

中性子過剰核物質中のストレンジネス(計 画研究AO2)

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鈴木隆敏(東京大学)

Physics of Neutron Stars and A02 issue

- Neutron stars
 - Various types of **signals** and **activities**
 - Identify neutron stars, information by **observation** (C01, D01)
 - **Structure**: “**Nuclear Matter**” from low to high ($\approx 9\rho_0$) densities
 - Low density: Atmosphere, Outer-crust, Inner-crust
 - High density: Outer-core, **Inner-core**
- What governs the structure of neutron stars?
 - **Gravity**: well known (general relativity), mass of neutron stars
 - **EoS of neutron matter** ($n + p + e^- + \mu^-$)
 - Density dependence of **symmetry energy** of nucleons
 - Experimental study with accelerators, etc. (B01-B03, D01)
 - Does the inner-core have **Exotic components**?
 - Contributions of **hyperons** ($\Lambda, \Sigma, \Xi, \dots$), **mesons** (π, K, \dots), etc.
 - Experimental studies at **J-PARC** (A01-02, D01)

Strangeness in Neutron Stars

- Strangeness appears in high density nuclear matter

Degree of nucleons and hyperons

$$\begin{aligned} n + p + e^- + \mu^- \\ \Lambda + \Sigma + \Xi \end{aligned}$$

Phenomenological BB interaction

$$V_y(\rho_x) = \frac{a_{xy}\rho_x}{\text{2-body}} + \frac{b_{xy}t_x t_y \rho_x}{\text{symmetry energy}} + \frac{c_{xy}\rho_x^{\gamma_{xy}}}{\text{3-body}}$$

γ_{xy} : Hardness of EoS

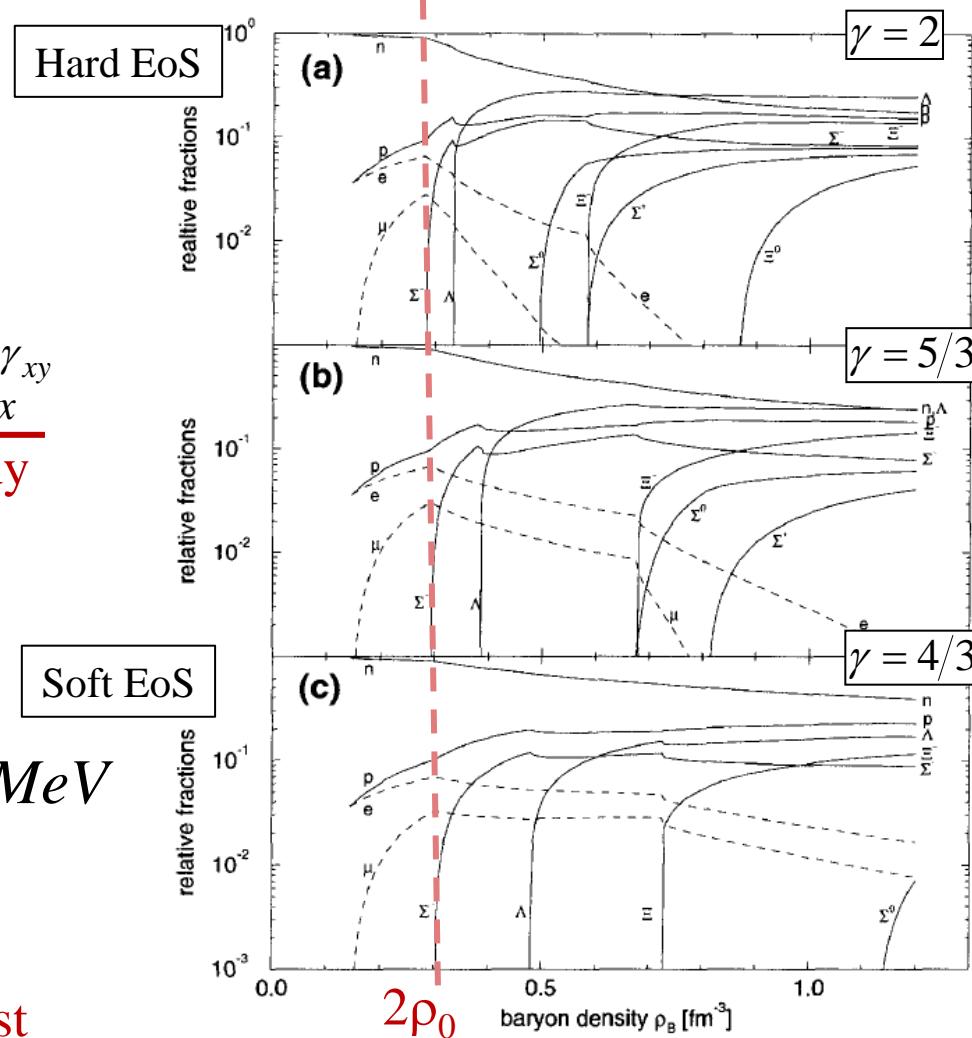
Hyperon-nucleon interaction

$$V_\Lambda(\rho_0) \approx V_\Sigma(\rho_0) \approx V_\Xi(\rho_0) \approx -30 \text{ MeV}$$

?

Hyperons appear around $2\rho_0$

Hyperon appearance looks robust

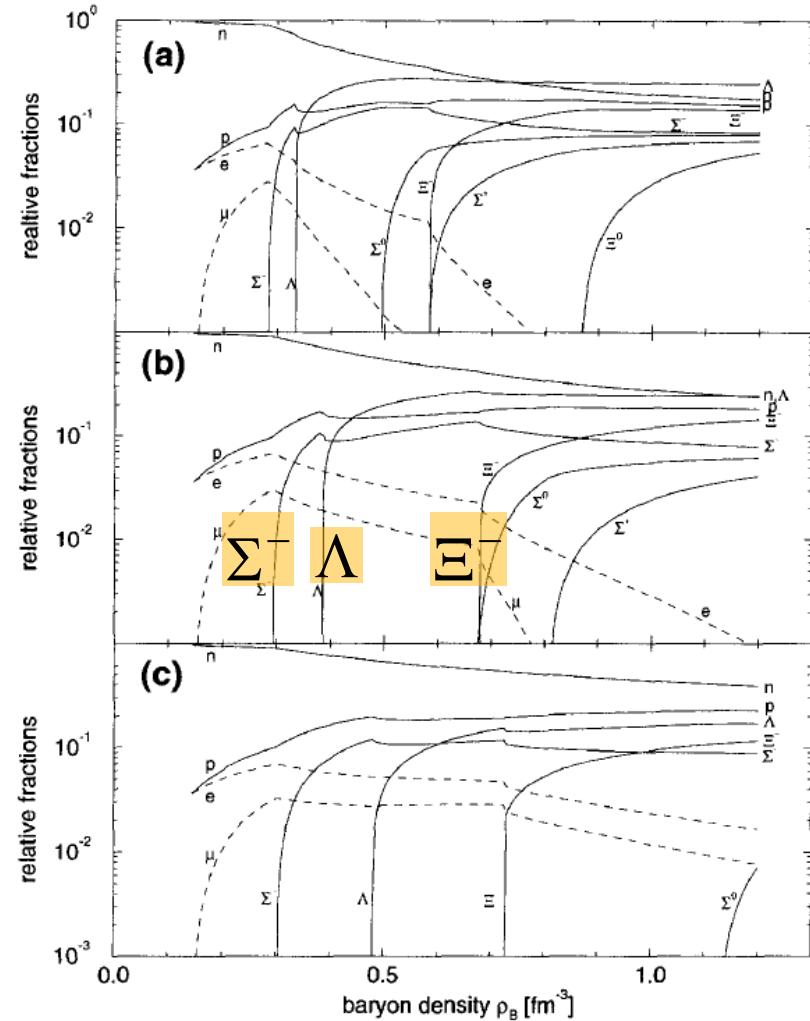


well fixed

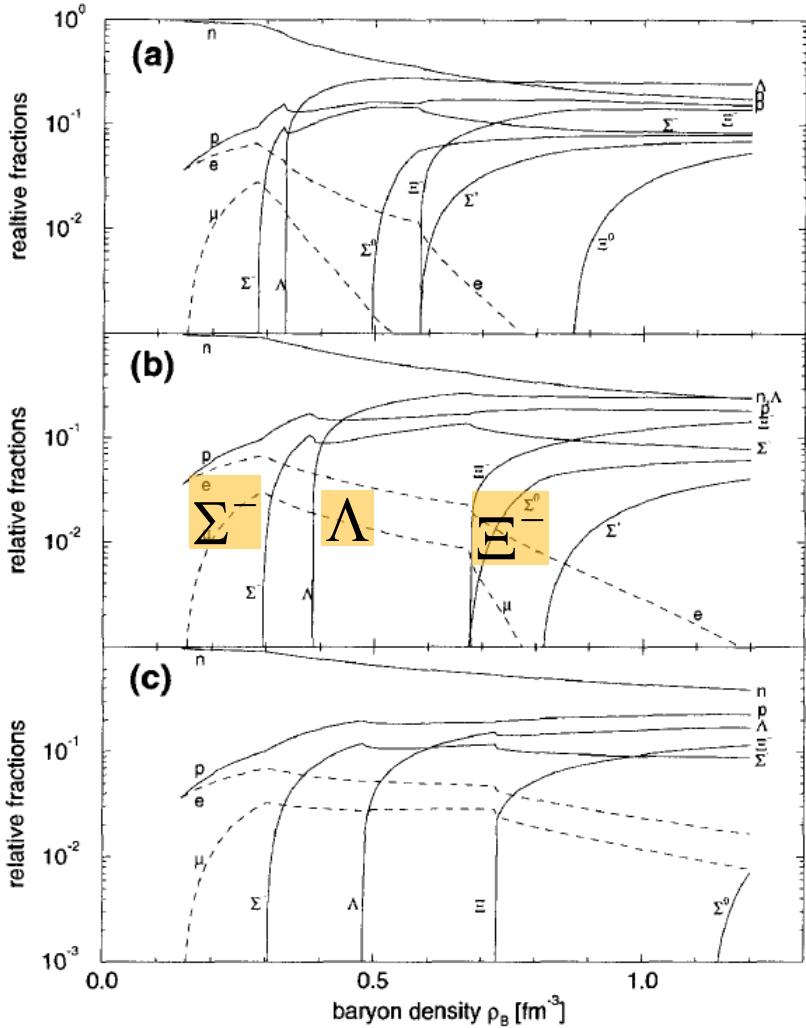
large ambiguity

- What controls the appearance of hyperons?

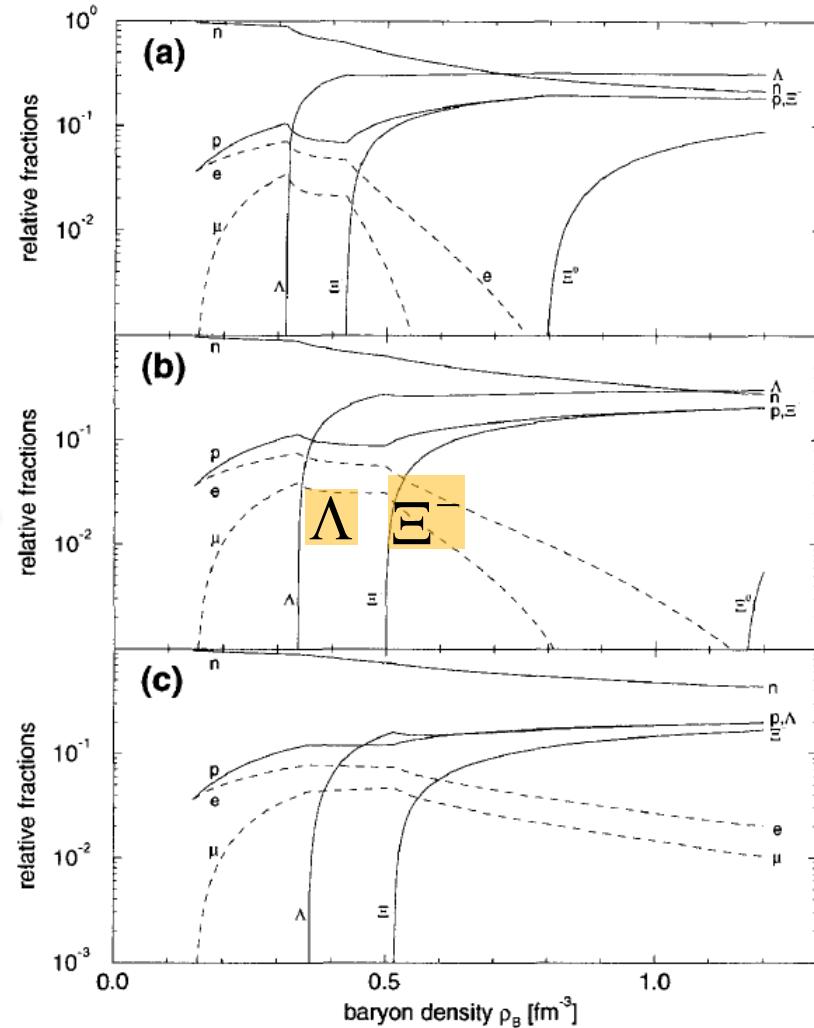
- Negative charge
 - Σ^- appears first
 - Appearance of Ξ^-
 - π^- and K^- condensation
- Mass of hyperons
 - Early appearance of Λ
- Attractive interaction
 - $V_{\Lambda n}$ is attractive
 - $\Sigma^- n$, $\Xi^- n$, $\Lambda \Lambda$ under study
 - π^- and K^- under study
- Symmetry energy
 - Large? for Σn ($I=1$)
 - Pauli blocking in quarks
 - None? for Λn ($I=0$)



- Example of effect of YN interaction

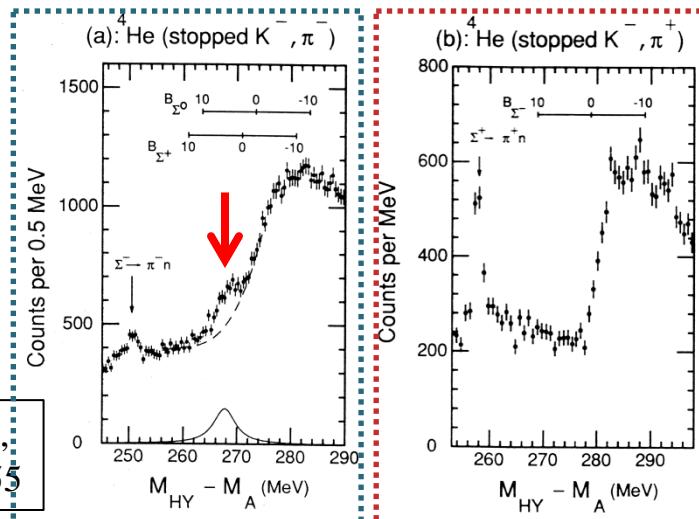


repulsive
ΣN int.

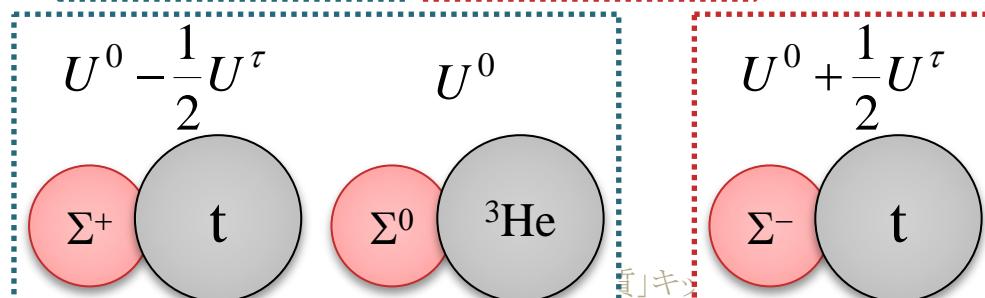
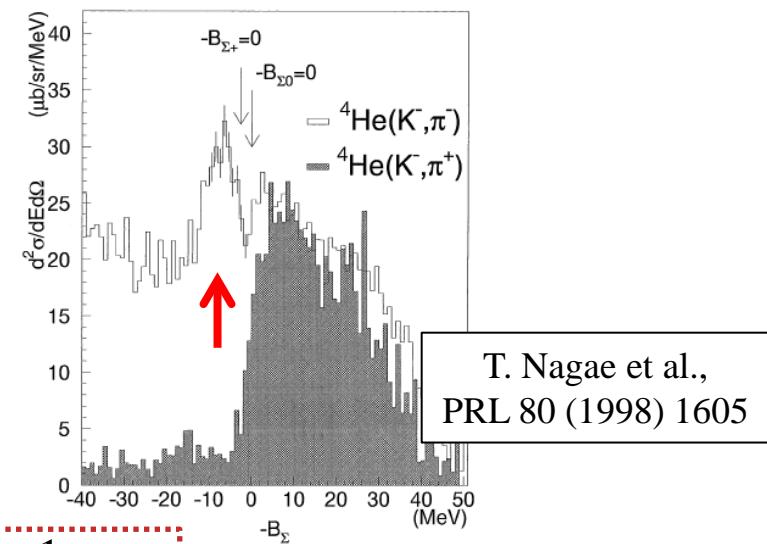


ΣN interaction (J-PARC E40)

- Sources of information on ΣN interaction
 - Σ hypernuclei
 - Clear bound state was observed in ${}^4\text{He}(\text{K}^-, \pi^-)$ reactions
 - Large decay width: strong ΣN - ΛN conversion in nuclei
 - No clear Σ bound state was observed in other hypernuclei

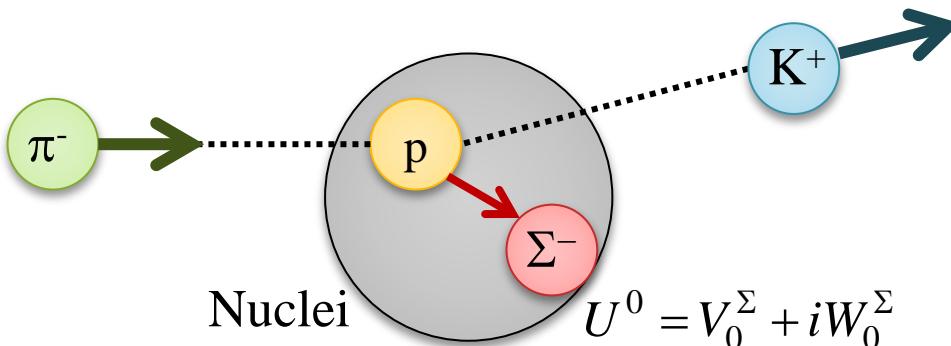


R.S. Hayano et al.,
 PLB 231 (1989) 355



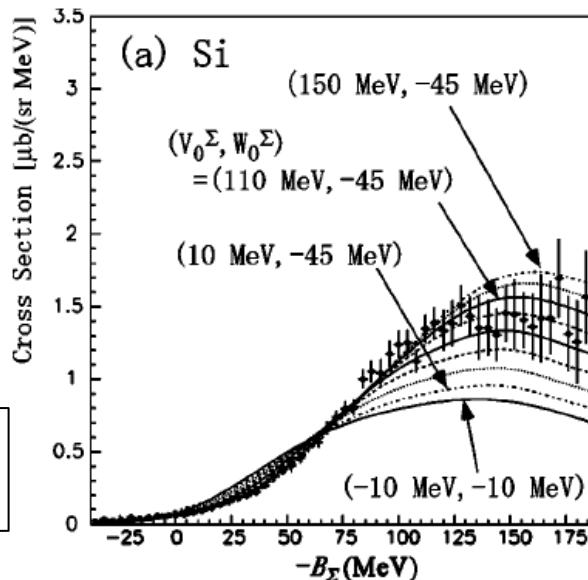
Lane term
 repulsive

- Quasi free Σ production off nuclei
 - Use quasi-free process to estimate ΣN interaction



Analysis with Green's function method

- Quasi-free $p(\pi^-, K^+) \Sigma^-$ reaction off C, Si, Ni, In and Bi



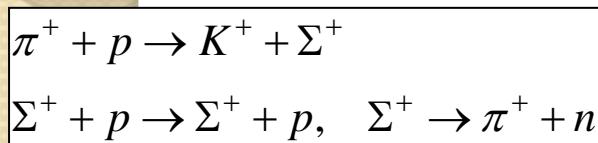
$U^\tau \propto A^{-1}$ negligible

U^0 isoscalar potential

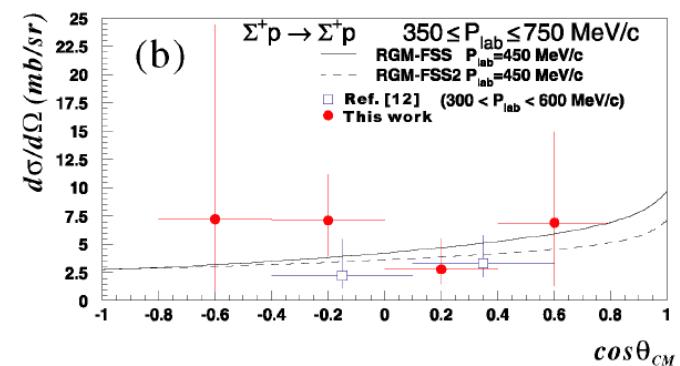
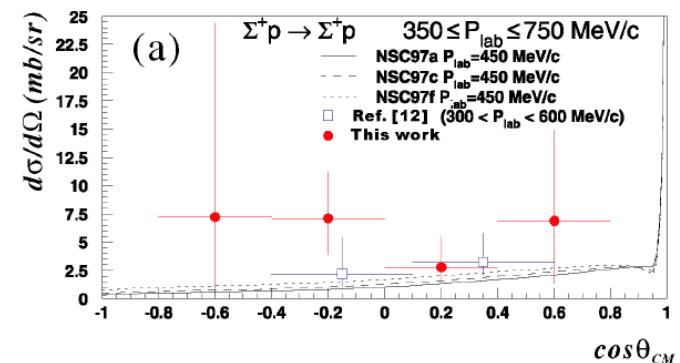
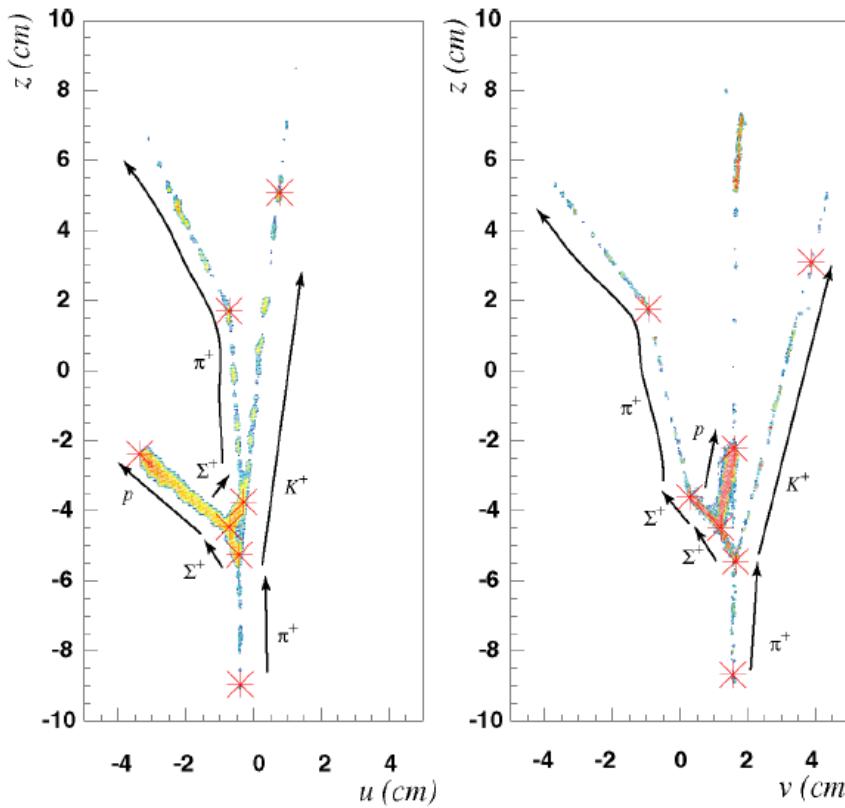
V_0^Σ : (strong) repulsive
is favored

P.K. Saha, H. Noumi et al.,
PRC 70 (2004) 044613

- ΣN elastic scattering
 - Most straight forward method to know ΣN interaction
 - Technically quite difficult (event selection and statistics)



$$U_{\Sigma^+ p} \approx U_{\Sigma^- n}$$

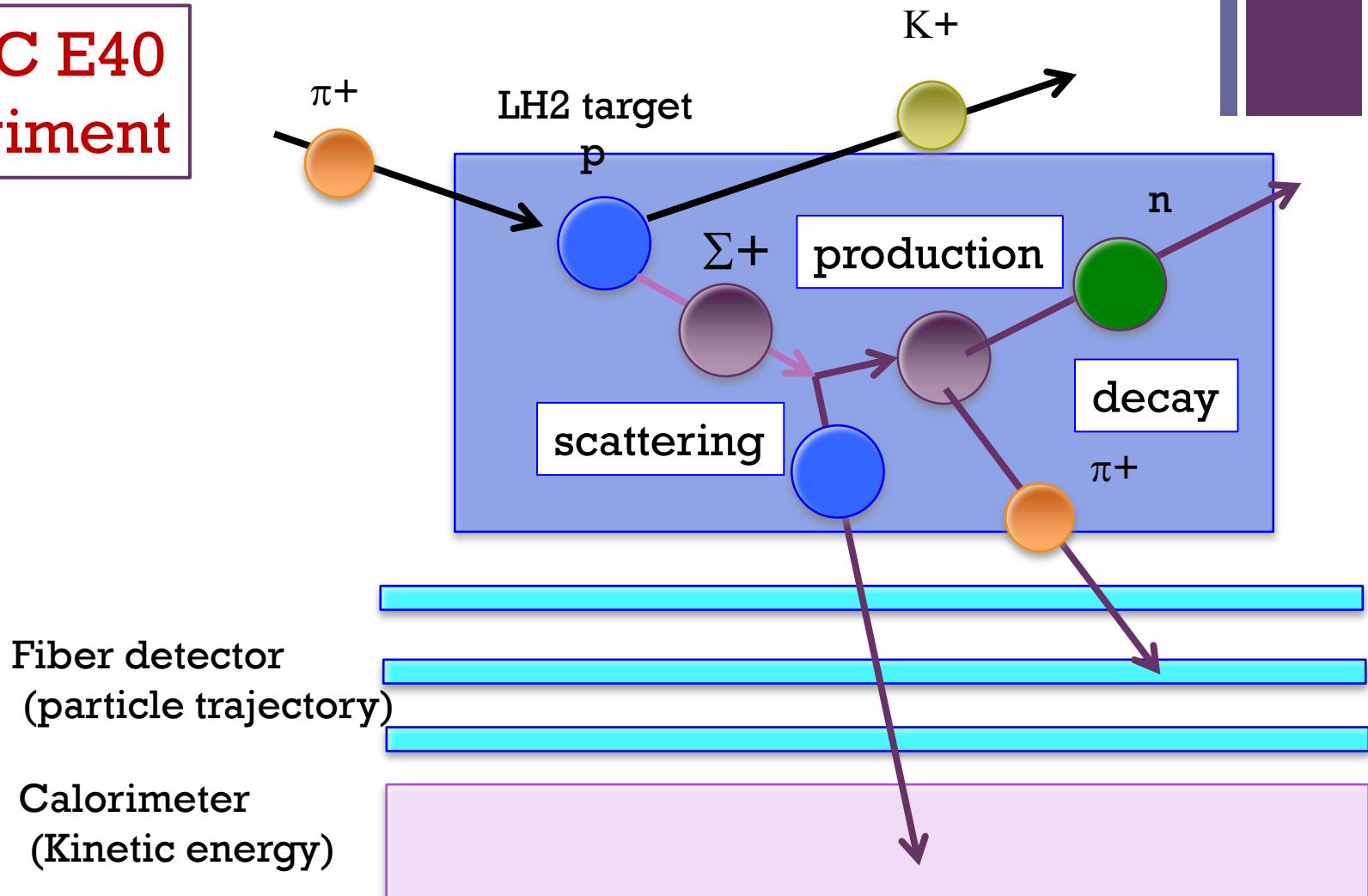


J.K. Ahn et al., NPA 761 (2005) 41



Principal of Σp scattering

J-PARC E40
experiment





Experimental setup at K1.8 beamline

10

■ Σ beam spectrometer

--- Σ beam part ---

- $\pi^\pm p \rightarrow K^\pm \Sigma^\pm$ reaction
- K1.8 spectrometer
- SKS spectrometer

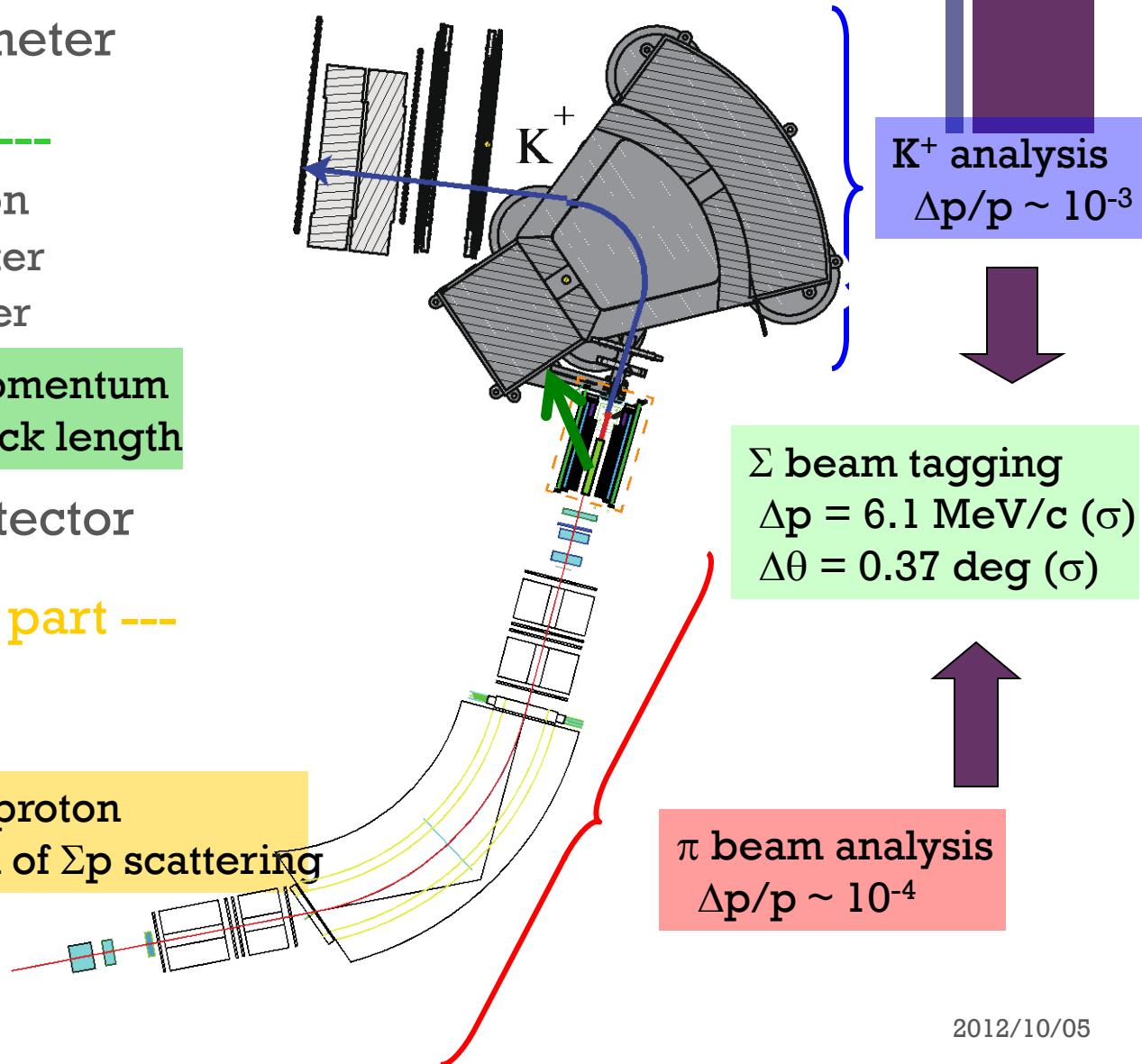
Σ beam momentum
 Σ beam track length

■ Σp scattering detector

--- Σp scattering part ---

- $\Sigma p \rightarrow \Sigma'' p''$
- $\Sigma p \rightarrow n'' \pi''$

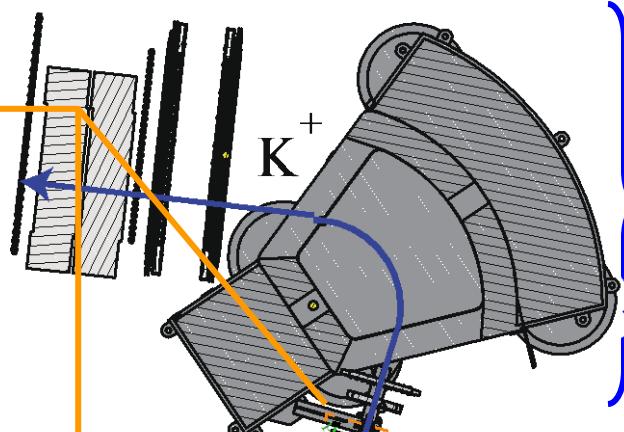
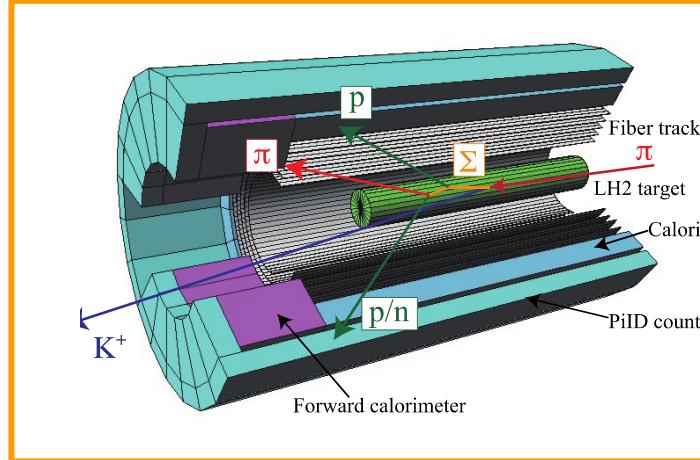
Detection of proton
Identification of Σp scattering



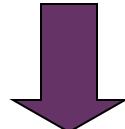


Experimental setup at K1.8 beamline

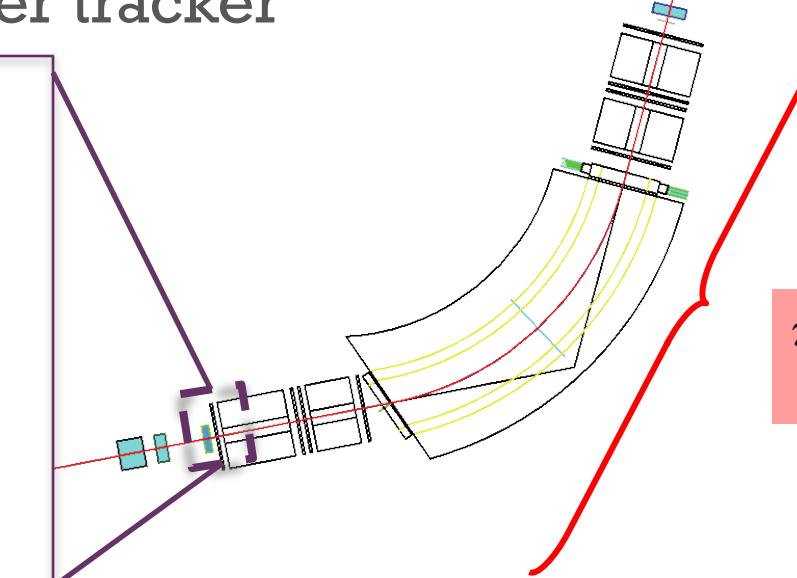
■ Scattered proton detector



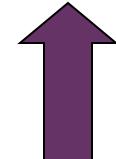
K^+ analysis
 $\Delta p/p \sim 10^{-3}$



■ Beamline fiber tracker



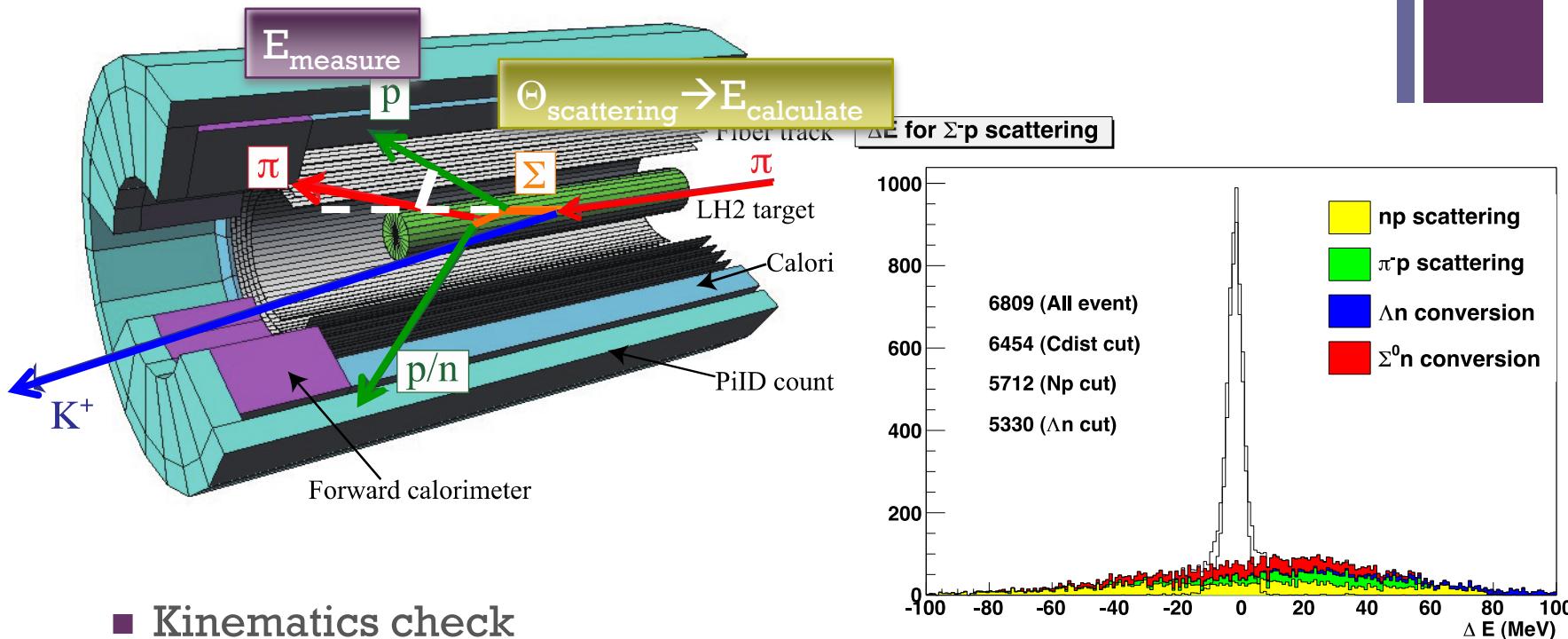
Σ beam tagging
 $\Delta p = 6.1 \text{ MeV}/c (\sigma)$
 $\Delta\theta = 0.37 \text{ deg} (\sigma)$



π beam analysis
 $\Delta p/p \sim 10^{-4}$



Identification of Σp scattering

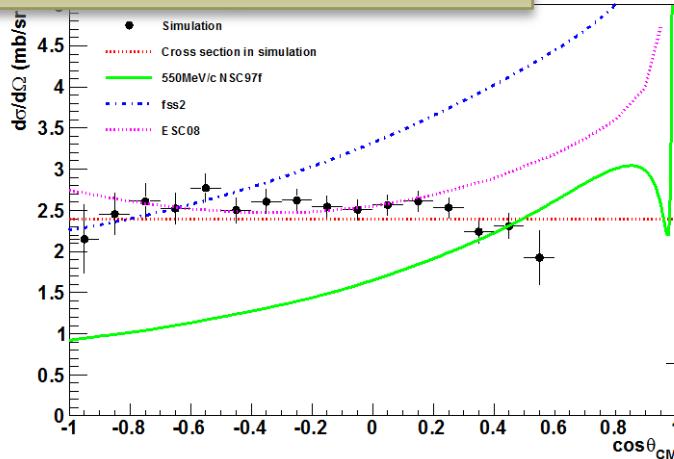


■ Kinematics check

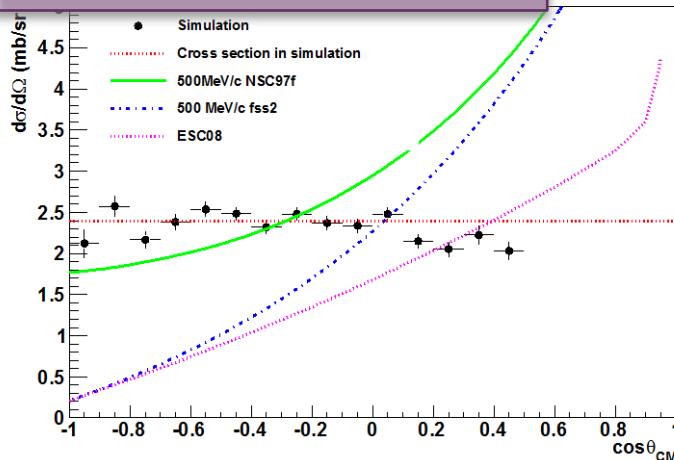
- $E_{\text{calculate}}$ (determined from θ assuming $\Sigma^- p$ scattering)
- E_{measure} (measured by Calorimeter)
- $\Delta E = E_{\text{measure}} - E_{\text{calculate}}$
 - ΔE should be 0 for $\Sigma^- p$ event.
 - Due to the kinematically overlapped region, there is a contamination of background around $\Delta E \sim 0$

+ Systematic study of ΣN interaction

$\Sigma^+ p \frac{d\sigma}{d\Omega} (p = 550 \text{ MeV}/c)$



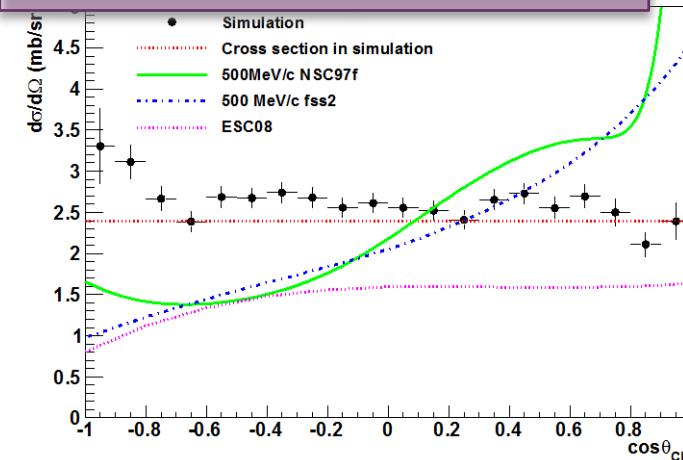
$\Sigma^- p \frac{d\sigma}{d\Omega} (p = 500 \text{ MeV}/c)$



Enough acceptance to test framework
higher wave contribution
long range region (meson picture)

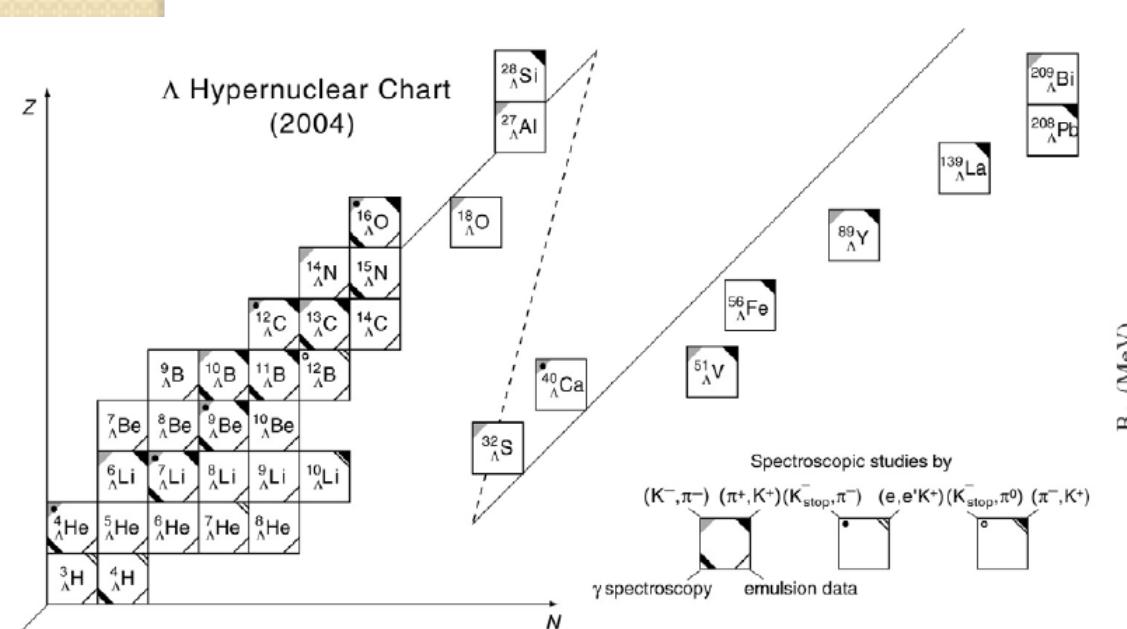
- We will measure the angular dependence of $d\sigma/d\Omega$ with 0.2 mb/sr precision.

$\Sigma^- p \rightarrow \Lambda n \frac{d\sigma}{d\Omega} (p = 500 \text{ MeV}/c)$

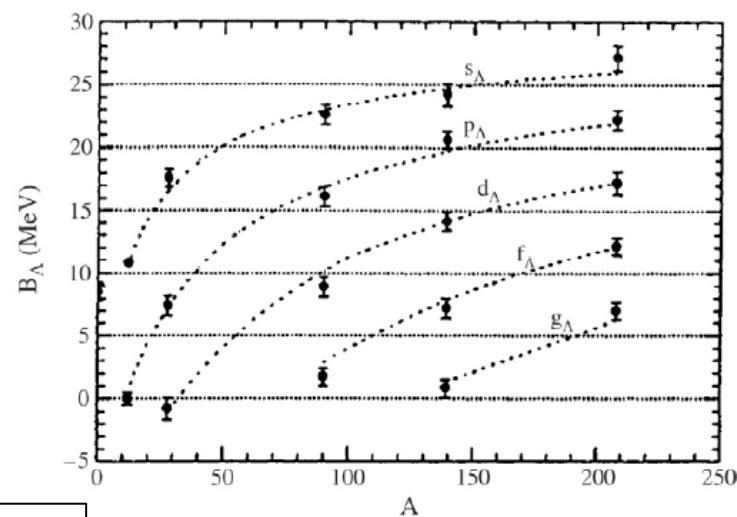


ΛN interaction (J-PARC E13/E10)

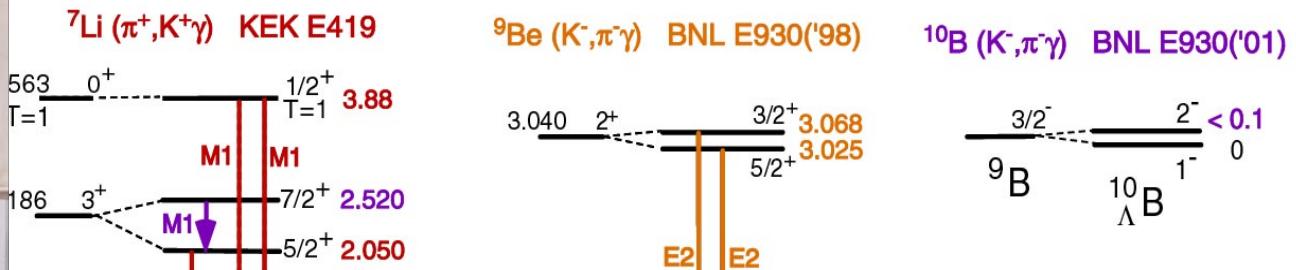
- Information on ΛN interaction from Λ hypernuclei
 - Light Λ hypernuclei
 - (K^-, π^-) and (π^+, K^+) reactions, and emulsion experiments
 - Medium to heavy Λ hypernuclei close to stability line
 - (K^-, π^-) and (π^+, K^+) reactions



$$V_0^\Lambda \approx -30 MeV$$



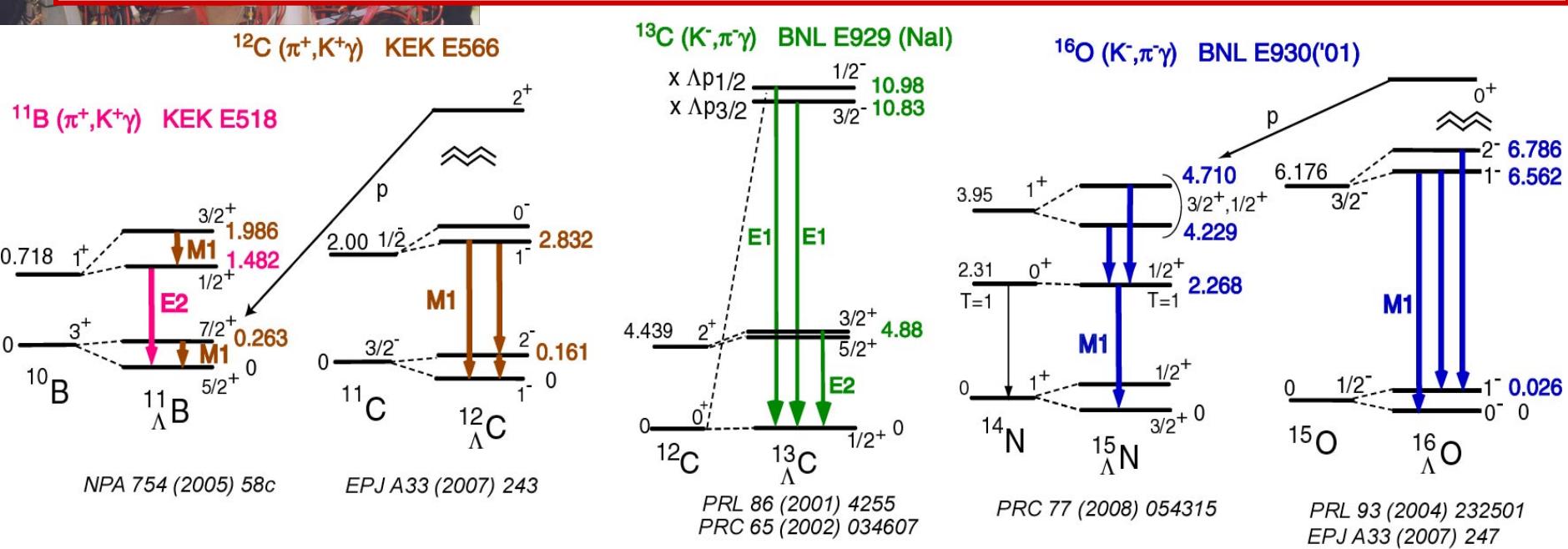
O. Hashimoto, H. Tamura, Prog. Part. Nucl. Phys. 57 (2006) 564



ΛN スピン依存相互作用を決定した

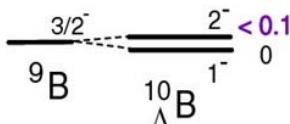
$$\Delta = 0.3 \sim 0.4, S_A = -0.01, S_N = -0.4, T = 0.03 \text{ MeV}$$

バリオン間力の理論模型に強い制限を与え、模型が改良された

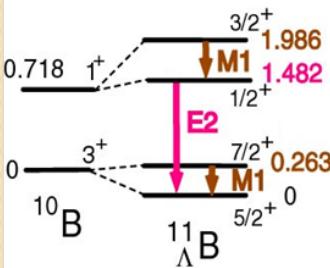


- Anomalies and missing information in ΛN interaction

^{10}B ($K^-, \pi^- \gamma$) BNL E930('01)

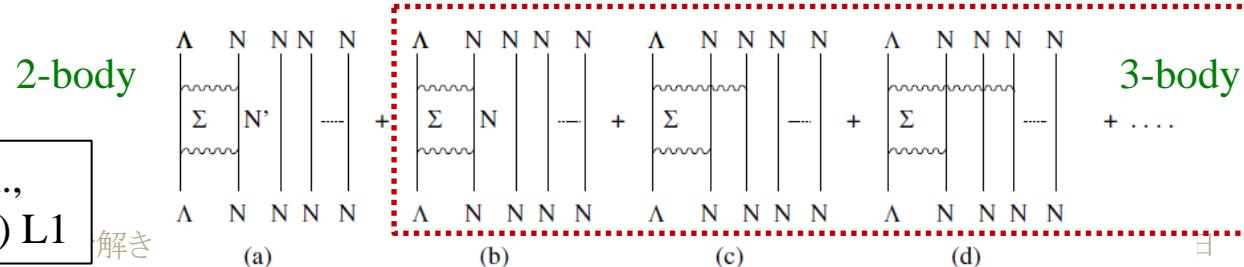


^{11}B ($\pi^+, K^+ \gamma$) KEK E518



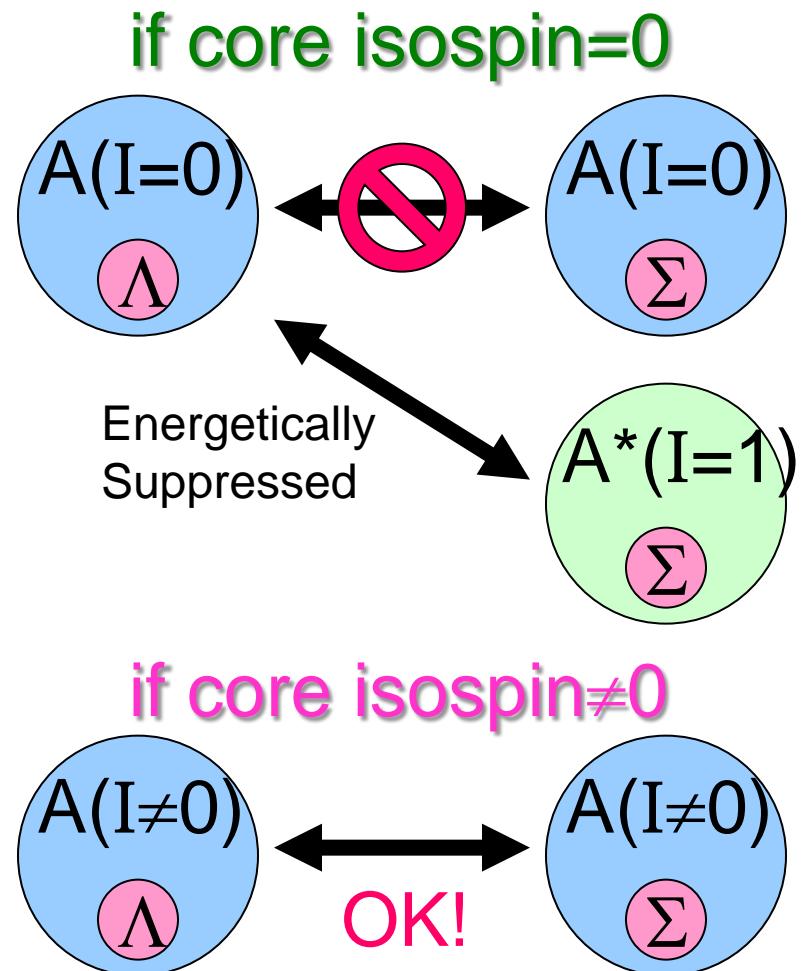
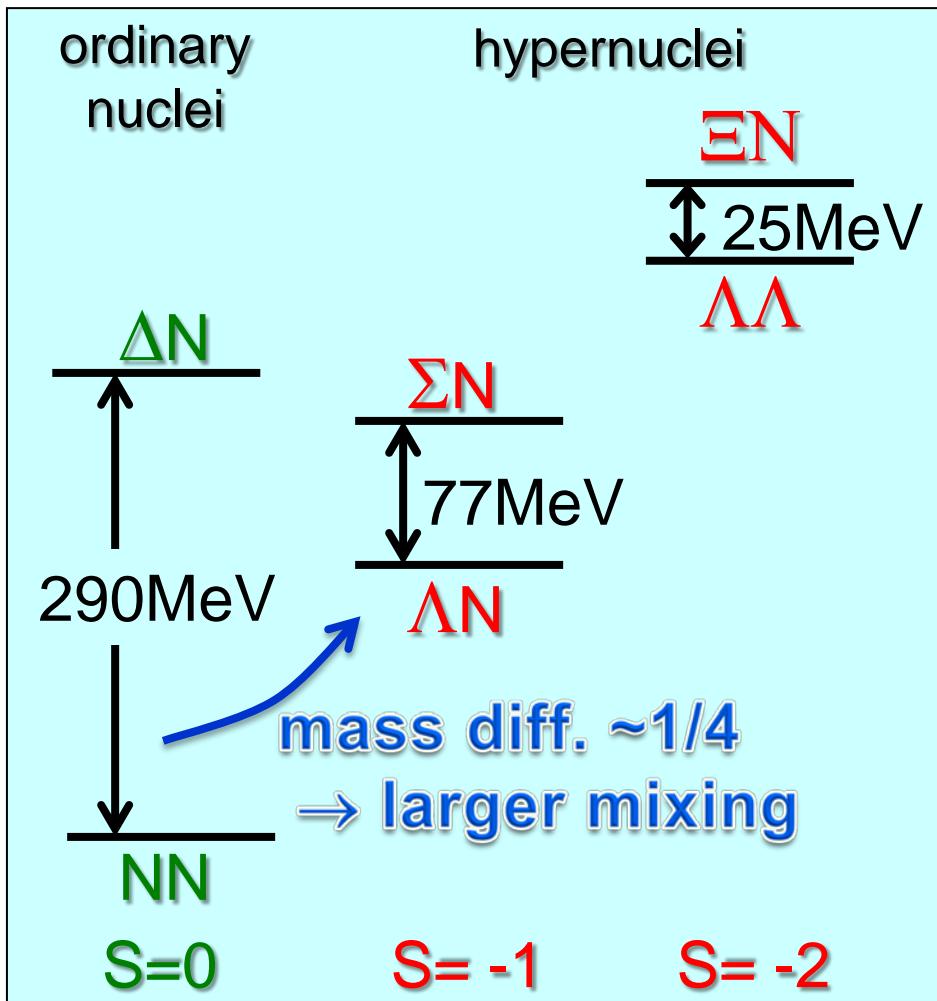
- γ -ray spectroscopy and shell-model parameterization were quite successful, but there are some anomalies
 - Quite small (<100keV) 1⁻ and 2⁻ splitting or level inversion in $^{10}\Lambda\text{B}$ hypernucleus
 - Measured 5/2⁺ and 1/2⁺ splitting in $^{11}\Lambda\text{B}$ hypernucleus is far from theoretical estimation (1020 keV)
 - One possible explanation is large ΛN - ΣN mixing contribution
- Many Λ hypernuclei were studied but only close to stability-line
 - We don't know ΛN interaction in neither neutron-rich environment nor proton-rich environment
 - ΛN - ΣN mixing may induce strong effects of neutron- and proton-rich environments

added coherently



ΛN - ΣN Mixing in Λ Hypernuclei

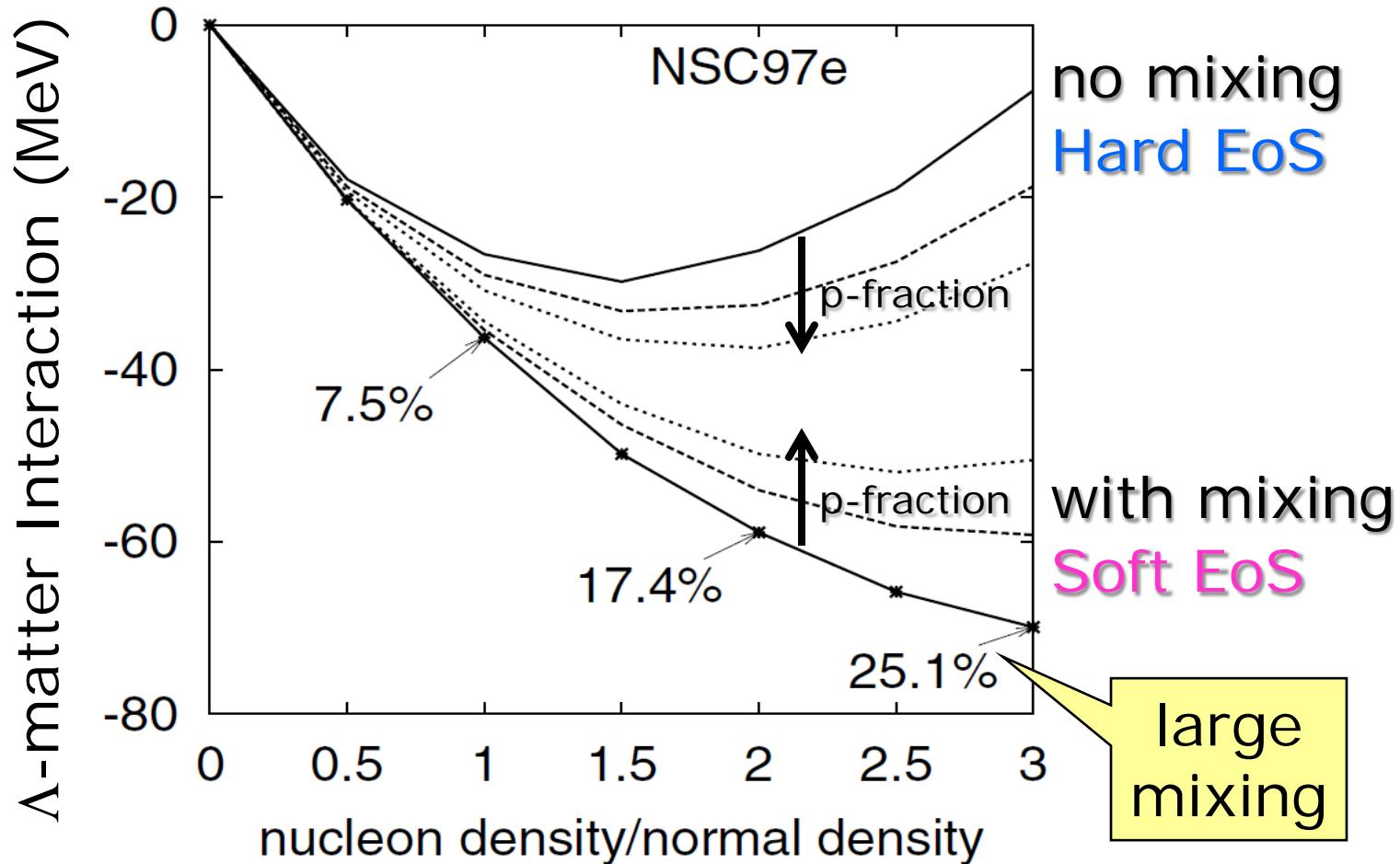
□ B.F. Gibson et al. PR C6 (1998) 433c



important in neutron-rich Λ -hypernuclei (large isospin)

Mixing and EoS in Neutron Stars

- Hyperon content in core of neutron stars
 - Degree of $\Lambda N - \Sigma^0 N$ mixing and EoS



Purpose of E13 experiment

Approved as DAY1, Second priority

ΛN interaction

Using (K^-, π^-) reaction at $p_K = 1.5$ (or 1.1) GeV/c

- (1) Charge symmetry breaking in ΛN interaction and spin-flip property in hypernuclear production

${}^4_{\Lambda}\text{He}$: Largest CSB is suggested but previous data is suspicious.
Easiest (100 hrs)

- (2) Radial dependence of ΛN interaction from sd-shell hypernuclei

${}^{19}_{\Lambda}\text{F}$: The first sd-shell hypernuclei (100 hrs)

- (3) Study $\Lambda N - \Sigma N$ coupling force

${}^{10}_{\Lambda}\text{B}$ and ${}^{11}_{\Lambda}\text{B}$: (100+200 hrs)

Inconsistency exist but previous data not enough.
Few-body approach as well as shell model possible.

g-factor of Λ in nucleus

- (4) Spin-flip B(M1) measurement and g_{Λ} in a nucleus

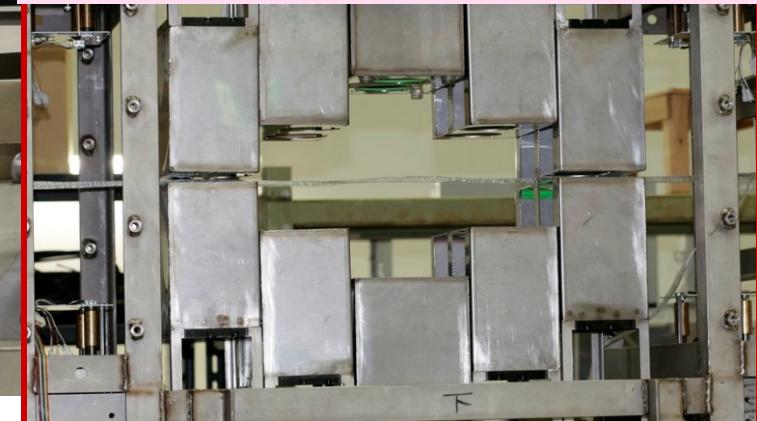
${}^7_{\Lambda}\text{Li}$: Least ambiguities exist and most reliable. (500 hrs)



Hyperball-Jに対するupgrade:

- 最下流部のさらなる高計数率化
- 波形解析による超高速読み出し法の開発・導入

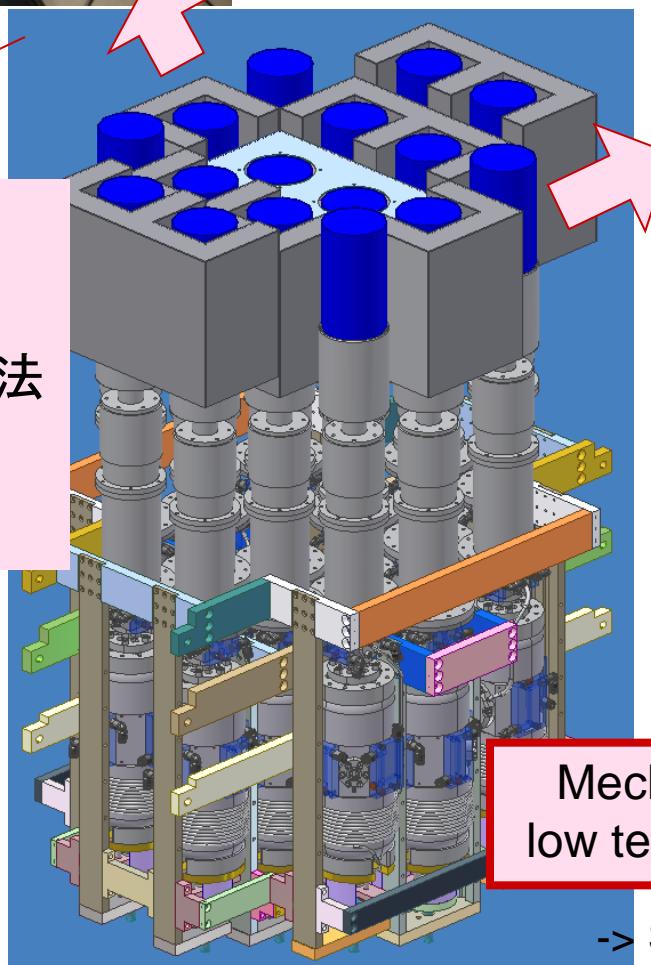
が必要



Hyperball-J

New generation Ge array

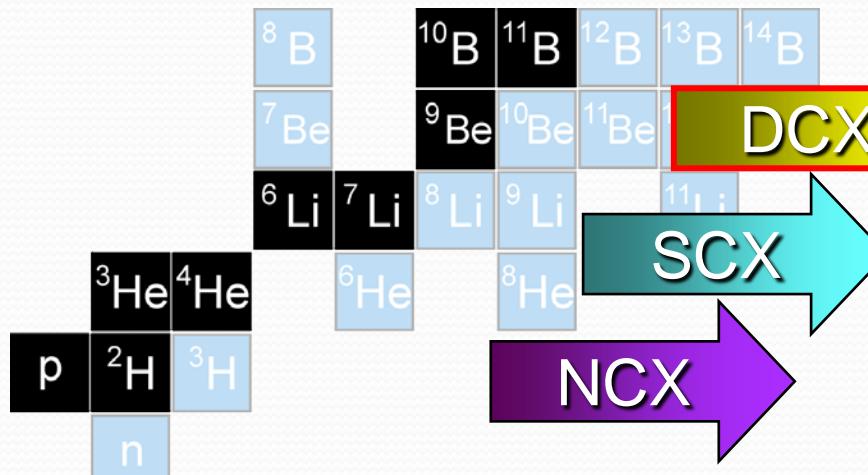
Lower half ↴



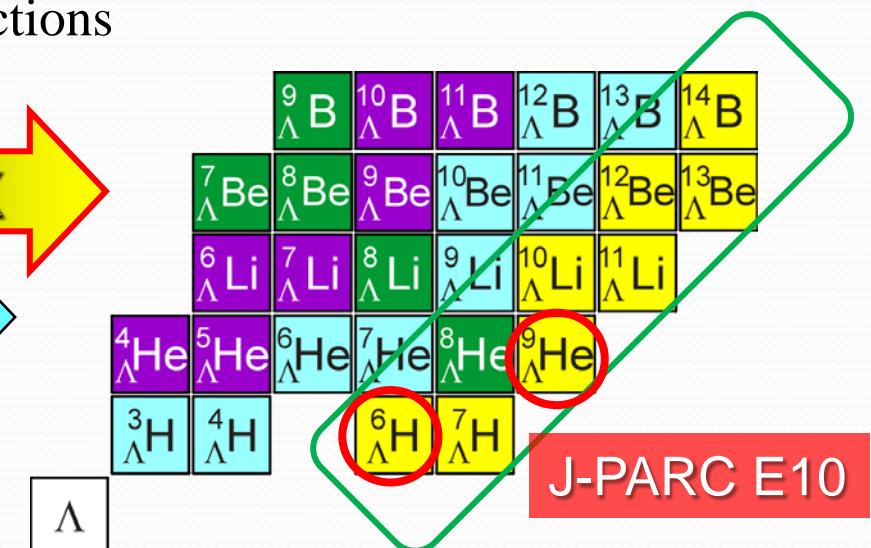
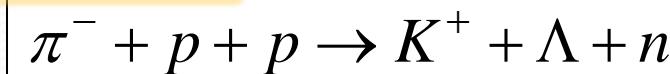
-> Suppress effects from radiation damage

E10: Produce neutron-rich hypernuclei

- Use double charge-exchange (DCX) reactions
 - Category of reactions to produce Λ hypernuclei
 - **NCX:** (π^+, K^+) and (K^-, π^-) reactions
 - **SCX:** ($e, e' K^+$), (π^-, K_S), (K^-, π^0) reactions, etc.
 - **DCX:** (π^-, K^+) and (K^-, π^+) reactions

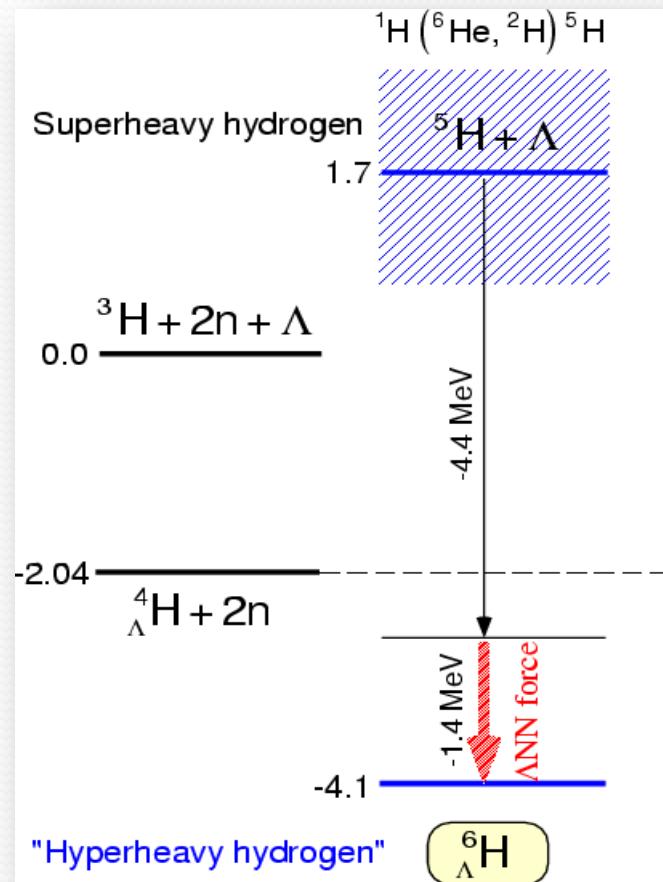


DCX reaction



the reaction has two-step nature
tiny cross section, $\sim 1/1000$ of NCX

- Investigation of ΛN - ΣN mixing
 - Precise measurement of binding energy of ${}^6_{\Lambda}H$



Suggestion of the calculation

Normal ΛN interaction

$$B_\Lambda \sim 4.4 \text{ MeV}$$

Coherent ΛN - ΣN mixing

$$B_\Lambda \sim 4.4 + 1.4 \text{ MeV}$$

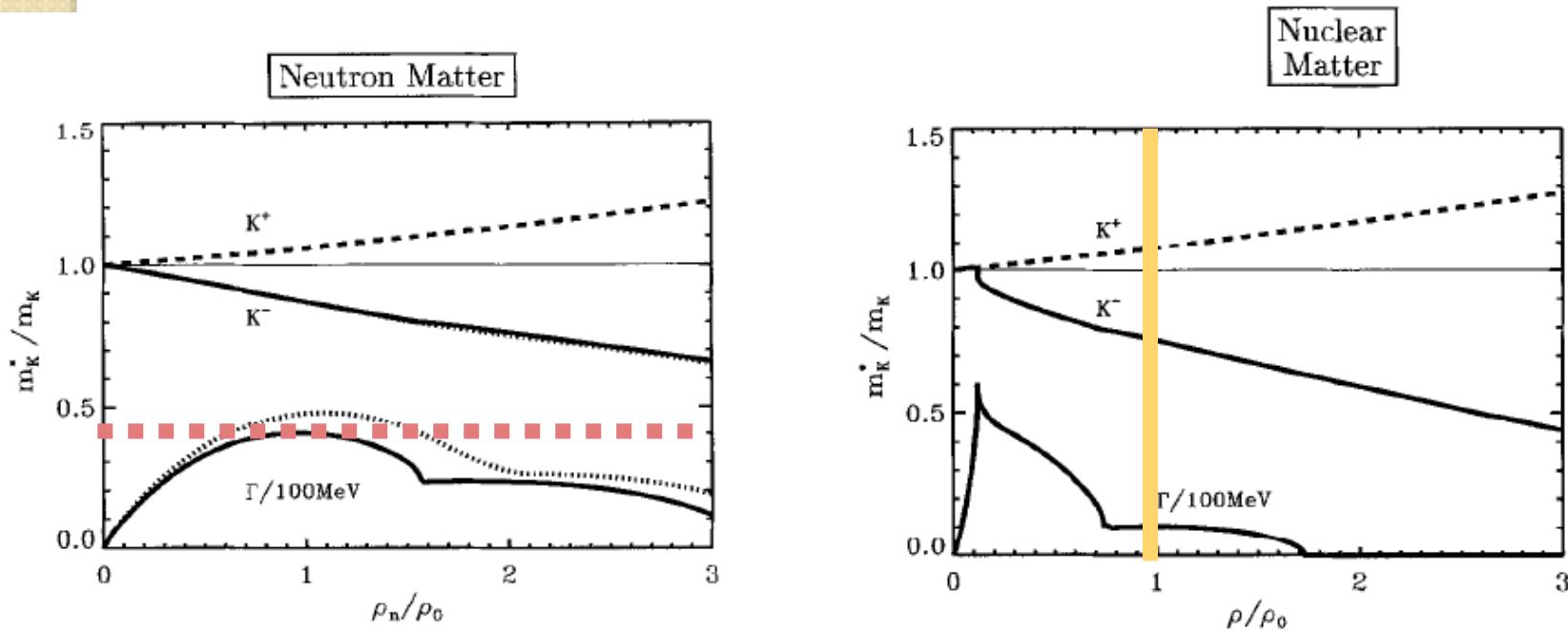
Difference is considerably large
experimentally accessible

Our basic idea

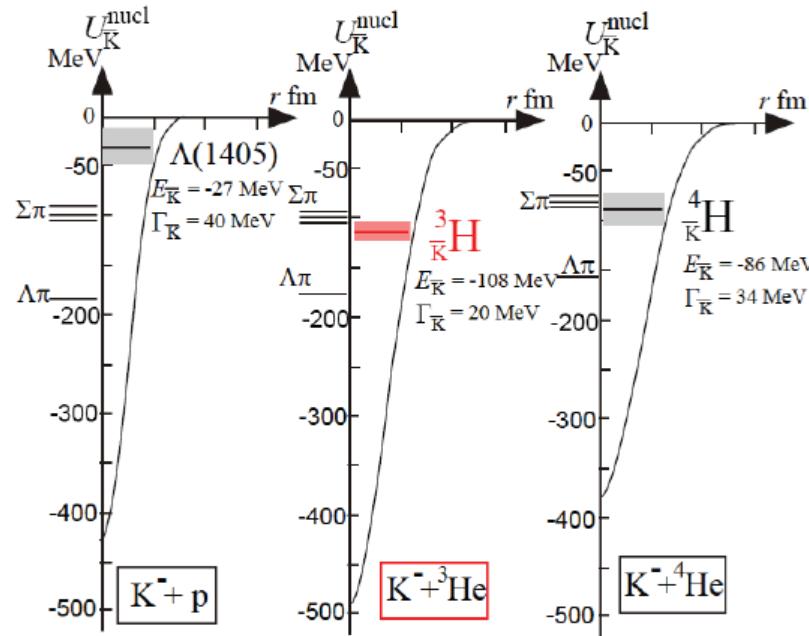
Precise measurement of B.E.
→ estimate mixing effect

KN interaction (J-PARC E15/E17/E31)

- Anti-kaon in high density nuclear matter
 - Decrease of effective mass m_K^*
 - K^- in Neutron Matter: threshold of condensation $\approx 200\text{MeV}/c^2$
 - in (Symmetric) Nuclear Matter: similar drop of m_K^*
 - Considerable drop ($>100\text{MeV}/c^2$) even at ρ_0
 - May be observed as strong attractive force in nuclei



- KN interaction and property of $\Lambda(1405)$
 - $\Lambda(1405)$: $S=-1$, $I=0$ and $J^\pi=1/2^-$ resonance
 - Mass is considerably lower than quark model prediction
 - Large fraction of KN hybrid (bound system)?
 - If $\Lambda(1405)$ is KN hybrid, it means strong KN attraction
 - Possible production of K-nucleus bound state

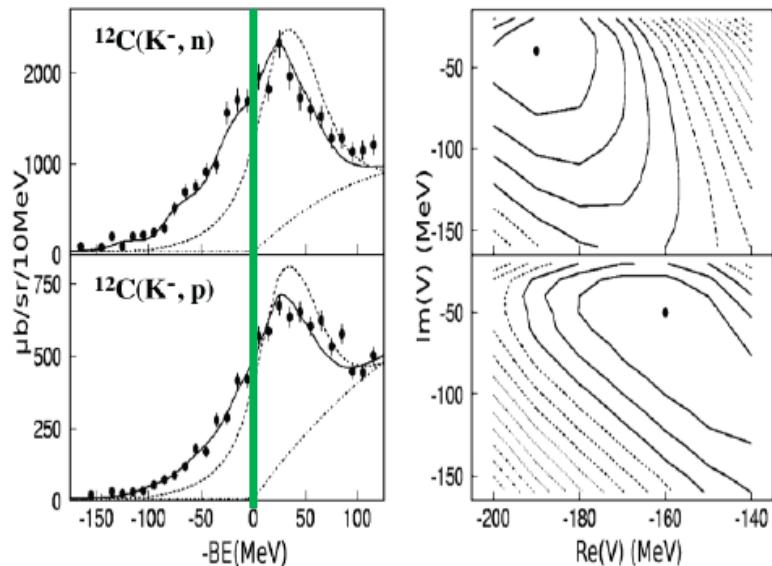


K^-

N

Knockout neutron and
embed K^- in nuclei

in-flight (K^-, n) reaction @ 1 GeV/c



fit = Green's function

T. Kishimoto et al., Prog. Theor. Phys. 118 (2007) 18

indicating very deep potential

- deep & wide KN pot. - lower background

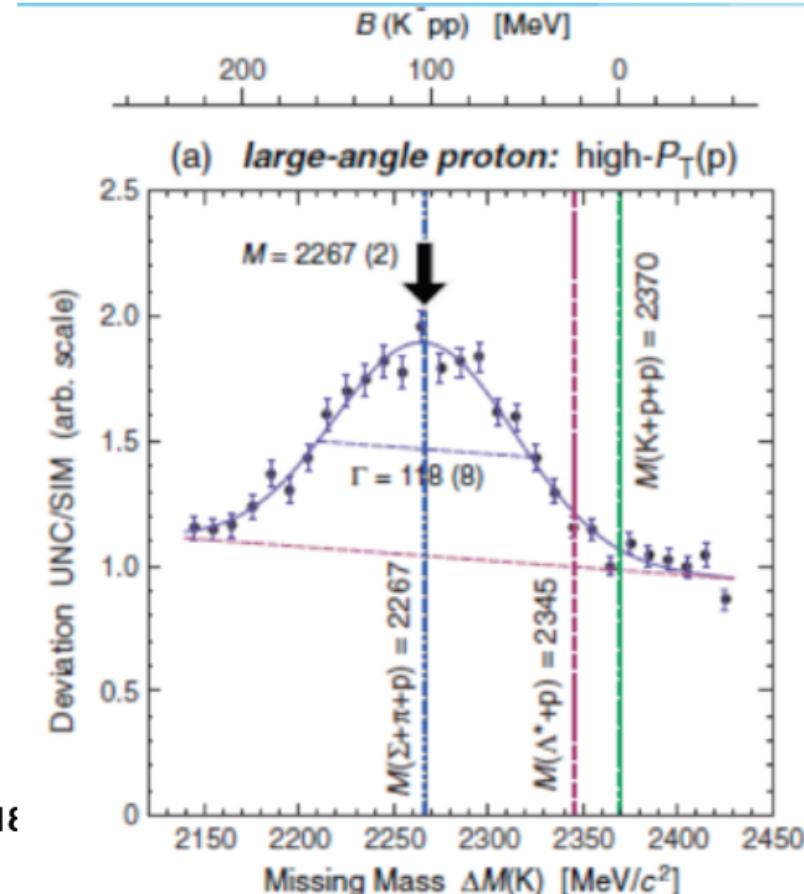
$Re(V) \sim 180$ MeV

$Im(V) \sim 50$ MeV

in-flight ensures ...

2N process suppressed

$p(p, K^+) X @ T = 2.85 \text{ GeV}$



T. Yamazaki et al., PRL104(2010)132502

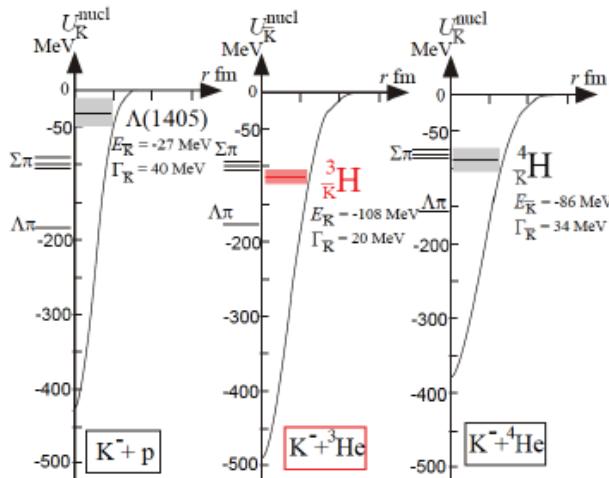
Kpp state? at

$M_x = 2267 \pm 3 \pm 5$ MeV !

$\Gamma_x = 118 \pm 8 \pm 10$ MeV !

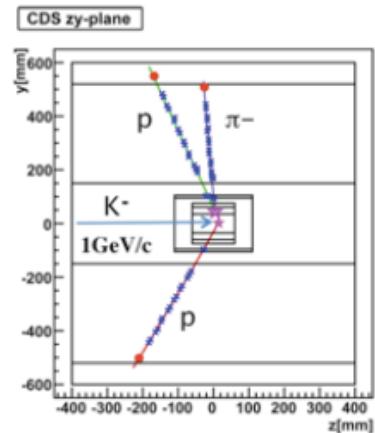
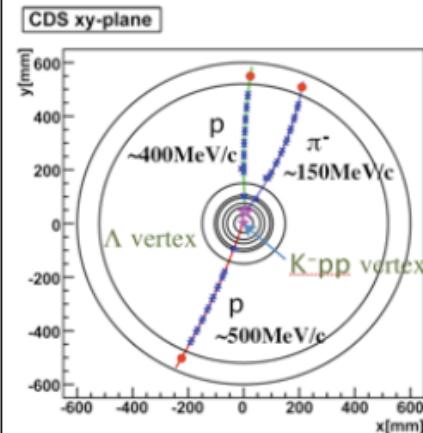
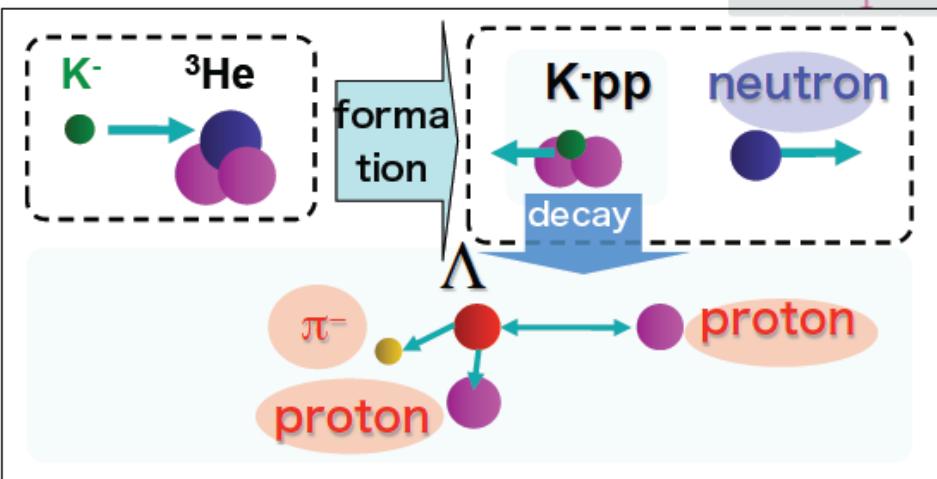
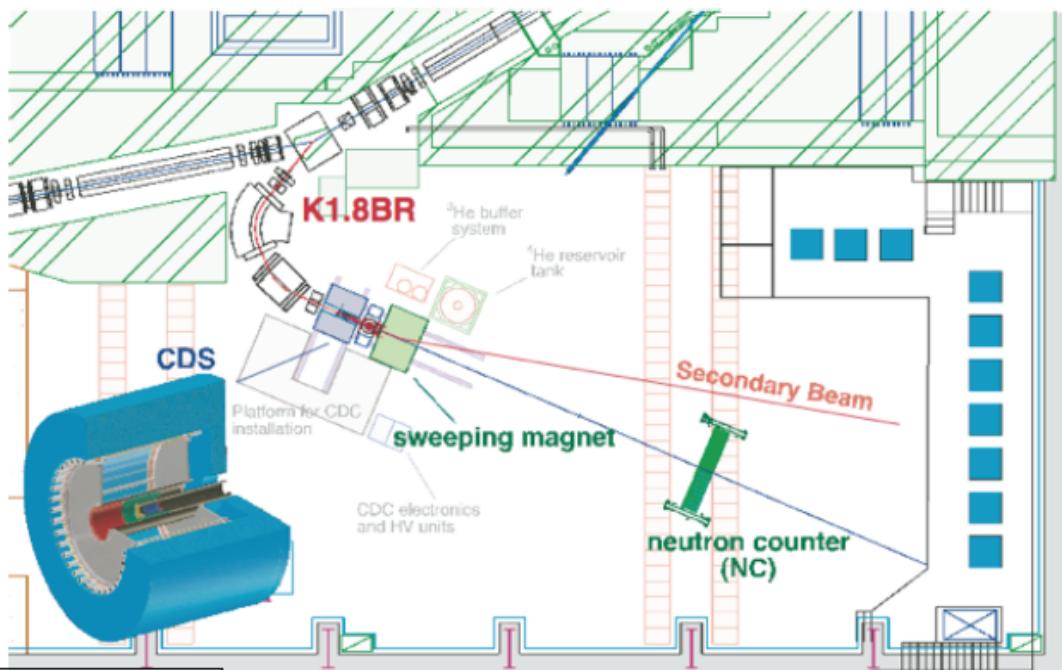


E15: $\bar{K}N$ interaction study by nuclear bound state

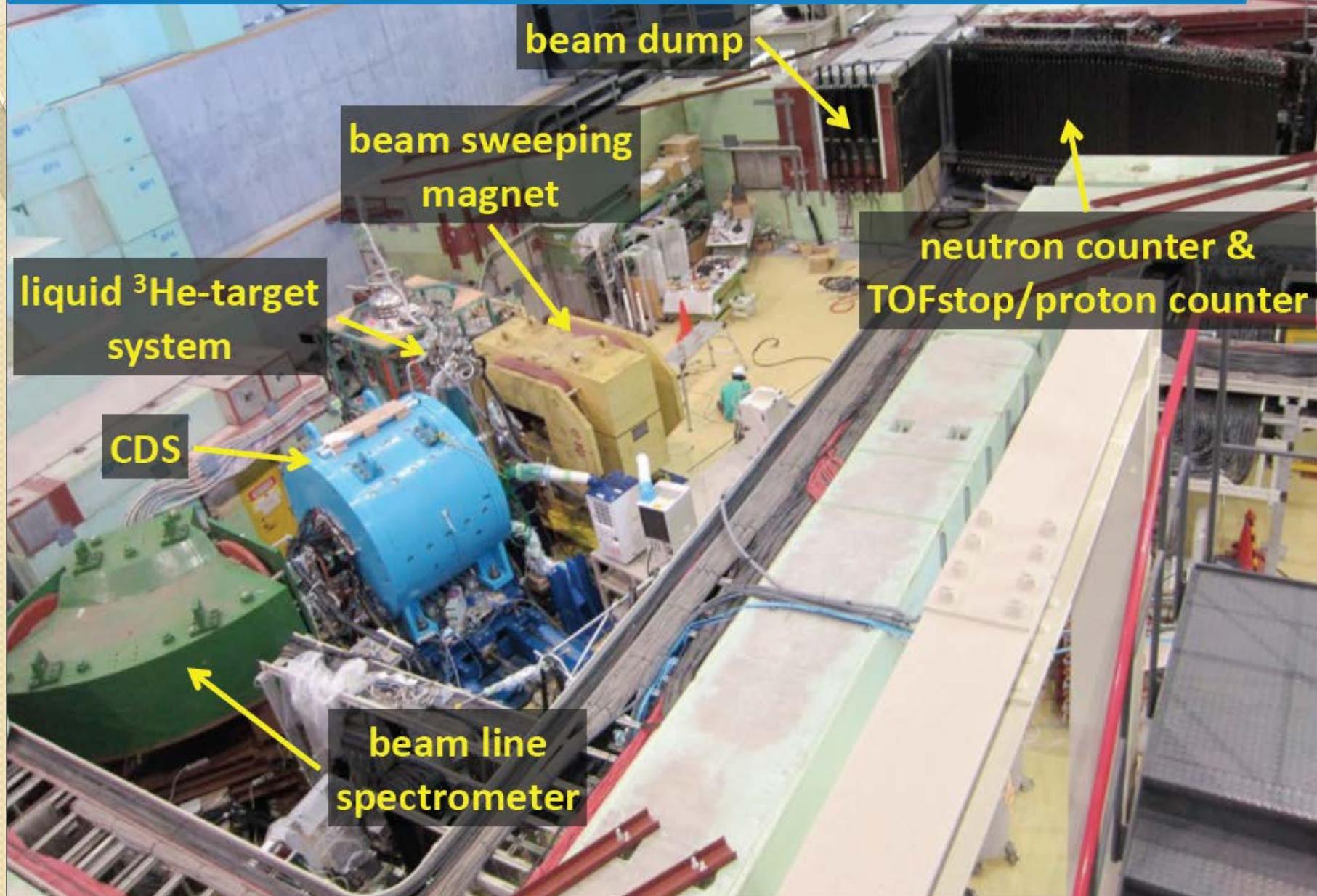


$K^- + {}^3\text{He} \rightarrow \text{"pp } K^- + n$

at 1 GeV/c by both
missing & invariant mass

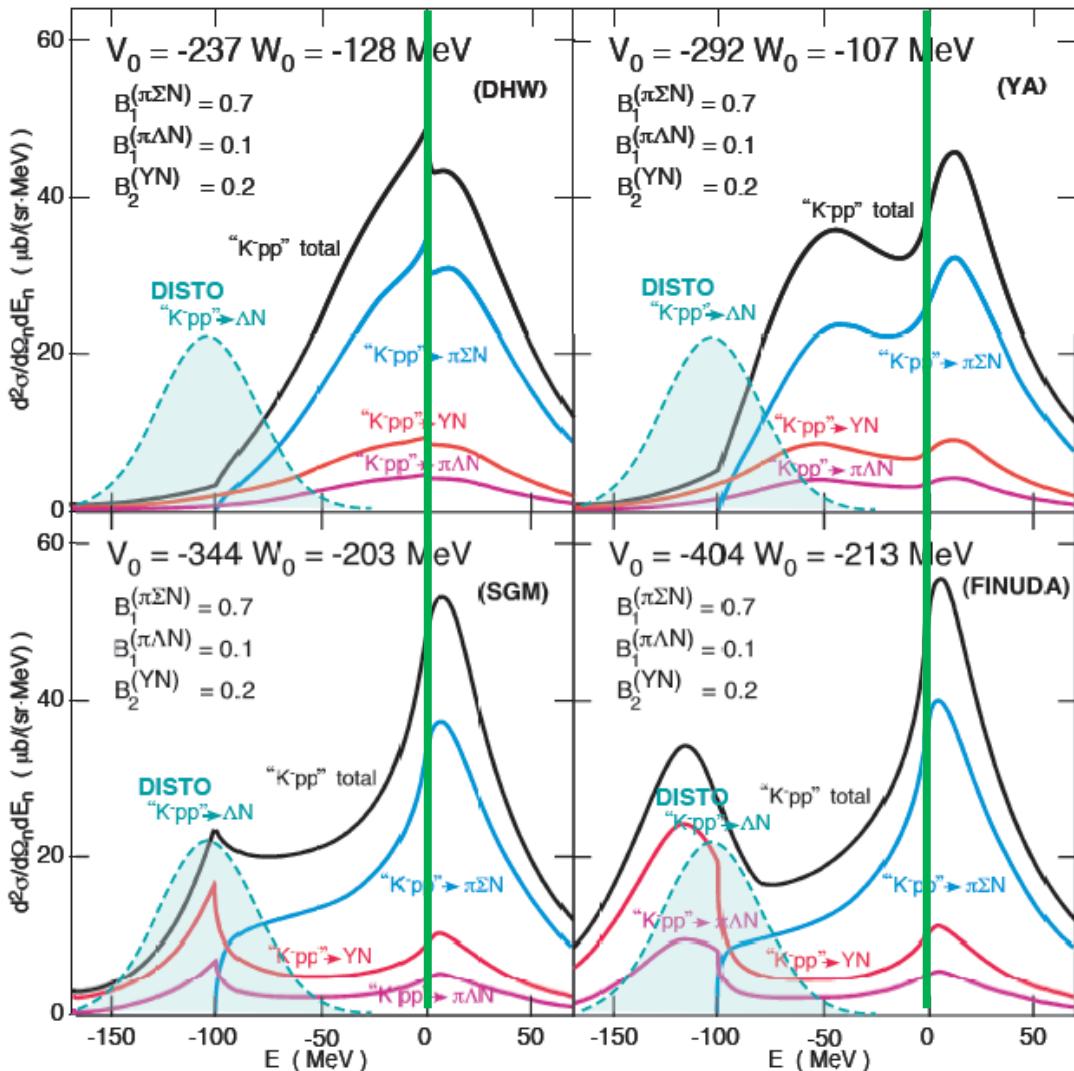


the completed K1.8BR spectrometer [RUN#43, Jun. 2012]





Koike-Harada vs DISTO



DHW: A. Dote, T. Hyodo, and W. Weise,
Nucl. Phys. A804, 197 (2008);
Phys. Rev. C79, 014003 (2009).

YA: T. Yamazaki and Y. Akaishi,
Phys. Lett. B535, 70 (2002);
Proc. Jpn. Academy, Series B 83, 144 (2007)

SGM: N.V. Shevchenko, A. Gal, and J. Mares,
Phys. Rev. Lett. 98, 082301 (2007);
N.V. Shevchenko, A. Gal, J. Mares, and J. Revai,
Phys. Rev. C76, 044004 (2007).

FINUDA: M. Agnello et al.,
Phys. Rev. Lett. 94, 212303 (2005).

DISTO

B_K~100MeV and Γ_K~100MeV

- only for Λp decay ch.

private communication

- does not fit in KH scheme

easy to observe,
if $d\sigma/d\Omega \gtrsim 1\text{mb/sr}$

Summary

- Does the inner-core have **Exotic components?**
 - One of important factors to determine **structure of neutron stars**
 - Candidates of Exotic components are **Hyperons** and **Kaons**
 - Negative charge, mass, **attractive interaction** and **symmetry energy** control appearance
- J-PARC is the best laboratory for the investigation
 - Several experiments are in progress and in preparation
 - **E40**: high statics measurements of ΣN elastic scattering
 - **E13**: investigation of ΛN - ΣN mixing by precise γ -ray spectroscopy
 - **E10**: ΛN interaction in neutron-rich environment and ΛN - ΣN mixing
 - **E15**: search for Kpp bound state and strongly attractive KN interaction