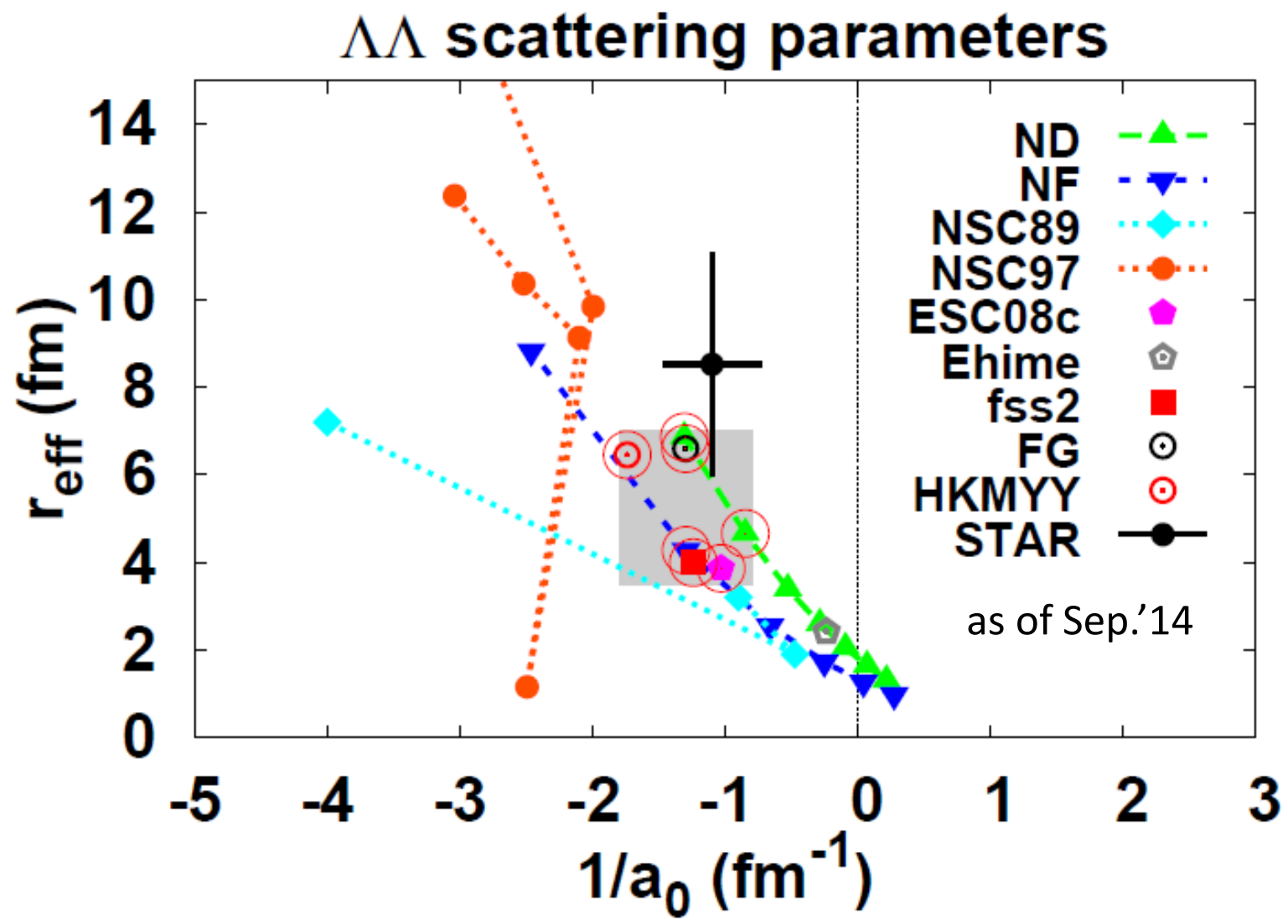
RHICにおける $\Lambda\Lambda$ 相互作用

森田健司
(京大基研)

共同研究者
大西明(京大基研)
古本猛憲(一関高専)

Ref.) Phys. Rev. C91, 024916 (2015)

散乱長 a_0 と有効レンジ r_{eff} への制限

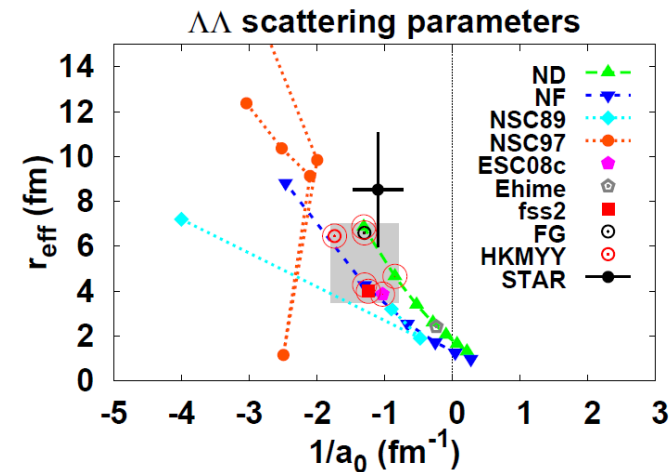


Outline

重イオン衝突 (RHIC, LHC) → $\Lambda\Lambda$ 相互作用

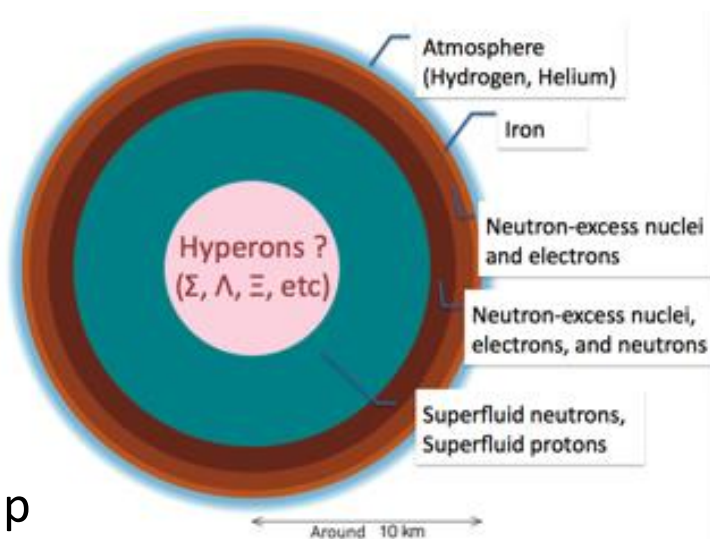
1. メカニズム
2. 重イオン衝突のダイナミクスの効果 (reminder)

3. $\Sigma^0 \rightarrow \Lambda \gamma$ の影響
4. 相関関数のロングテール
5. STARとの違い
6. まとめと展望



$\Lambda\Lambda$ 相互作用の重要性

■ 中性子星におけるハイペロン混合



状態方程式の軟化

観測との整合性？

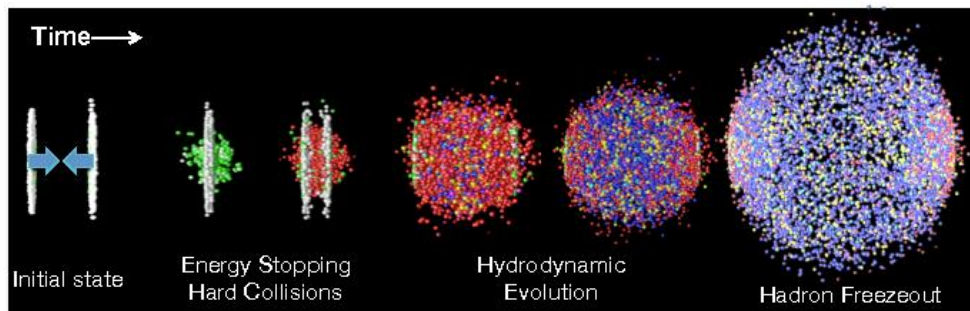
相互作用の影響は？

■ “ H ”-ダイバリオン

● Λ - Λ 束縛状態？ (Lattice・重い π)

● 共鳴状態があるかも？

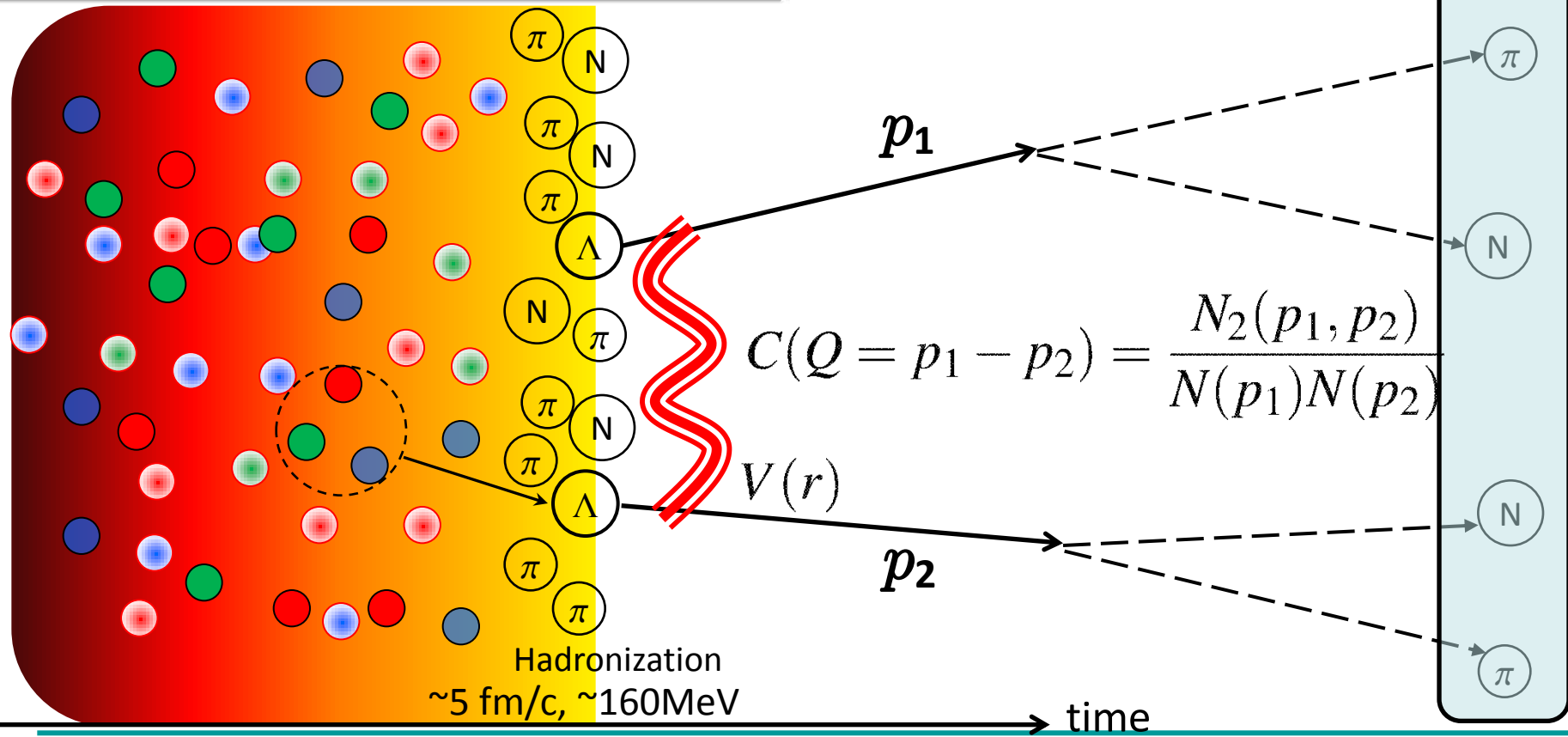
重イオン衝突での $\Lambda\Lambda$ 生成・相関



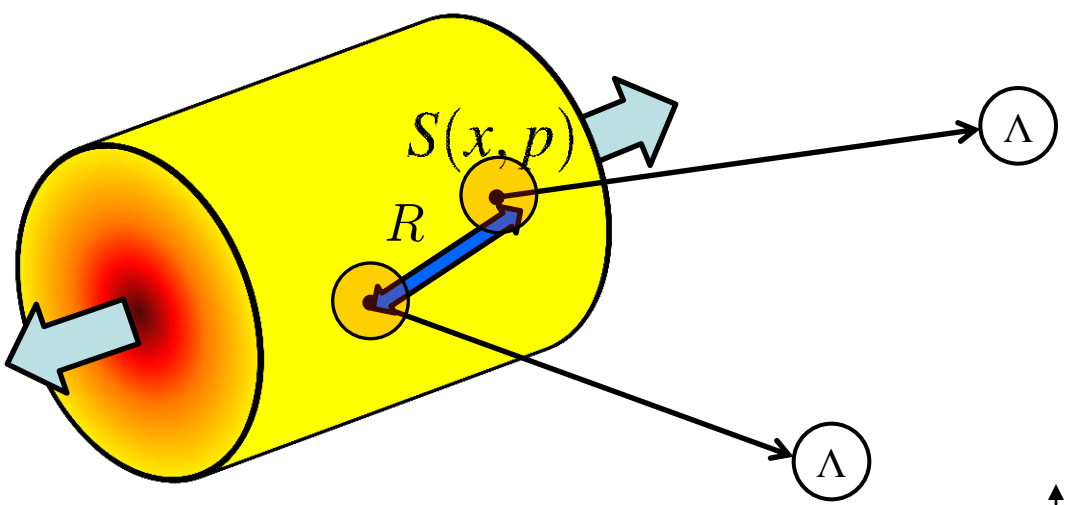
Au+Au 200A GeV

$$\frac{dN_{\Lambda}}{dy} \simeq 0.6 - 13 \quad \times 10^8 \text{ events}$$

(60-80%) (0-5%)



$\Lambda\Lambda$ 相関関数



熱的粒子源：ランダム位相

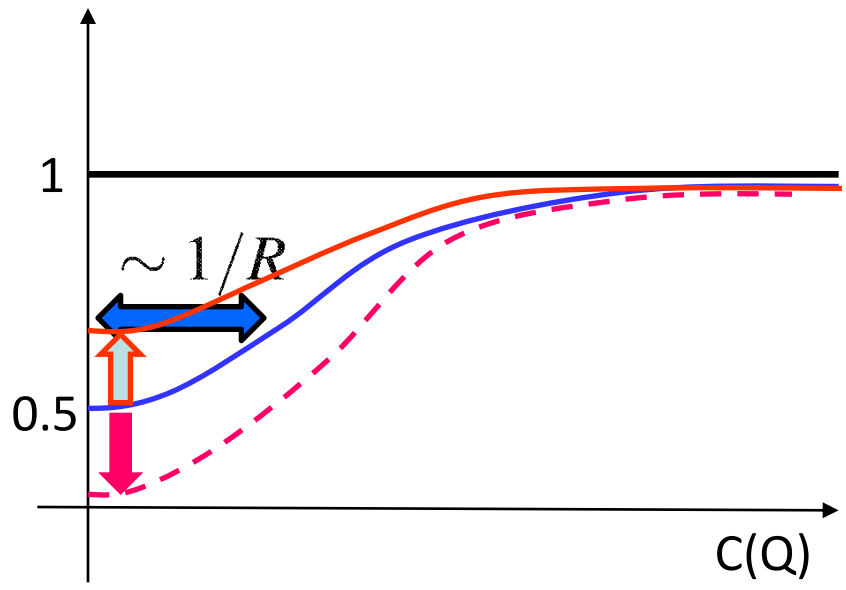
同種粒子起因の相関 (HBT)

$c(Q)$: 粒子源サイズ

$\Lambda\Lambda$ 相互作用 : No Coulomb!

$c(Q)$ への影響 : effective range
 $r_{\text{eff}} \sim R$

repulsive / attractive interaction
での違い



手法：熱的粒子源モデル + $V_{\Lambda\Lambda}$

Formula (Gong et al., '91)

$$C_2(Q, K) = \frac{\int d^4x_1 d^4x_2 S(x_1, K) S(x_2, K) |\Psi_{12}(Q, x_1 - x_2 - (t_2 - t_1)K/m)|^2}{\int d^4x_1 d^4x_2 S(x_1, k_1) S(x_1, k_2)}$$

ソース関数 (放出点分布)



熱的粒子源モデル

(理想：流体シミュレーション)

- ◆ 静的・球対称
- ◆ 球対称 + 膨張
- ◆ 円筒対称 + 縦ブースト不変性 + 横膨張

$\Lambda\Lambda$ 相対波動関数 (S-wave)



ポテンシャル (2 or 3-range Gaussian Fit)

Meson exchange models (Nijmegen model D, F, Soft Core89/97, ESC08)

Phenomenological (Ehime)

Quark model (fss2)

Fit to $_{\Lambda\Lambda}^6\text{He}$ (Nagara) Filikhin-Gal (FG)
Hiyama et al. (HKMYY)

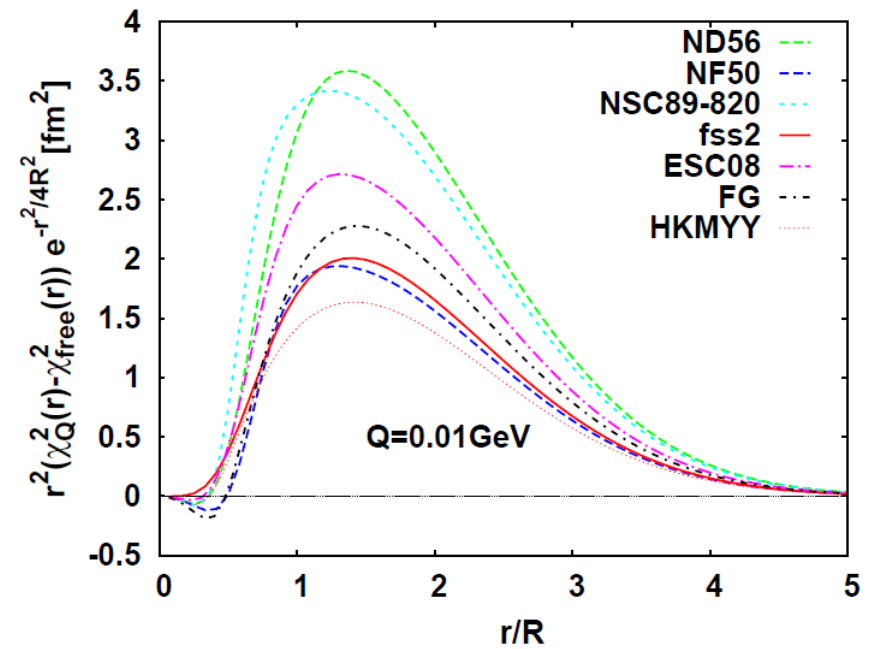
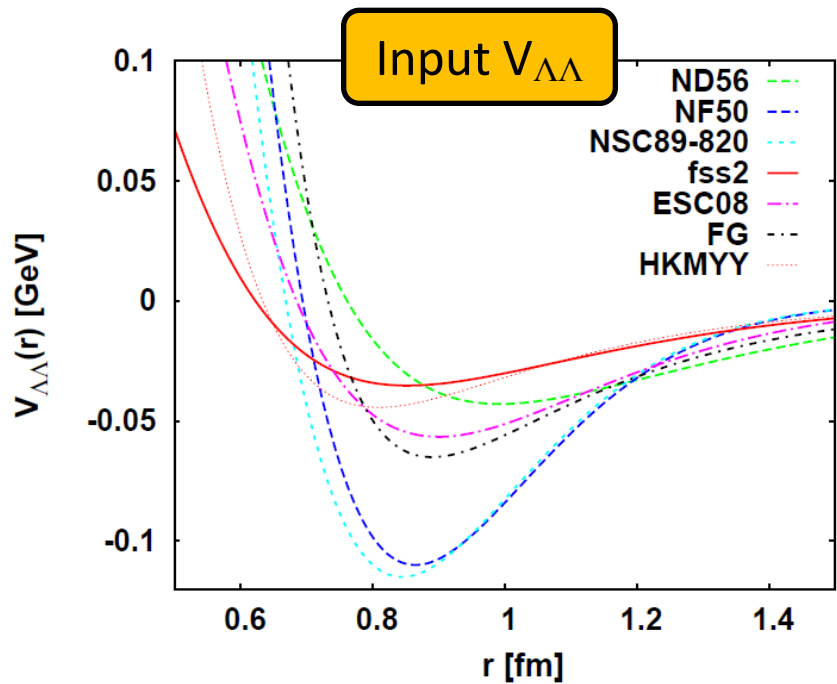
ポテンシャルと波動関数

球対称静的ソース

$$C_{\text{stat}}(Q) = 1 - \frac{1}{2}e^{-Q^2 R^2} + \frac{1}{4\sqrt{\pi}R^3} \int_0^\infty dr r^2 e^{-\frac{r^2}{4R^2}} \underbrace{[\chi_Q(r)]^2 - [j_0(Qr/2)]^2}_{\text{Interaction : deviation from free w.f.}}$$

HBT (size R)

S-wave w.f.

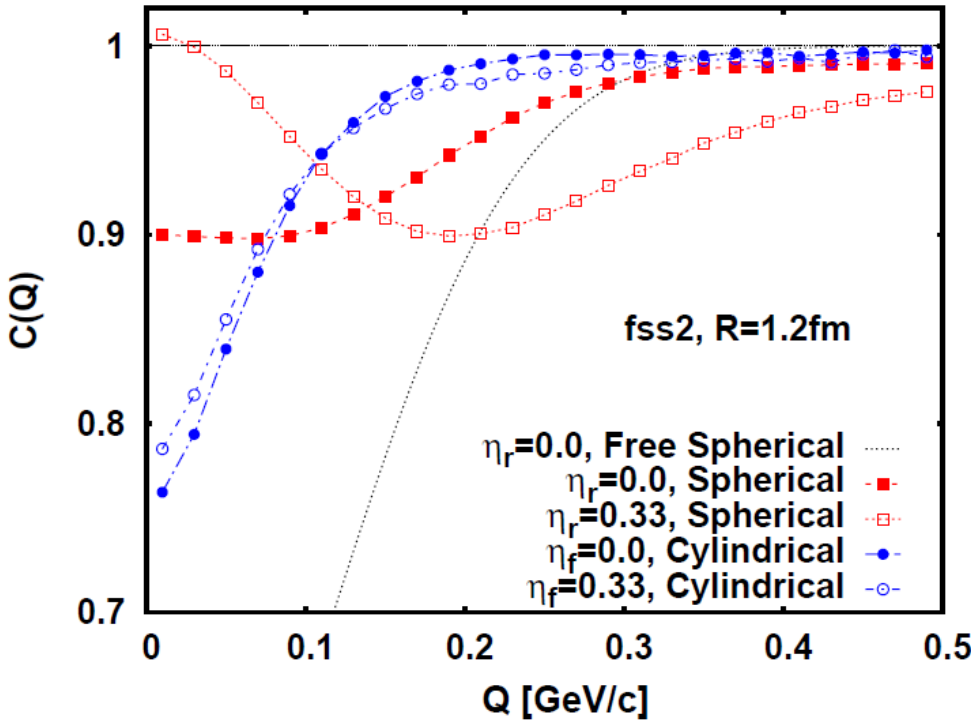


膨張粒子源モデル

S.Chapman et al., '95

$$S(x, k) \propto m_T \cosh(y - Y_L) n_F(u \cdot k / T) \exp \left[-\frac{(\tau - \tau_0)}{2(\Delta\tau)^2} - \frac{x^2 + y^2}{2R^2} \right]$$
$$\sim \exp \left[-\frac{\gamma_T M_T}{T} \cosh(y - Y_L) \right] \exp \left[\frac{\gamma v_T k_T \cos \phi}{T} \right]$$

ブースト不変でも縦に有限サイズ $R_L \sim \tau_0 \sqrt{\frac{T}{M_T}}$

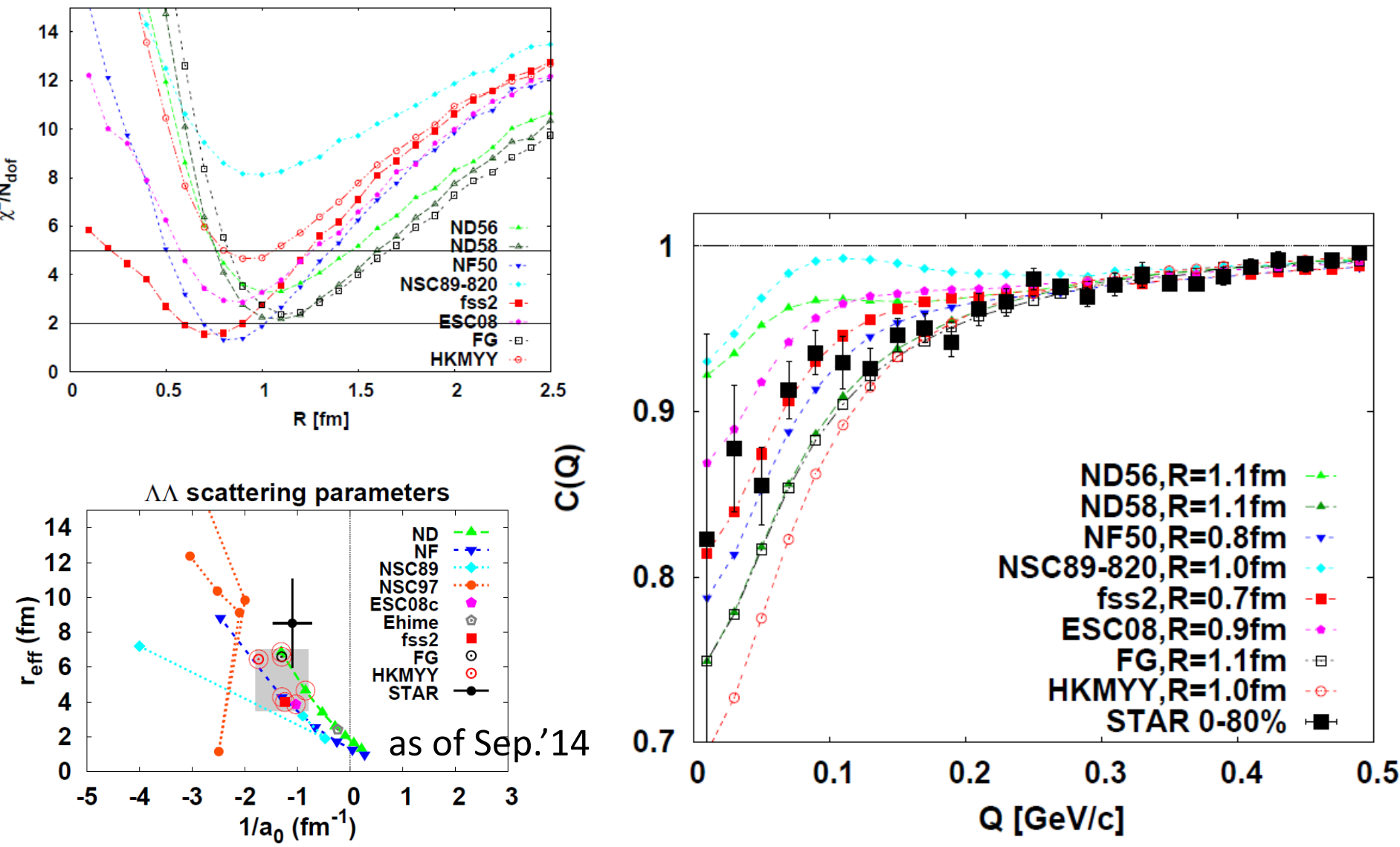


Width of C(Q) (Q=|p₁-p₂|) : effective 3d size

Large R_L : narrow width

Transverse flow: fit to p_t distribution of Λ (STAR '12)

$V_{\Lambda\Lambda}$: 膨張粒子源モデルによるFit



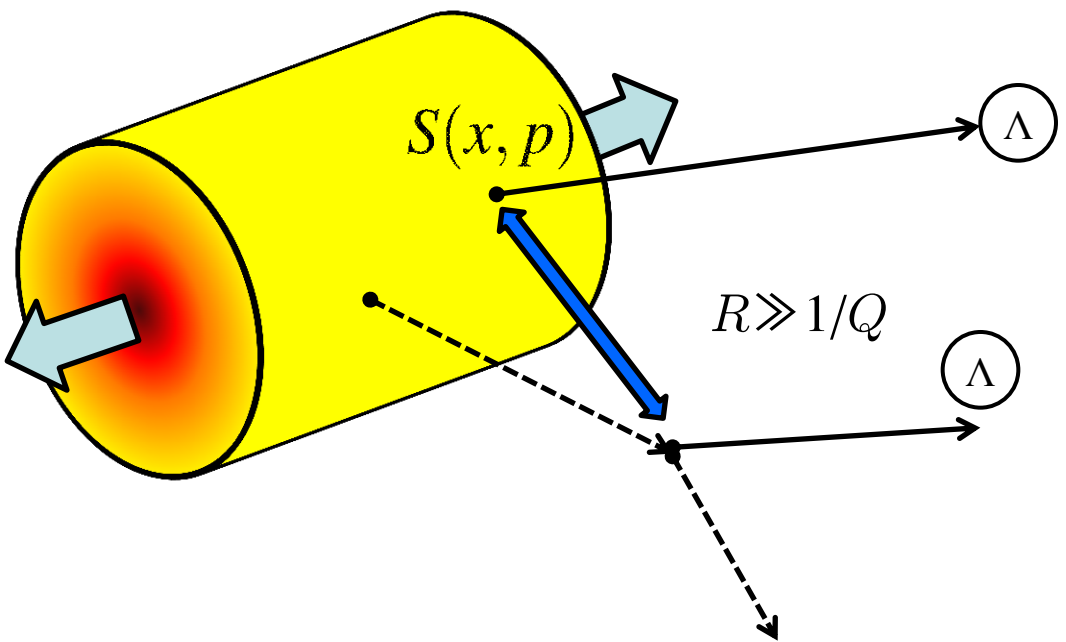
Feed-Down の寄与

■ $\Xi \rightarrow \Lambda + \pi$: subtracted

● DCA < 0.4cm from the primary vertex

■ $\Sigma^0 \rightarrow \Lambda + \gamma$

$$C(Q) \rightarrow 1 + \left(\frac{\Lambda^{\text{dir}}}{\Lambda^{\text{tot}}} \right)^2 (C(Q) - 1)$$



$\uparrow$
 $(0.67)^2$

熱的生成とconsistent

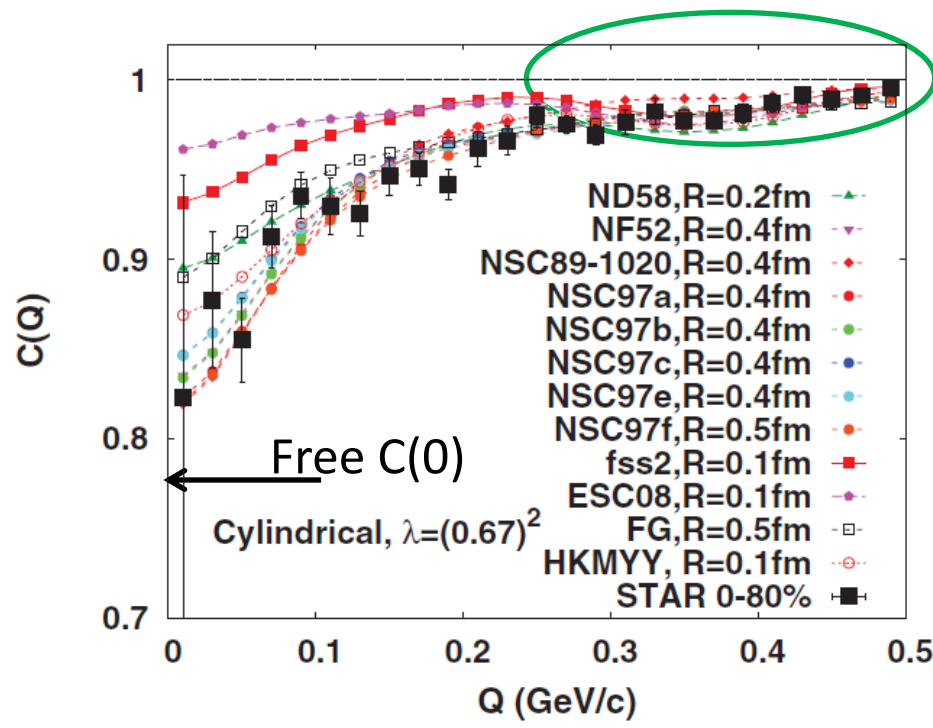
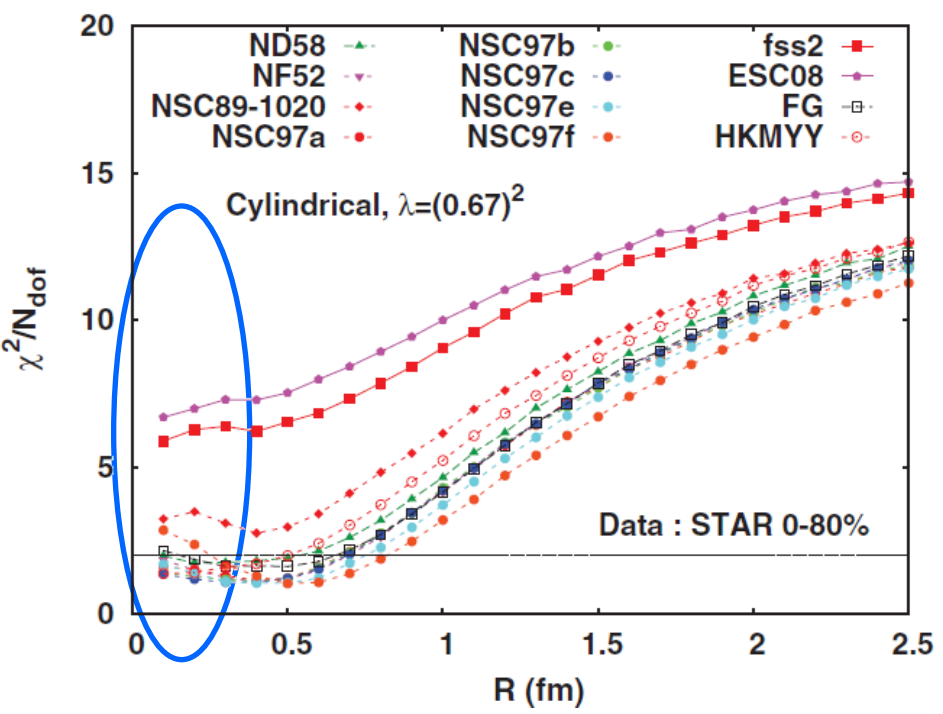
$$\Sigma^0/\Lambda = 0.278 \text{ (p+Be data)}, \Xi/\Lambda = 0.85 \text{ (RHIC)} \rightarrow \Lambda^{\text{dir}}/\Lambda^{\text{tot}} = 0.67$$

(0.52 if including Ξ)

Feed-Down の寄与

$$C(Q) \rightarrow 1 + (0.67)^2 (C(Q) - 1)$$

Low Qでの相互作用依存性→弱くなる



非常に小さいサイズ？

long tail を合わせるのに必要。。。。

ロングテールを記述する相関？

別起源の相関

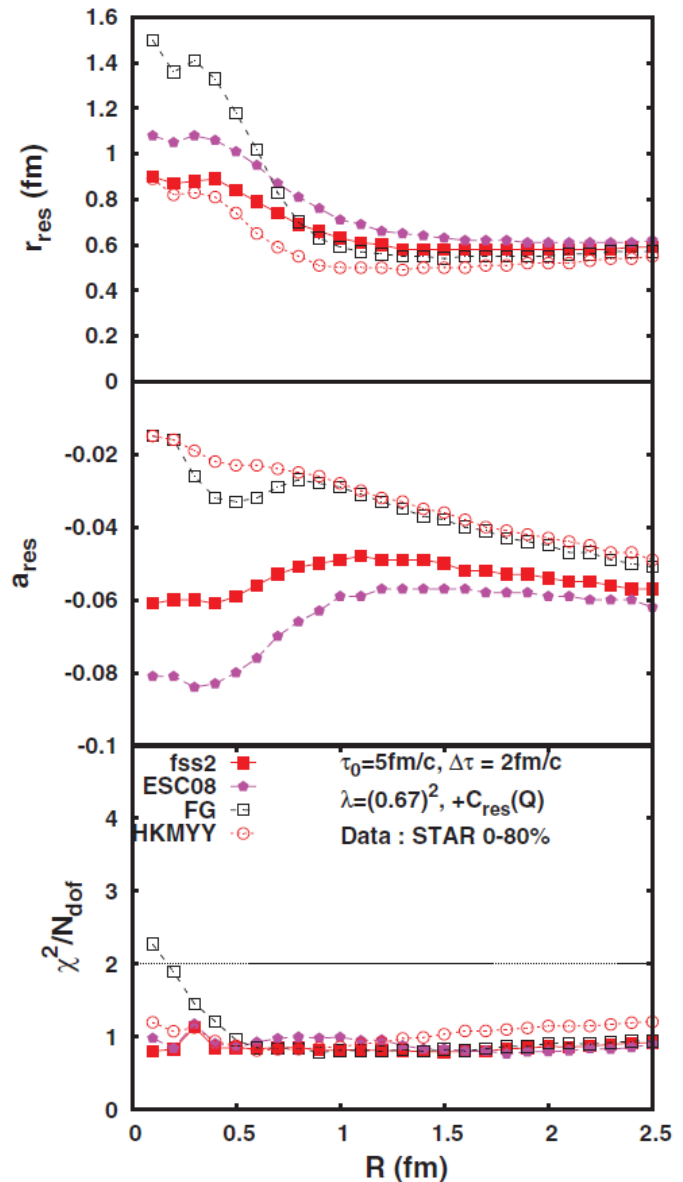
$$C(Q) \rightarrow C(Q) + a_{\text{res}} e^{-r_{\text{res}}^2 Q^2}$$

それぞれのRについて
($a_{\text{res}}, r_{\text{res}}$) で $\min \chi^2$ 探索

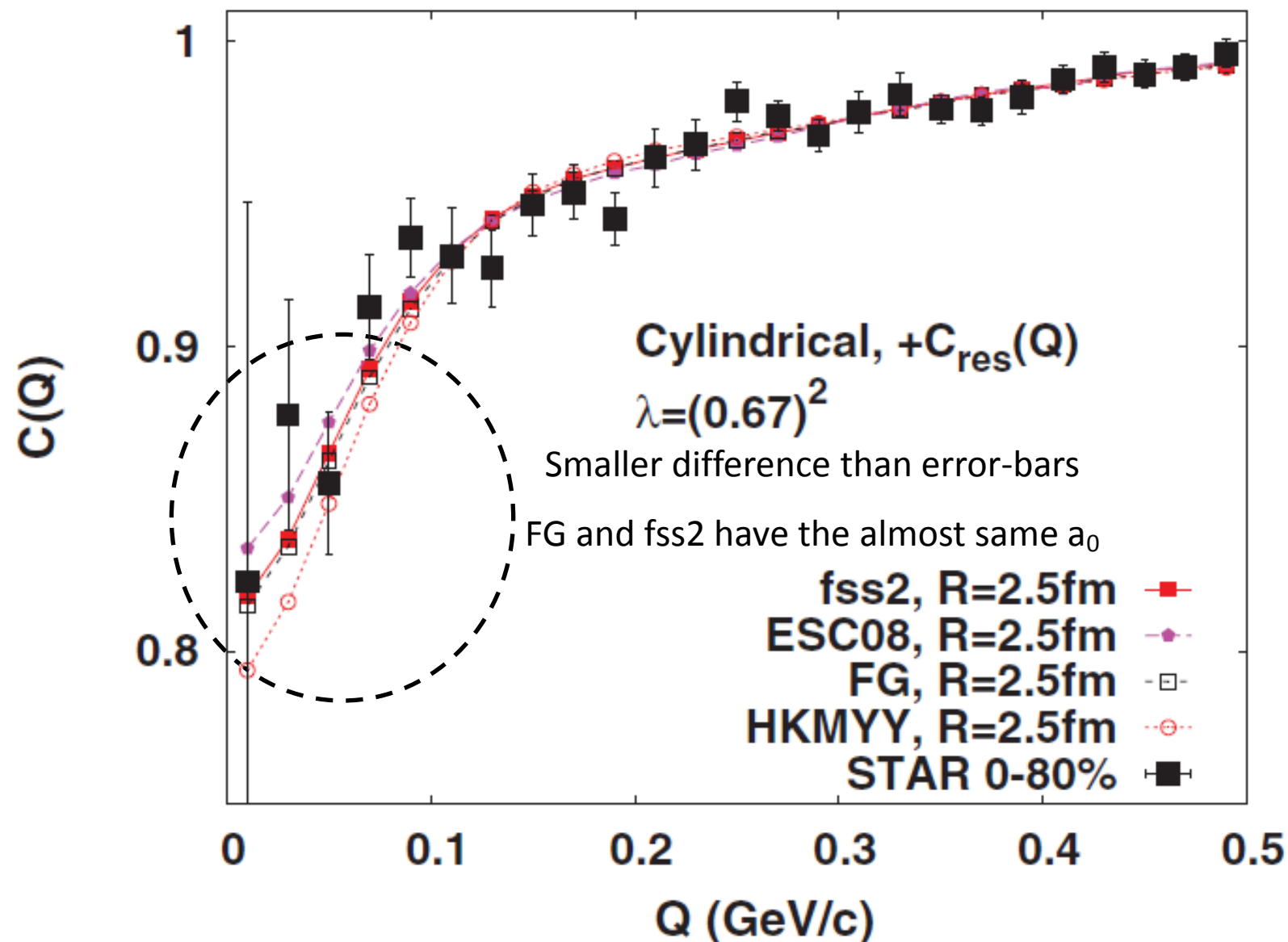
Rによらず $\chi^2/N_{\text{dof}} \sim 1$

Small R : Large $r_{\text{res}} > R$

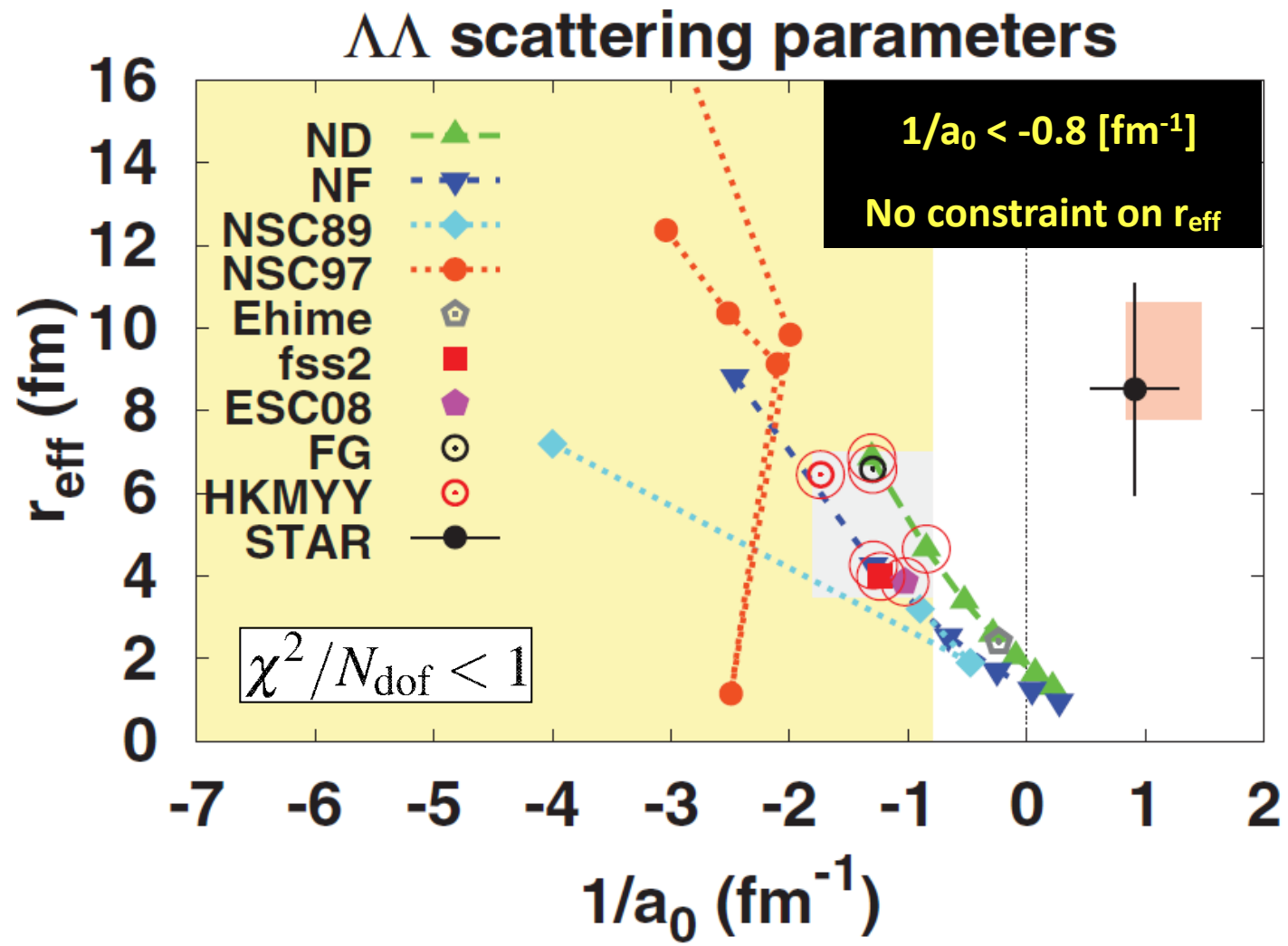
Low Q behavior $\longleftrightarrow$ $\Lambda\Lambda$ Interaction



Fit結果例

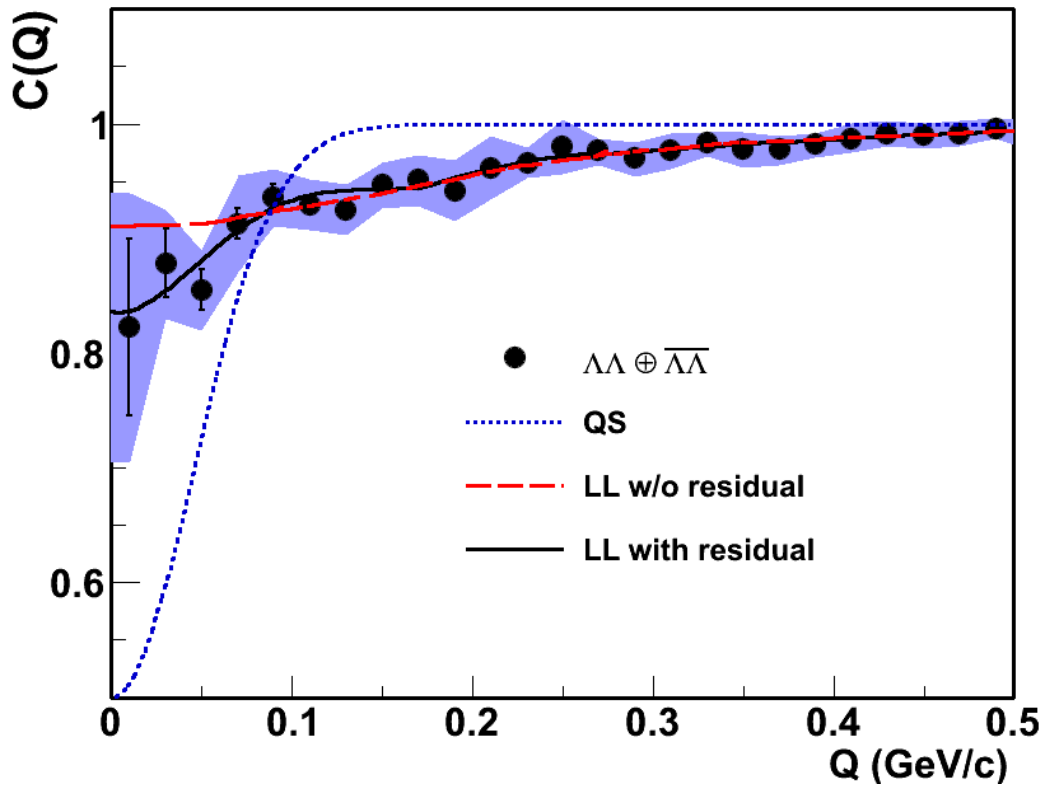


Constraints on a_0 and r_{eff}



STAR Coll.の解析

PRL114 '15



Interaction: Lednicky-Lyuboshitz model

Fit w/ 7 parameters

$a_0 > 0$: 引力

$$a_0 = -1.10 \pm 0.37^{+0.68}_{-0.08} \text{ fm}$$
$$r_{\text{eff}} = 8.52 \pm 2.56^{+2.09}_{-0.74} \text{ fm}$$
$$\chi^2/N_{\text{dof}} = 0.56$$

$$C_{\text{fit}}(Q) = N \left[1 + \lambda \left\{ -\frac{1}{2} e^{-r_0^2 Q^2} + \chi(a_0, r_{\text{eff}}) \right\} + a_{\text{res}} \exp(-r_{\text{res}}^2 Q^2) \right]$$

	HBT (size r_0)	Interaction	Residual correlation	
1.006	0.18	2.96 fm	-0.044	0.43 fm

まとめと展望

- 重イオン衝突(QGP生成)から $\Lambda\Lambda$ interaction
- 膨張の効果
- 崩壊から来る寄与の除去が重要
 - うまくできれば強い制限
 - できなければごく弱い制限
- $\Lambda\Lambda$ 散乱長 : $1/a_0 < -0.8 \text{ fm}^{-1}$
- 他のチャンネルにも応用可能
- 統計の問題 : LHCに期待
(Λp は綺麗なデータがあり)

Backup

TABLE I: $\Lambda\Lambda$ potentials. The scattering length (a_0) and effective range (r_{eff}) are fitted using a two-range gaussian potential, $V_{\Lambda\Lambda}(r) = V_1 \exp(-r^2/\mu_1^2) + V_2 \exp(-r^2/\mu_2^2)$.

Model	a_0 (fm)	r_{eff} (fm)	μ_1 (fm)	V_1 (MeV)	μ_2 (fm)	V_2 (MeV)	Ref.
ND46	4.621	1.300	1.0	-144.89	0.45	127.87	[18] $r_c = 0.46$ fm
ND48	14.394	1.633	1.0	-150.83	0.45	355.09	[18] $r_c = 0.48$ fm
ND50	-10.629	2.042	1.0	-151.54	0.45	587.21	[18] $r_c = 0.50$ fm
ND52	-3.483	2.592	1.0	-150.29	0.45	840.55	[18] $r_c = 0.52$ fm
ND54	-1.893	3.389	1.0	-147.65	0.45	1114.72	[18] $r_c = 0.54$ fm
ND56	-1.179	4.656	1.0	-144.26	0.45	1413.75	[18] $r_c = 0.56$ fm
ND58	-0.764	6.863	1.0	-137.74	0.45	1666.78	[18] $r_c = 0.58$ fm
NF42	3.659	0.975	0.6	-878.97	0.45	1048.58	[19] $r_c = 0.42$ fm
NF44	23.956	1.258	0.6	-1066.98	0.45	1646.65	[19] $r_c = 0.44$ fm
NF46	-3.960	1.721	0.6	-1327.26	0.45	2561.56	[19] $r_c = 0.46$ fm
NF48	-1.511	2.549	0.6	-1647.40	0.45	3888.96	[19] $r_c = 0.48$ fm
NF50	-0.772	4.271	0.6	-2007.35	0.45	5678.97	[19] $r_c = 0.50$ fm
NF52	-0.406	8.828	0.6	-2276.73	0.45	7415.56	[19] $r_c = 0.52$ fm
NSC89-1020	-0.250	7.200	1.0	-22.89	0.45	67.45	[20] $m_{\text{cut}} = 1020$ MeV
NSC89-920	-2.100	1.900	0.6	-1080.35	0.45	2039.54	[20] $m_{\text{cut}} = 920$ MeV
NSC89-820	-1.110	3.200	0.6	-1904.41	0.45	4996.93	[20] $m_{\text{cut}} = 820$ MeV
NSC97a	-0.329	12.370	1.0	-69.45	0.45	653.86	[21]
NSC97b	-0.397	10.360	1.0	-78.42	0.45	741.76	[21]
NSC97c	-0.476	9.130	1.0	-91.80	0.45	914.67	[21]
NSC97d	-0.401	1.150	0.4	-445.77	0.30	373.64	[21]
NSC97e	-0.501	9.840	1.0	-110.45	0.45	1309.55	[21]
NSC97f	-0.350	16.330	1.0	-106.53	0.45	1469.33	[21]
Ehime	-4.21	2.41	1.0	-146.6	0.45	720.9	[23]
fss2	-0.81	3.99	0.92	-103.9	0.41	658.2	[25]
ESC08	-0.97	3.86	0.80	-293.66	0.45	1429.27	[22]

TABLE II: $\Lambda\Lambda$ potentials from Nagara event. The scattering length (a_0) and effective range (r_{eff}) are fitted using a three-range gaussian potential, $V_{\Lambda\Lambda}(r) = V_1 \exp(-r^2/\mu_1^2) + V_2 \exp(-r^2/\mu_2^2) + V_3 \exp(-r^2/\mu_3^2)$.

Model	a_0 (fm)	r_{eff} (fm)	μ_1 (fm)	V_1 (MeV)	μ_2 (fm)	V_2 (MeV)	μ_3 (fm)	V_3 (MeV)	Ref.
HKMY	-0.575	6.45	1.342	-10.96	0.777	-141.75	0.35	2136.6	[3]
FG	-0.77	6.59	1.342	-21.49	0.777	-250.13	0.35	9324.0	[2]

$$N = 1.006 \pm 0.001$$

$$\lambda = 0.18 \pm 0.05^{+0.12}_{-0.06}$$

$$a_0 = -1.10 \pm 0.37^{+0.68}_{-0.08}$$

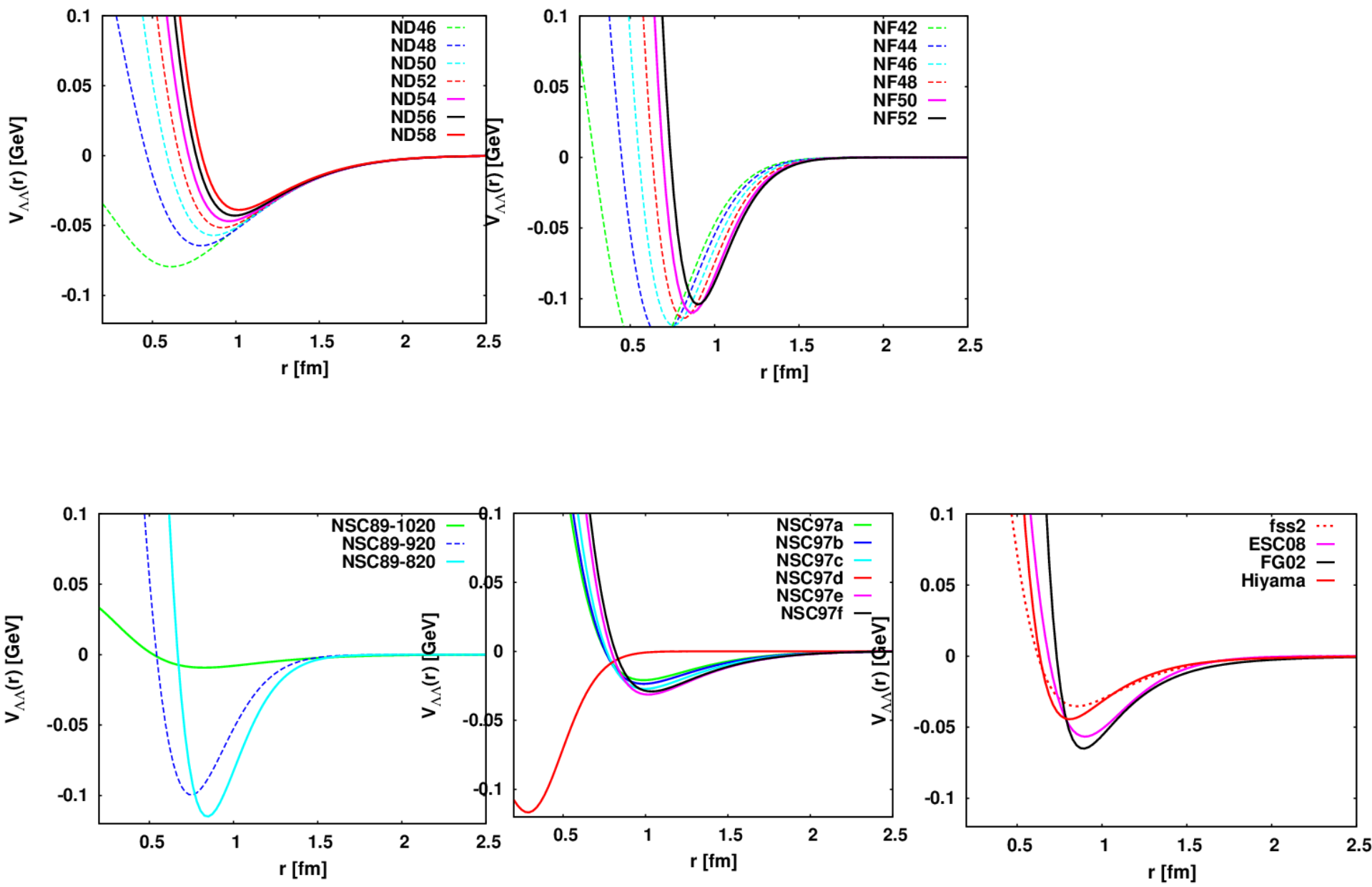
$$r_{\text{eff}} = 8.52 \pm 2.56^{+2.09}_{-0.74}$$

$$r_0 = 2.96 \pm 0.38^{+0.96}_{-0.02}$$

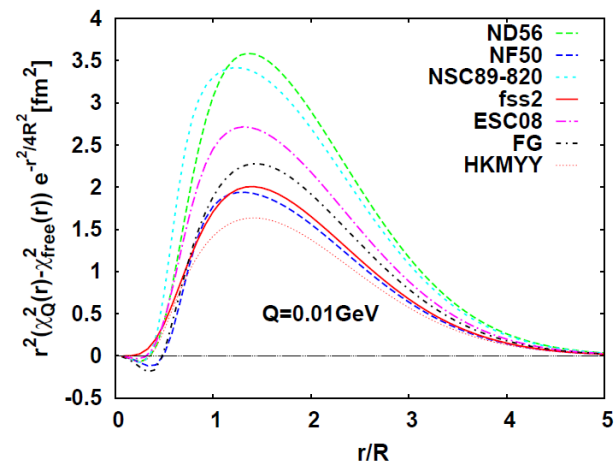
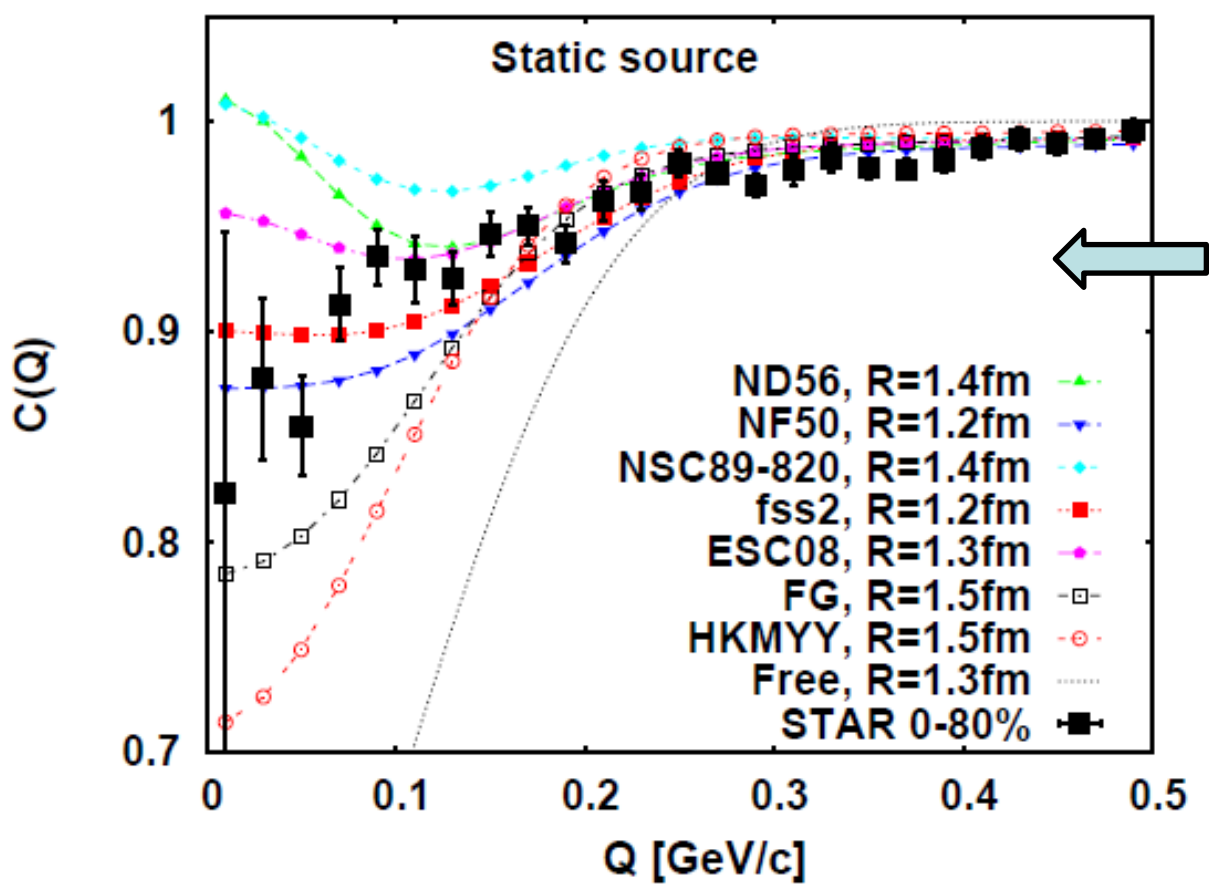
$$a_{\text{res}} = -0.044 \pm 0.004^{+0.048}_{-0.009}$$

$$r_{\text{res}} = 0.43 \pm 0.04^{+0.43}_{-0.03}$$

$$\chi^2/N_{\text{dof}} = 0.56$$



Results from the Static Source



Strong attraction
is reflected onto
 $C(Q)$

Larger variation among potentials than data error-bars

Size : determined from $\min. \chi^2$

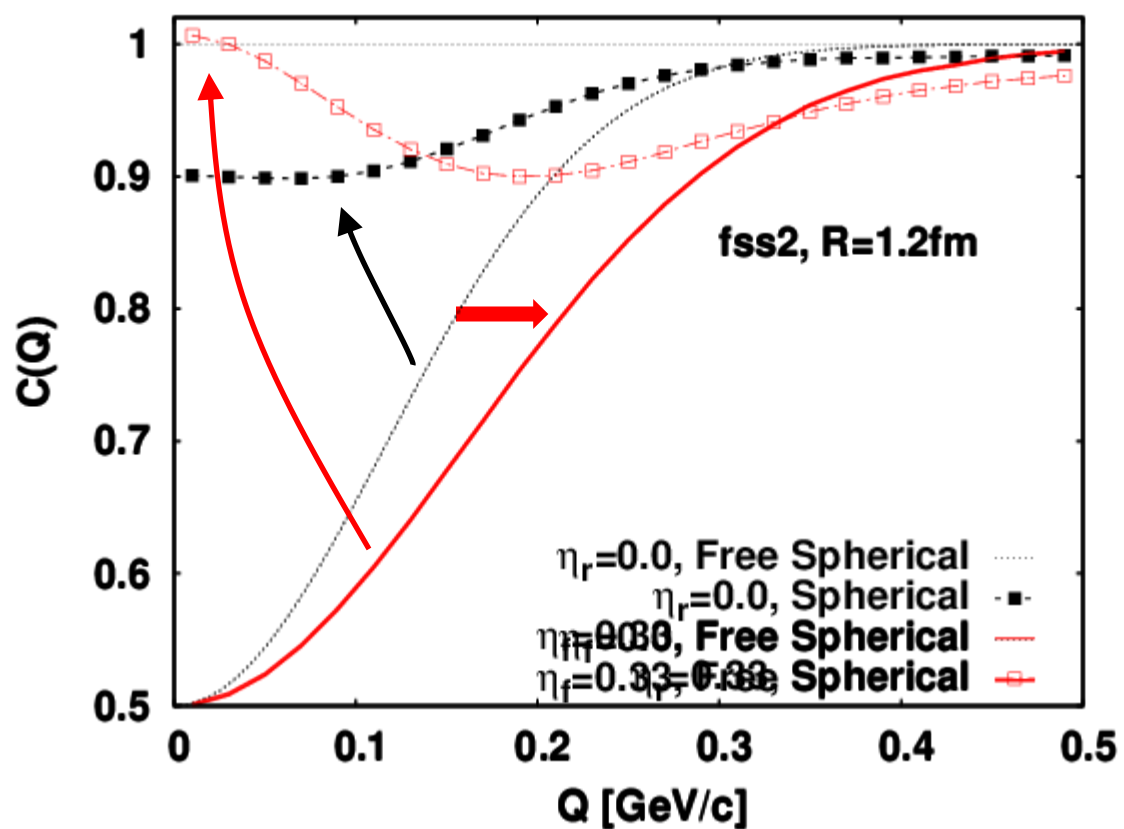
Collectivity Deforms Source Function

$$C_2(Q, K) = \frac{\int d^4x_1 d^4x_2 S(x_1, K) S(x_2, K) |\Psi_{12}(Q, x_1 - x_2 - (t_2 - t_1)K/m)|^2}{\int d^4x_1 d^4x_2 S(x_1, k_1) S(x_1, k_2)}$$

Influence on the best-fit potentials?

C(Q) is fairly sensitive
to interaction

Illustration : Spherical Source



Expansion makes the apparent source size smaller

Attraction in $C(Q)$ becomes stronger