

# 中性子星内における $\Sigma^-$ 出現 の可能性

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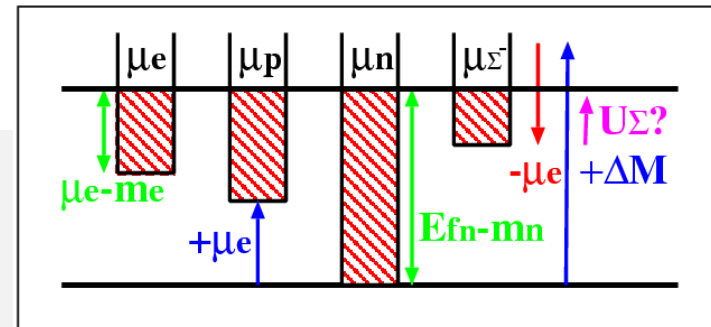
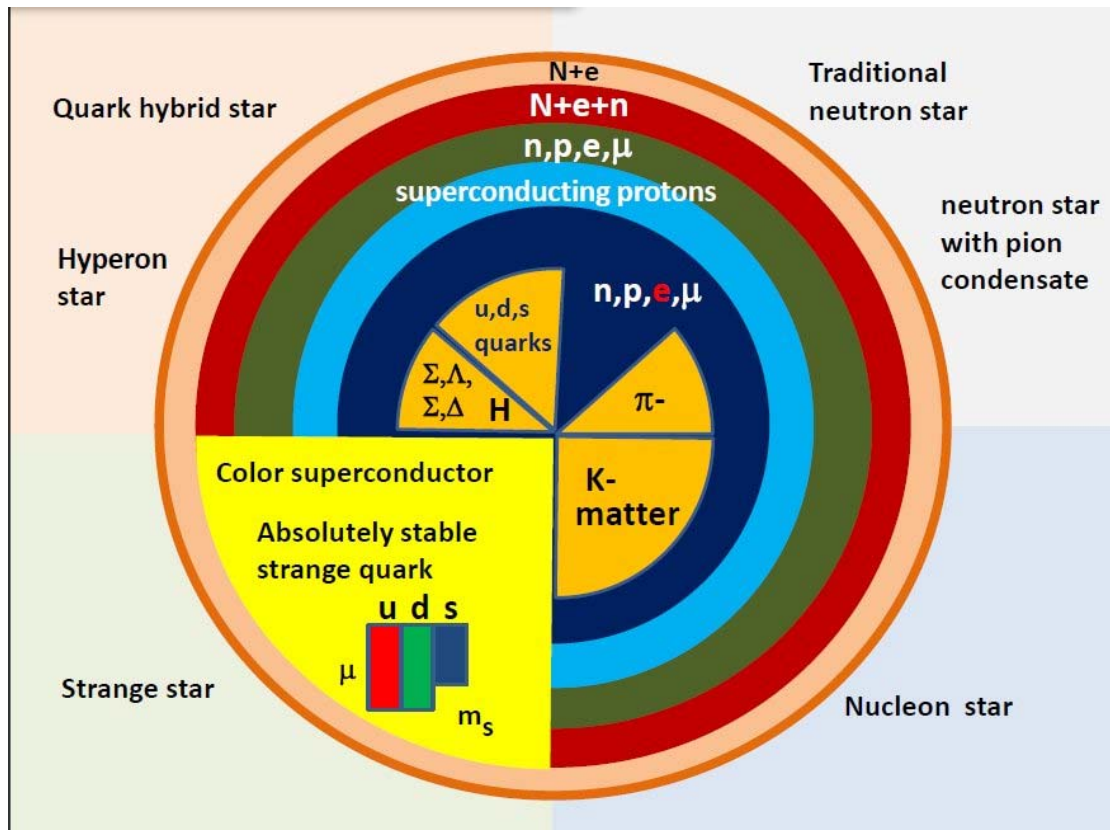
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# Introduction

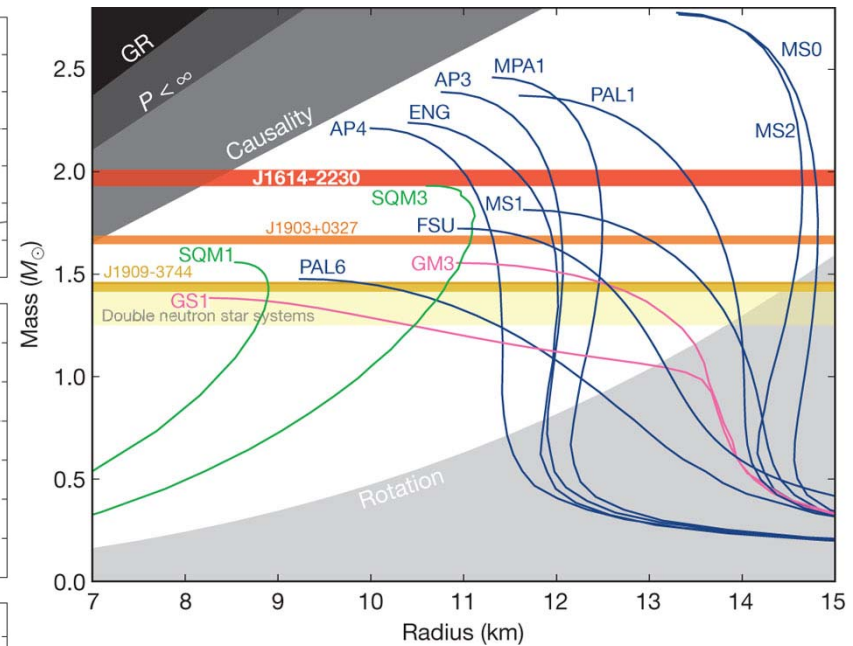
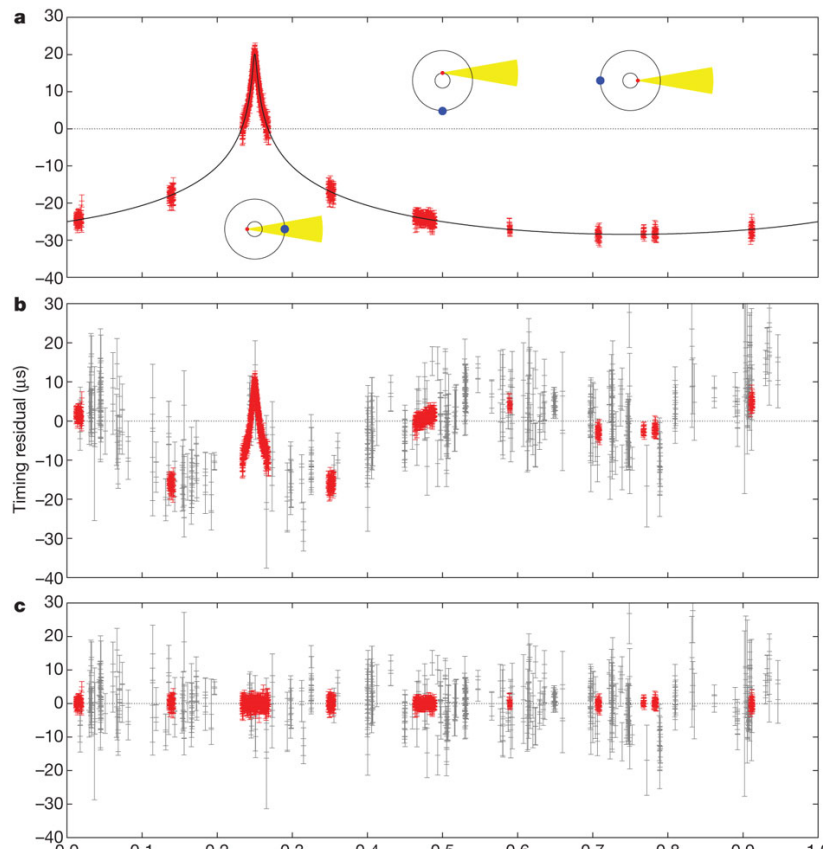
► Ambiguity in dense inner region of NS



- New DOF such as Exotic particle soften NS EOS
- Emergence of hyperon: needs to treat “finite” hypernuclear systems in order to determine hyperon potential
- Framework which enable to treat finite/infinite system →RMF model

# Introduction

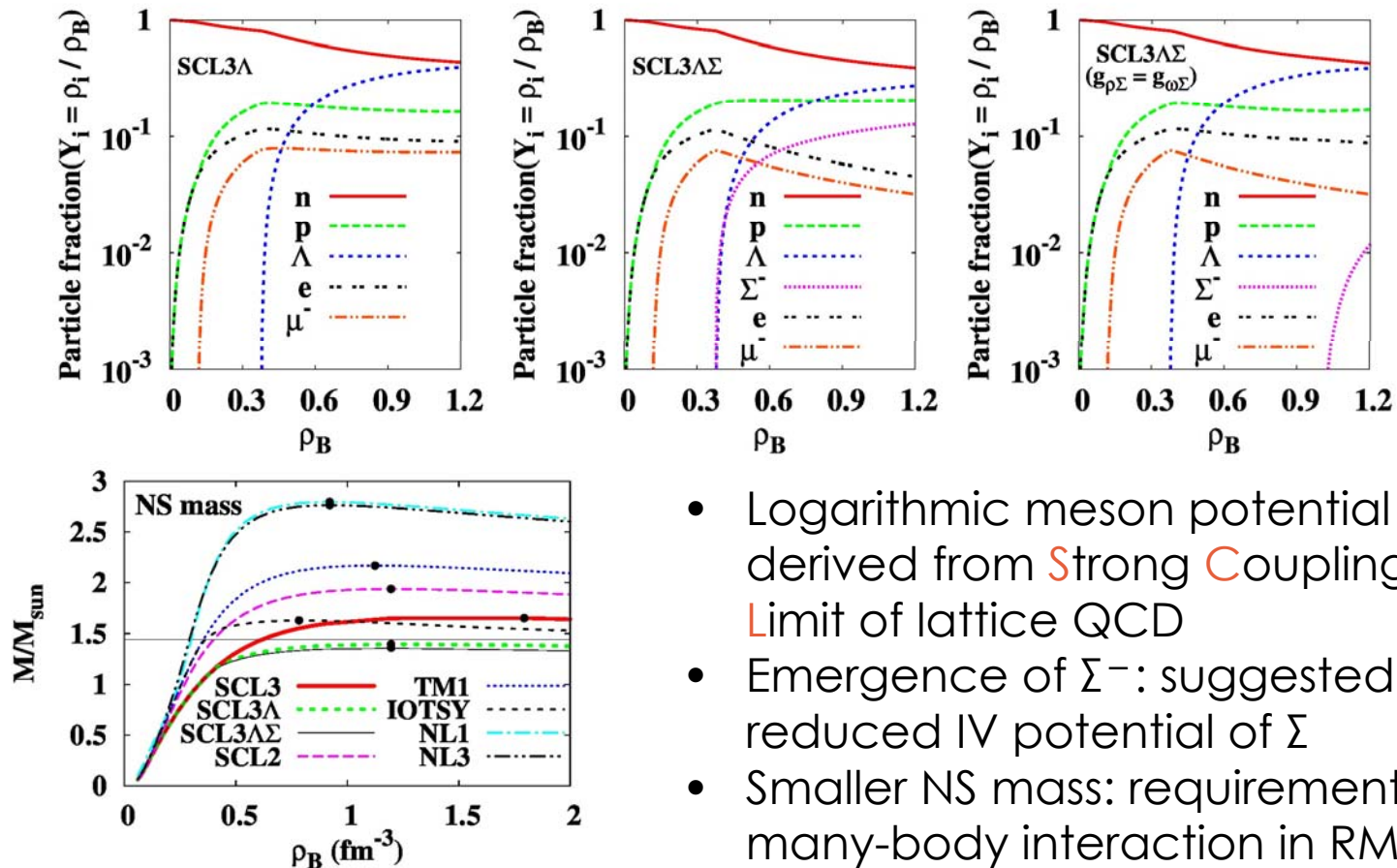
## ► Observation of massive neutron star



“Most EOS curves involving exotic matter, such as kaon condensates or hyperons tend to predict maximum masses well below  $2.0 M_{\odot}$ .”

# Personal motivation

- ▶ Previous works on **SCL3** RMF model

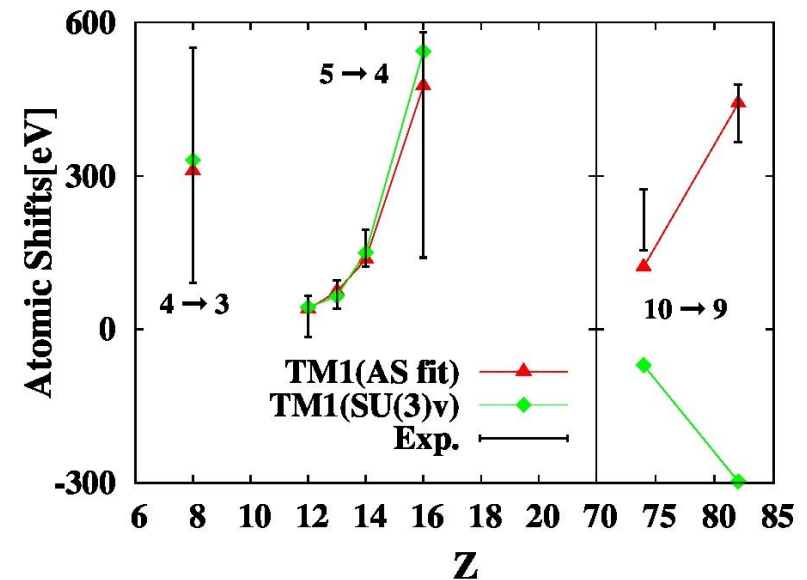
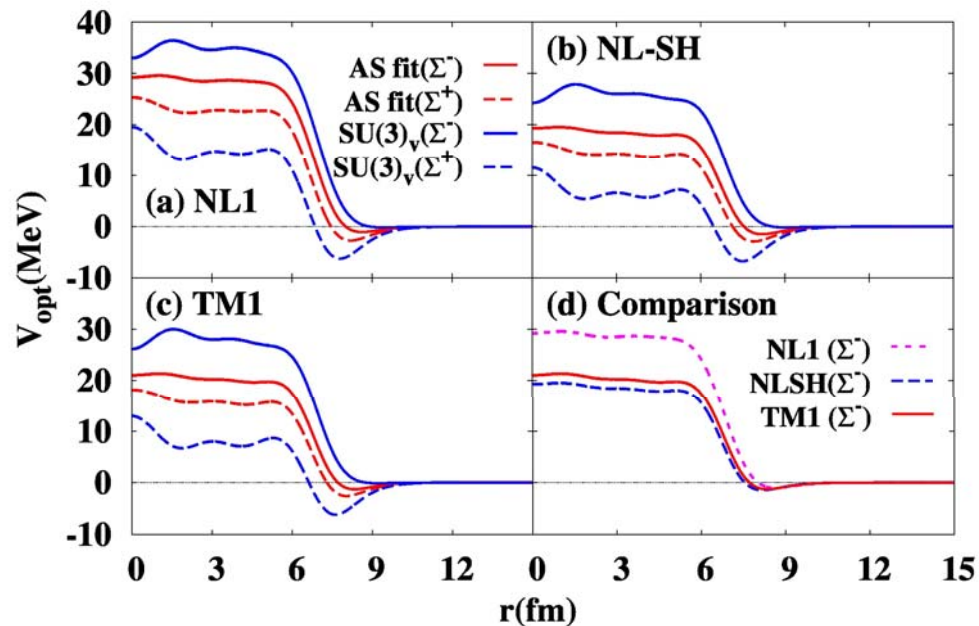


- Logarithmic meson potential derived from **S**trong **C**oupling **L**imit of lattice QCD
- Emergence of  $\Sigma^-$ : suggested from reduced IV potential of  $\Sigma$
- Smaller NS mass: requirement of many-body interaction in RMF?

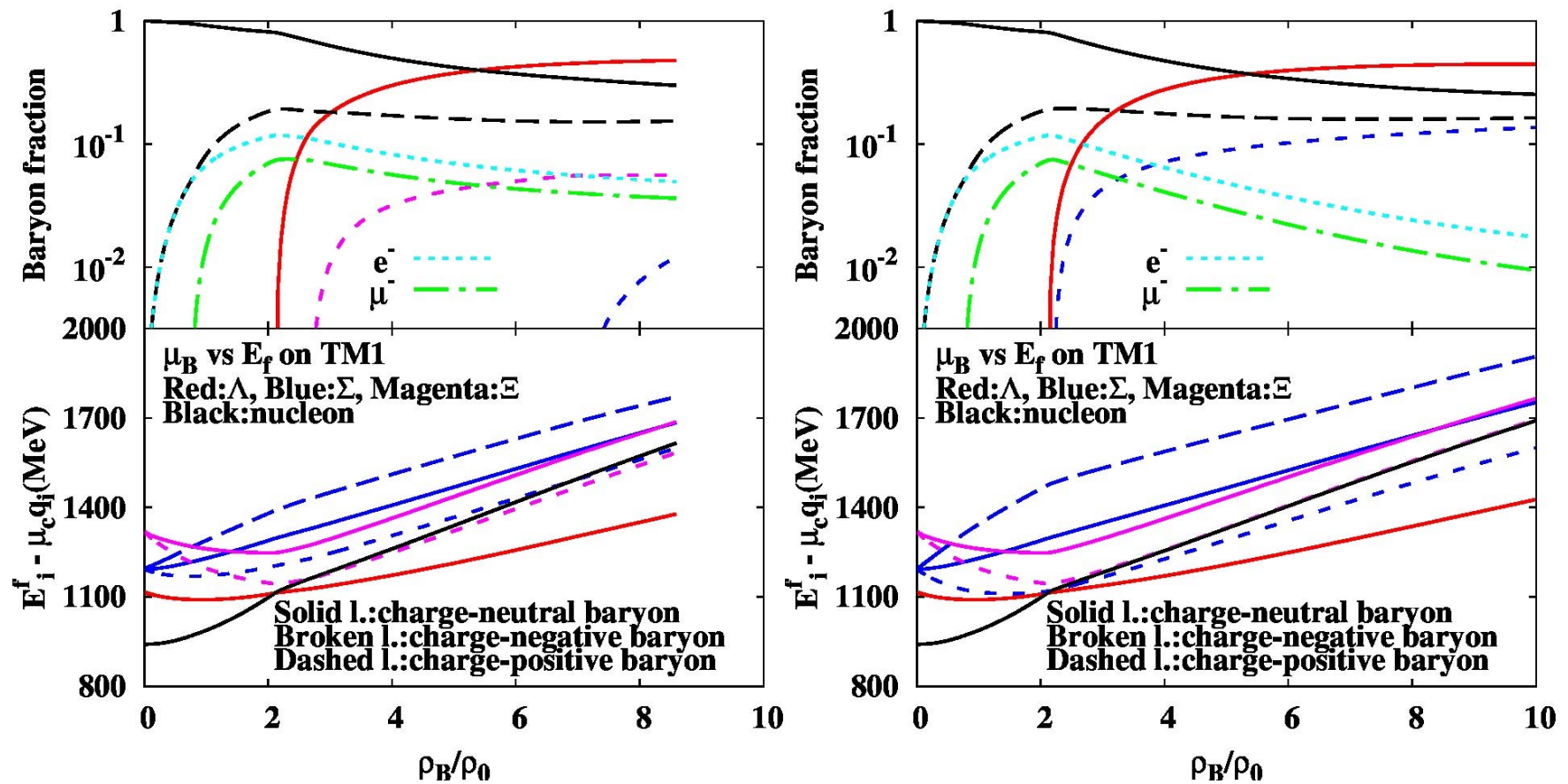
# Hypernuclear Symmetry Energy

- ▶ Atomic shift of  $\Sigma^-$ : Attractive potential around surface
- ▶ Difference between  $\Sigma^+$  and  $\Sigma^-$  potentials  $\rightarrow$  proportional to "Hypernuclear" symmetry energy: 15MeV(SU(3))  $\rightarrow$  5MeV(fit)

Optical  $\Sigma$  potentials in Pb



# Baryon Fraction



$\Sigma^-$  may survive in NS matter even if  $\Xi^-$  is taken into account ( $U_{\Xi}(^{12}\text{C}-\Xi^-) = -14\text{MeV}$ ).

# Physical motivation

- ▶ More massive NS than  $2M_{\odot}$ :  
How can be explained in RMF model?  
→  $n=3$  meson-baryon couplings
- ▶ Can  $\Sigma^-$  survive in NS EOS in case we consider above  
 $n=3$  couplings and how large effects does  $\Sigma^-$  take  
from those couplings

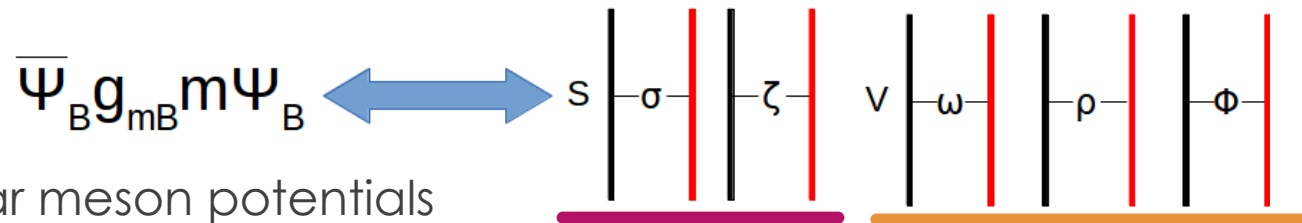
# RMF model

- ▶  $n=B/2+M+D=2$  RMF Lagrangian + meson potentials

$$\mathcal{L} = \sum_i \bar{\psi} [i\partial - \underline{M}_i^* - \gamma_\mu \underline{V}^\mu] \psi - \frac{1}{4} F_{\mu\nu} F^{\mu\nu}$$

$$+ \frac{1}{2} \partial_\mu \varphi_\sigma \partial^\mu \varphi_\sigma - \frac{m_\sigma^2}{2} \varphi_\sigma^2 + \frac{1}{2} \partial_\mu \varphi_\zeta \partial^\mu \varphi_\zeta - \frac{m_\zeta^2}{2} \varphi_\zeta^2 - V_{\sigma\zeta}$$

$$- \frac{1}{4} W_{\mu\nu} W^{\mu\nu} + \frac{m_\omega^2}{2} \omega_\mu \omega^\mu - \frac{1}{4} R_{\mu\nu} R^{\mu\nu} + \frac{m_\rho^2}{2} \rho_\mu \rho^\mu - \frac{1}{4} \phi_{\mu\nu} \phi^{\mu\nu} + \frac{m_\phi^2}{2} \phi_\mu \phi^\mu + \frac{C_{\omega^4}}{4} (\omega_\nu \omega^\nu)^2$$



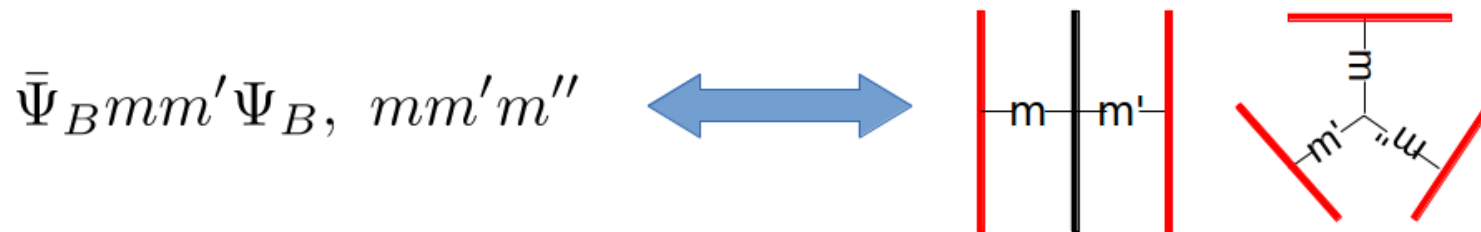
- ▶ Non-linear meson potentials

- ▶ Interactions among mesons in finite/infinite system
- ▶ Higher order terms than  $n=3$  meson-baryon couplings



# n=3 coupling in RMF

- ▶  $n=B/2+m+D=3$  couplings



- ▶ Number  $n$  corresponds to the number of baryons participating in each interaction

- ▶ Isoscalar type couplings

$$\bar{\Psi}_B \left( \frac{g_{\sigma\sigma B} \varphi_\sigma^2}{f_\pi} \right) \Psi_B, \bar{\Psi}_B \left( \frac{g_{\sigma\omega B} \varphi_\sigma \omega}{f_\pi} \right) \Psi_B, \bar{\Psi}_B \left( \frac{g_{\omega\omega B} \omega^2}{f_\pi} \right) \Psi_B, \frac{C_{\sigma\omega^2}}{2} f_\pi \varphi_\sigma \omega^2$$

- ▶ Isovector type couplings

$$\bar{\Psi}_B \left( \frac{g_{\rho\sigma B} \rho \varphi_\sigma}{f_\pi} \right) \Psi_B, \bar{\Psi}_B \left( \frac{g_{\omega\rho B} \omega \rho}{f_\pi} \right) \Psi_B, \bar{\Psi}_B \left( \frac{g_{\rho\rho B} \rho^2}{f_\pi} \right) \Psi_B, \frac{C_{\sigma\rho^2}}{2} f_\pi \varphi_\sigma \rho^2$$

それぞれ scalar、vector 型結合

# Saturation parameters in RMF

▶ L,Kの具体的な表式

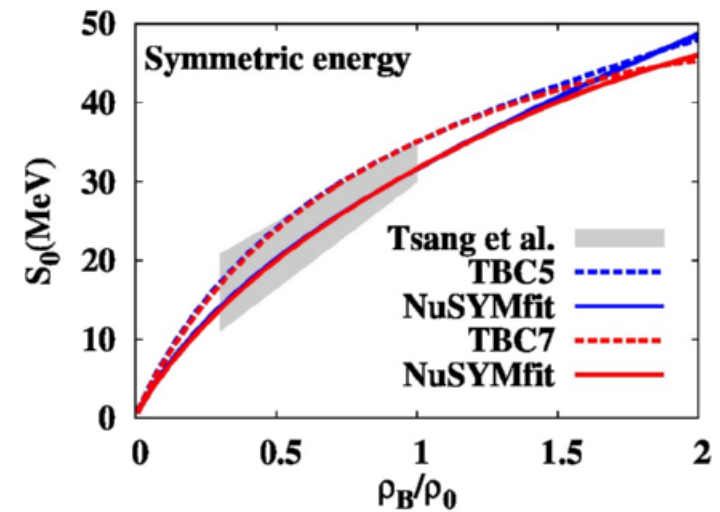
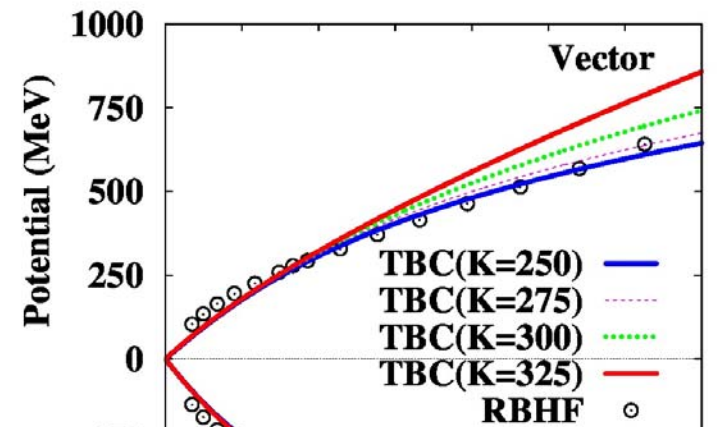
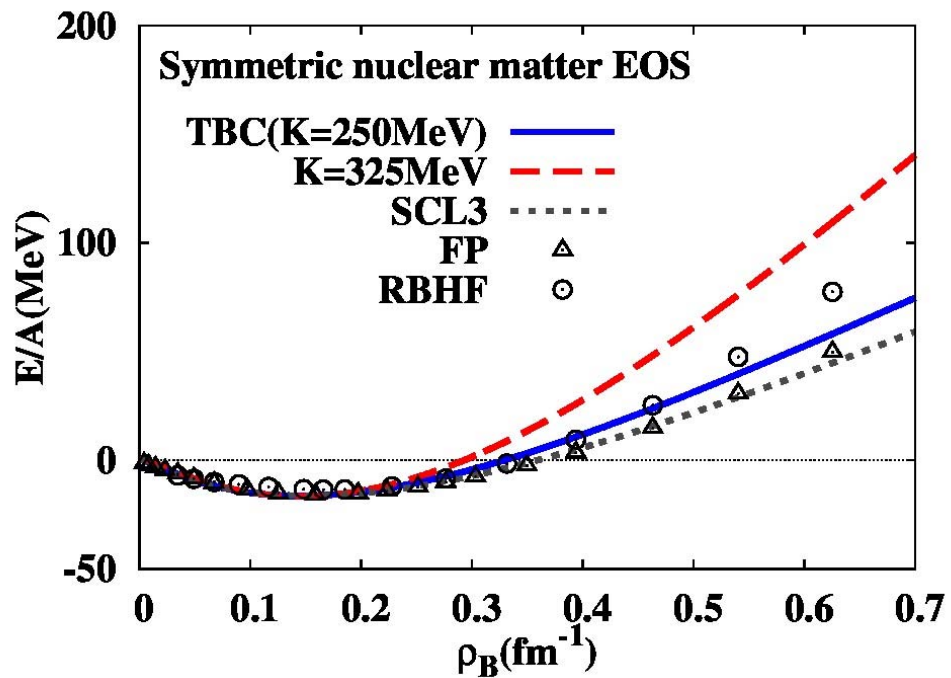
$$L_{\infty} \frac{dE}{d\rho_B} = \sum_i \frac{d\rho_{Bi}}{d\rho_B} \frac{dE}{d\rho_{Bi}} = \frac{1}{\rho_B} \left[ \frac{1}{2} \sum_i \left\{ V_i^0 + \sqrt{k_{Fi}^2 + (M_i^*)^2} \right\} - E \right]$$

$$K_{\infty} = \frac{9\rho_0}{4} \sum_j \left\{ \frac{\pi^2}{k_{Fj}} \frac{1}{\sqrt{k_{Fj}^2 + (M_j^*)^2}} + \sum_i \sum_m \frac{\partial f_m}{\partial \rho_{Bj}} \frac{\partial}{\partial f_m} \left( V_i^0 + \sqrt{k_{Fi}^2 + (M_i^*)^2} \right) \right\}$$

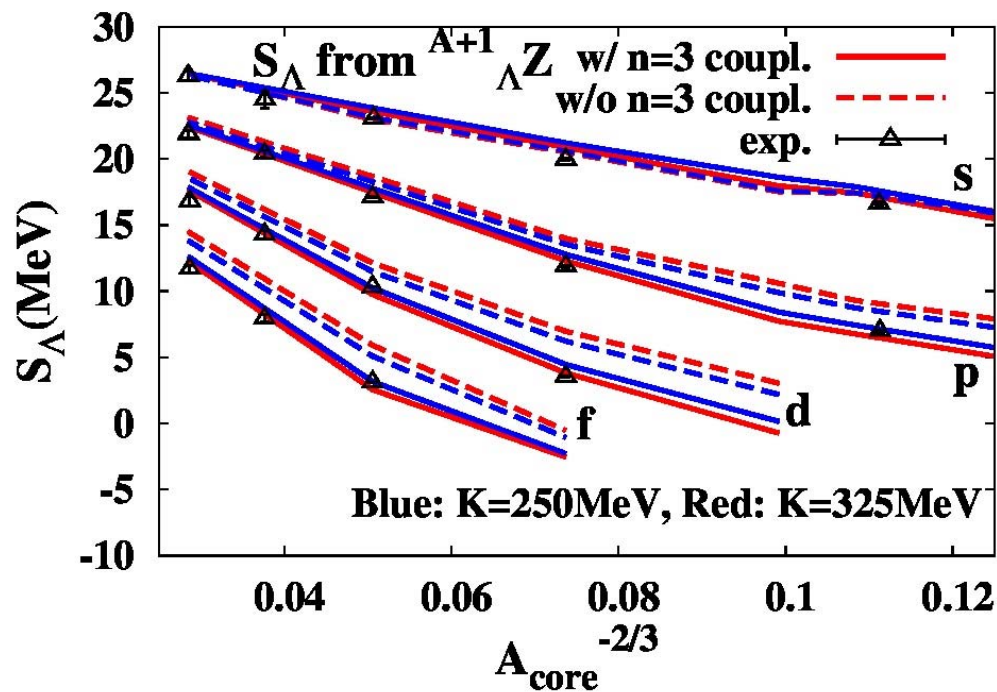
$$= \frac{9\rho_0}{4} \sum_j \left[ \frac{\pi^2}{k_{Fj}} \frac{1}{\sqrt{k_{Fj}^2 + (M_j^*)^2}} + \frac{d\omega}{d\rho_{Bj}} \sum_i \frac{1}{f_{\pi}} \left( g_{\omega i} f_{\pi} - g_{\sigma \omega i} \varphi_{\sigma} + \frac{2g_{\omega \omega i} \omega M_i^*}{\sqrt{k_{Fi}^2 + (M_i^*)^2}} \right) \right. \\ \left. + \frac{d\varphi_{\sigma}}{d\rho_{Bj}} \sum_i \frac{1}{f_{\pi}} \left\{ -g_{\sigma \omega i} \omega + \frac{(-g_{\sigma i} f_{\pi} + 2g_{\sigma \sigma i} \varphi_{\sigma}) M_i^*}{\sqrt{k_{Fi}^2 + (M_i^*)^2}} \right\} \right]$$

- L,K can be constraints to n=3 coupling parameters
- Coupling strengths: Decidable based on also B/E and charge rms radii

# Parameter setting

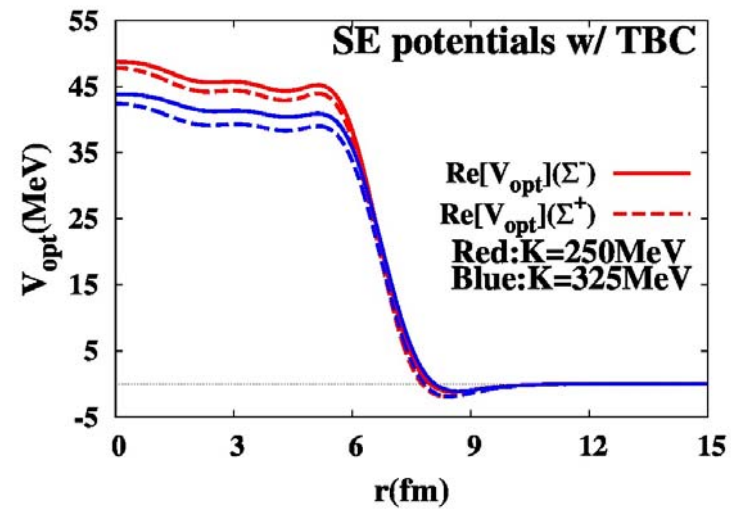
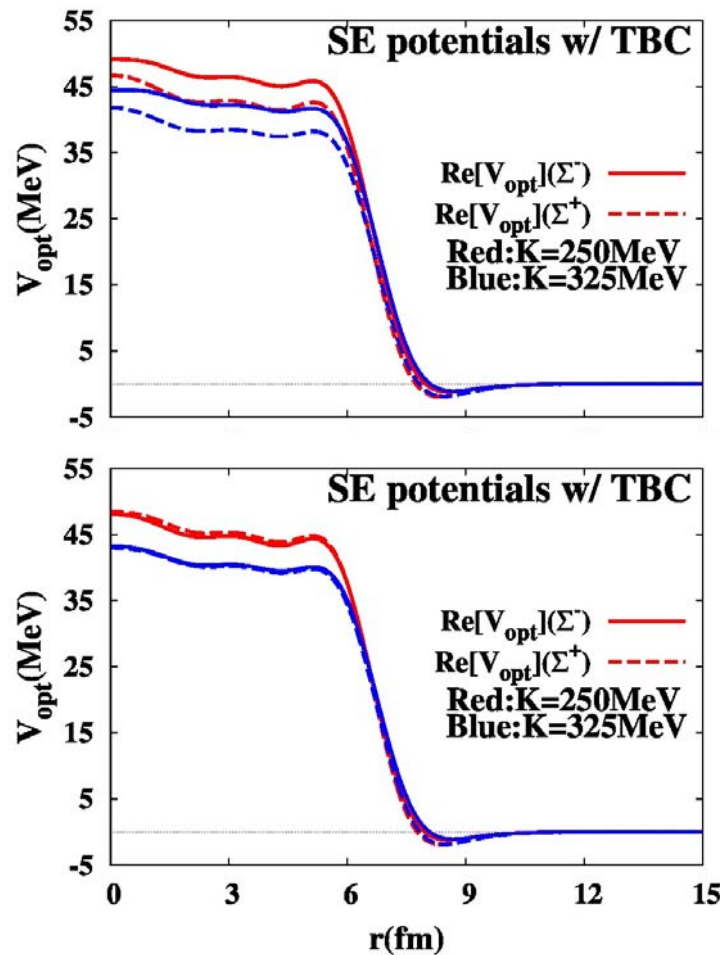


# Results: $S_\Lambda$ from single $\Lambda$ hypernuclei



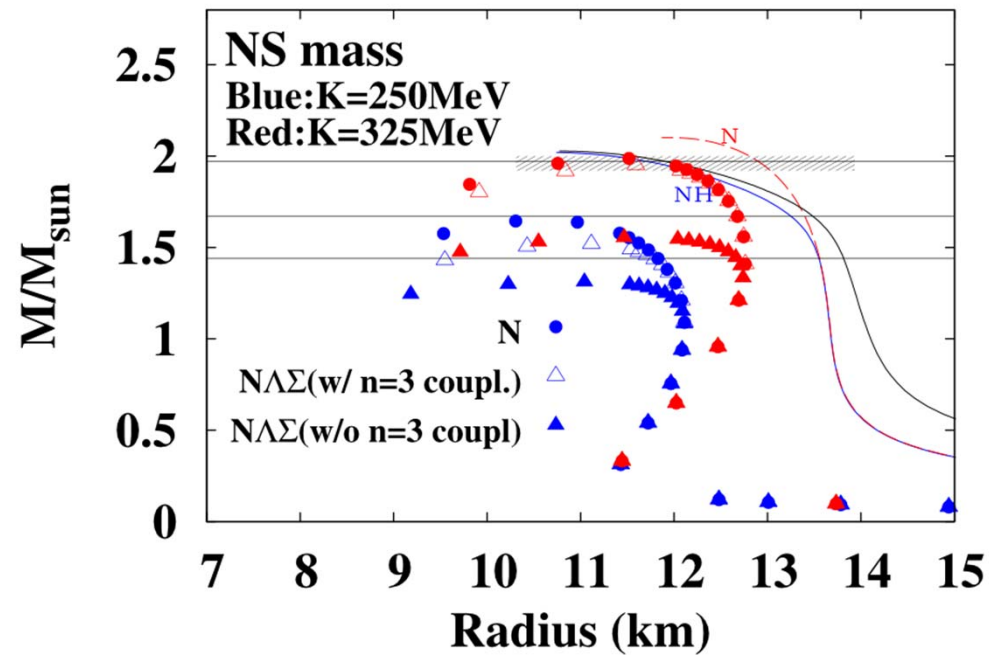
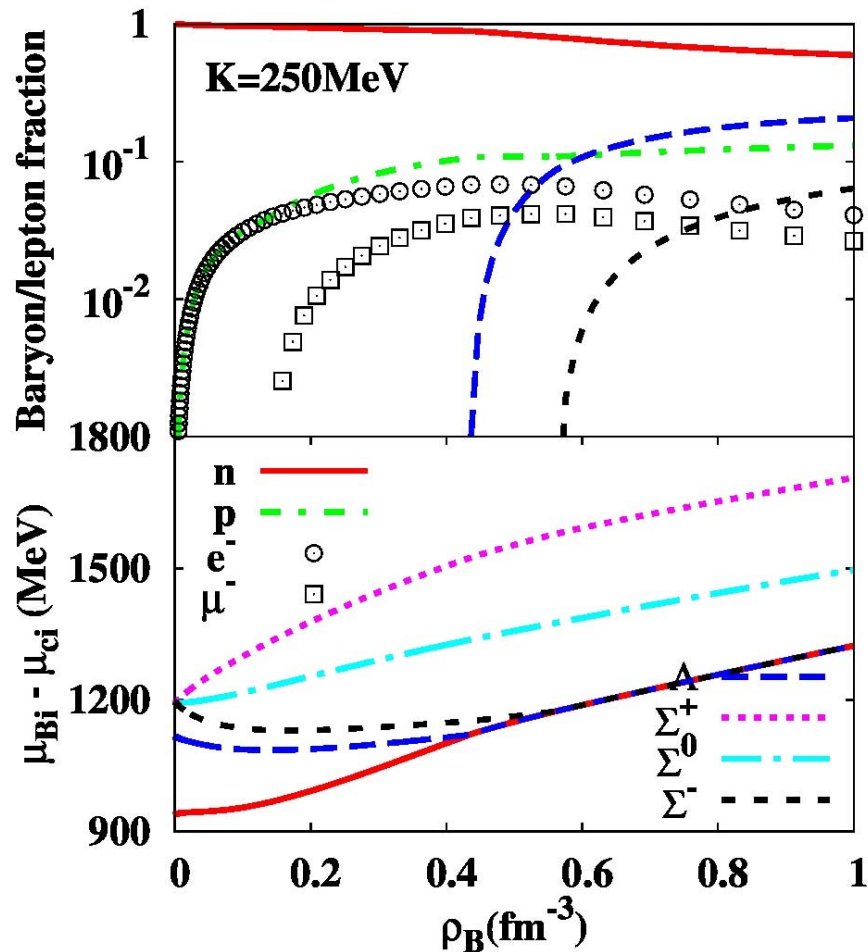
- Role of incompressibility to  $S_\Lambda \Rightarrow$  Almost free if we decide n=3 couplings where  $\Lambda$  participate  
 $\Leftrightarrow$  We can determine the strengths of n=3 couplings of  $\Lambda$  enough to each parameter set characterized by the incompressibility.

# Results: $\Sigma^-$ optical potential



Smaller hypernuclear  
symmetry energy:  $5 \sim 0\text{MeV}$   
suggested from factorized  
 $n=3$  coupling of  $\Sigma^-$

# Baryon Fraction plot with $\mu$



- Emergence of hyperons is pushed to higher  $\rho_B$  region.
- Calculated NS mass can exceed  $2M_{\odot}$  although hyperons are taken into account. (K=325 MeV)

# Summary

- ▶  $\Sigma^-$  may allow to emerge in NS matter EOS if we respect the experimental results of  $\Sigma^-$  atomic shifts.
- ▶ Emergence of  $\Xi^-$  may suppressed if we apply SU(3)<sub>v</sub> symmetry to coupling constants.
- ▶ n=3 couplings in RMF model:  
well characterized by incompressibility and constrained by reproducing experimental/  
values of nuclear/hypernuclear systems.
- ▶ Calculated NS mass excess  $2M_{\odot}$  (K=325MeV).
- ▶ Even if we introduced n=3 coupling to  $\Sigma^-$ ,  $\Sigma^-$  may appear in NS EOS as substitute of negative particles.



Thank you  
for your listening.

THAT'S ALL.



# Introduction

## Baryon and lepton fractions in NS matter

